# ON NON-RADIAL SINGULAR SOLUTIONS OF SUPERCRITICAL BI-HARMONIC EQUATIONS 

ZONGMING GUO, JUNCHENG WEI, AND WEN YANG


#### Abstract

We develop gluing method for fourth order ODEs and construct infinitely many non-radial singular solutions for a bi-harmonic equation with supercritical exponent.


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## 1. Introduction

In this paper we are concerned with positive singular solution of the following bi-harmonic equation:

$$
\begin{equation*}
\Delta^{2} u=u^{p} \quad \text { in } \mathbb{R}^{n}, \quad n \geq 6 \tag{1.1}
\end{equation*}
$$

where $p>\frac{n+4}{n-4}$.
Equation (1.1) arises both in physics and geometry. In the past decades there have been lots of research activities in classifying solutions to (1.1). When $1<p \leq$ $\frac{n+4}{n-4}$, all nonnegative solutions to (1.1) have been completely classified (Lin ([18]) and Wei- $\mathrm{Xu}([19]))$ : if $p<\frac{n+4}{n-4}$ (1.1) admits no nontrivial nonnegative regular solution, while $p=\frac{n+4}{n-4}$, i.e. the critical case, any positive regular solution of (1.1) can be written as the following form:

$$
u_{\lambda, \xi}=[n(n-4)(n-2)(n+2)]^{-\frac{n-4}{8}}\left(\frac{\lambda}{1+\lambda^{2}|x-\xi|^{2}}\right)^{\frac{n-4}{2}}, \xi \in \mathbb{R}^{n}
$$

However, the question on complete classification of positive regular solutions of (1.1) in the supercritical case, i.e. $p>\frac{n+4}{n-4}$, remains largely open.

The structure of positive radial solutions of (1.1) with $p>\frac{n+4}{n-4}$ has been studied by Gazzola-Grunau [7] and Guo-Wei [10]. For the following fourth order ODE:

$$
\left\{\begin{array}{l}
\Delta^{2} u(r)=u^{p}(r), \quad r \in[0, \infty)  \tag{1.2}\\
u(0)=a, u^{\prime \prime}(0)=b, u^{\prime}(0)=u^{\prime \prime \prime}(0)=0
\end{array}\right.
$$

it is known from [7] that for any $a>0$ there is a unique $b_{0}:=b_{0}(a)<0$ such that the unique solution $u_{a, b_{0}}$ of (1.2) satisfies that $u_{a, b_{0}} \in C^{4}(0, \infty), u_{a, b_{0}}^{\prime}(r)<0$ and

$$
\lim _{r \rightarrow \infty} r^{\alpha} u_{a, b_{0}}(r)=K_{0}^{\frac{1}{p-1}}
$$

where $\alpha=\frac{4}{p-1}$ and
$K_{0}=\frac{8}{(p-1)^{4}}\left[(n-2)(n-4)(p-1)^{3}+2\left(n^{2}-10 n+20\right)(p-1)^{2}-16(n-4)(p-1)+32\right]$.
This implies that $u_{a, b_{0}}(r)>0$ for all $r>0$ and $u_{a, b_{0}}(r) \rightarrow 0$ as $r \rightarrow \infty$. Moreover, it is known from [10] that if $5 \leq n \leq 12$ or if $n \geq 13$ and $\frac{n+4}{n-4}<p<p_{c}(n)$,
$u_{a, b_{0}}-K_{0}^{\frac{1}{p-1}} r^{-\alpha}$ changes sign infinitely many times in $(0, \infty)$, if $n \geq 13$ and $p \geq$ $p_{c}(n)$, then $u(r)<K_{0}^{\frac{1}{p-1}} r^{-\alpha}$ for all $r>0$ and the solutions are strictly ordered with respect to the initial value $a=u_{a, b_{0}}(0)$. Here $p_{c}(n)$ refers to the unique value of $p>\frac{n+4}{n-4}$ such that

$$
p_{c}(n)= \begin{cases}+\infty, & \text { if } 4 \leq n \leq 12, \\ \frac{n+2-\sqrt{n^{2}+4-n \sqrt{n^{2}-8 n+32}}}{n-6-\sqrt{n^{2}+4-n \sqrt{n^{2}-8 n+32}},} & \text { if } n \geq 13 .\end{cases}
$$

Very recently, Dávila-Dupaigne-Wang-Wei [5] proved that all stable or finite Morse index solutions of the equation (1.1) are trivial provided $1<p<p_{c}(n)$. According to the result by [10] and [16] all radial solutions are stable when $p \geq p_{c}(n)$. Thus the result in [5] is sharp.

We now turn to the singular solutions of (1.1). It is easily seen that

$$
\begin{equation*}
u_{s}(x):=K_{0}^{\frac{1}{p-1}}|x|^{-\frac{4}{p-1}} \tag{1.3}
\end{equation*}
$$

is a singular solution of (1.1), in other words, $u_{s}$ satisfies the equation

$$
\begin{equation*}
\Delta^{2} u=u^{p}, \quad u>0 \text { in } \mathbb{R}^{n} \backslash\{0\} \tag{1.4}
\end{equation*}
$$

As far as we know, the radial singular solution in (1.3) is the only singular solution to (1.4) known so far. The question we shall address in this paper is whether or not there are non-radial singular solutions to (1.4). To this end, we first discuss the corresponding second order Lane-Emden equation

$$
\begin{equation*}
\Delta u+u^{p}=0, \quad u>0 \quad \text { in } \mathbb{R}^{n} \tag{1.5}
\end{equation*}
$$

Equation (1.5) has been widely studied in the past decades. We refer to [1], [2], [3], [6], [9], [11], [12], [14], [15], [17], [22] and the references therein. Farina ([6]) proved that if $\frac{n+2}{n-2}<p<p^{c}(n)$, the Morse index of any regular solution $u$ of (1.5) is $\infty$. Here $p^{c}(n)$ is the Joseph-Lundgren exponent ([15]):

$$
p^{c}(n)= \begin{cases}+\infty, & \text { if } 2 \leq n \leq 10 \\ \frac{(n-2)^{2}-4 n+8 \sqrt{n-1}}{(n-2)(n-10)}, & \text { if } n \geq 11\end{cases}
$$

In [3], Dancer, Du and Guo showed that if $\Omega_{0}$ is a bounded domain containing $0, u$ is a solution of (1.5) in $\Omega_{0} \backslash\{0\}$; $u$ has finite Morse index and $\frac{n+2}{n-2}<p<p^{c}(n)$, then $x=0$ must be a removable singularity of $u$. They also show that if $\Omega_{0}$ is a bounded domain containing $0, u$ is a solution of (1.5) in $\mathbb{R}^{n} \backslash \Omega_{0}$ that has finite Morse index and $\frac{n+2}{n-2}<p<p^{c}(n)$, then $u$ must be a fast decay solution. It is easily seen that (1.5) has a radial singular solution

$$
u^{s}(x):=u^{s}(r)=\left[\frac{2}{p-1}\left(n-2-\frac{2}{p-1}\right)\right]^{\frac{1}{p-1}}|x|^{-\frac{2}{p-1}}
$$

In a recent paper [4], Dancer, Guo and Wei obtained infinitely many positive nonradial singular solutions of (1.5) provided $p \in\left(\frac{n+1}{n-3}, p^{c}(n-1)\right)$. The proof of [4] is via a gluing of outer and inner solutions.

The main result in this paper is the following theorem.
Theorem 1.1. Let $n \geq 6$. Assume that

$$
\frac{n+3}{n-5}<p<p_{c}(n-1)
$$

Then (1.1) admits infinitely many non-radial singular solutions.
The proof of Theorem 1.1 is via a gluing of inner and outer solutions, as in [4]. In the second order case, one glues $\left(u(r), u^{\prime}(r)\right)$ at some intermediate point. However, since equation (1.1) is fourth order, we have to match the inner solution and outer solution up to the third derivative $\left(u(r), u^{\prime}(r), u^{\prime \prime}(r), u^{\prime \prime \prime}(r)\right)$. There are some essential obstructions appearing in matching the inner and outer solutions. As far as we know this seems to be the first paper in gluing inner and outer solutions for fourth order ODE problems.

In the following, we sketch the idea of proving Theorem 1.1. After performing a separation of variables for a solution $u$ of (1.1): $u(x)=r^{-\alpha} w(\theta)$, finding a nonradial singular solutions of (1.1) is equivalent to finding a non-constant solution of the following equation:

$$
\begin{equation*}
\Delta_{S^{n-1}}^{2} w+k_{1}(n) \Delta_{S^{n-1}} w+k_{0}(n) w=w^{p} \tag{1.6}
\end{equation*}
$$

where

$$
\begin{gathered}
k_{0}(n)=(n-4-\alpha)(n-2-\alpha)(2+\alpha) \alpha \\
k_{1}(n)=-[(n-4-\alpha)(2+\alpha)+(n-2-\alpha) \alpha]
\end{gathered}
$$

It is clear that $w(\theta)=\left[k_{0}(n)\right]^{\frac{1}{p-1}}$ is the constant solution of (1.6), which provides the radial singular solution of (1.1) and it is given in (1.3).

In order to construct positive non-radial singular solutions of (1.1), we need to find positive non-constant solutions of (1.6), which is a fourth order inhomogeneous nonlinear ODE, therefore, we shall construct infinitely many positive non-constant radially symmetric solutions of (1.6), i.e. solutions that only depend on the geodesic distance $\theta \in[0, \pi)$. We only consider the simple case $w(\theta)=w(\pi-\theta)$ for $0 \leq \theta \leq \frac{\pi}{2}$. In this case, (1.6) can be written as the following form:

$$
\left\{\begin{array}{l}
T_{1} w(\theta)+k_{1}(n) T_{2} w(\theta)+k_{0}(n) w=w^{p}, \quad w(\theta)>0 \quad 0<\theta<\frac{\pi}{2},  \tag{1.7}\\
w_{\theta}^{\prime}(0), \quad w_{\theta}^{\prime \prime \prime}(0) \text { exist, } \quad w_{\theta}^{\prime}\left(\frac{\pi}{2}\right)=w_{\theta}^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

where $T_{1}, T_{2}$ are the differential operators defined as:

$$
T_{1} w(\theta)=\frac{1}{\sin ^{n-2} \theta} \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d}{d \theta}\left(\frac{1}{\sin ^{n-2} \theta} \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d w(\theta)}{d \theta}\right)\right)\right)
$$

and

$$
T_{2} w(\theta)=\frac{1}{\sin ^{n-2} \theta} \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d w(\theta)}{d \theta}\right)
$$

A key observation is that

$$
\begin{equation*}
w_{*}(\theta)=A_{p}[\sin \theta]^{-\alpha}, \quad \theta \in\left(0, \frac{\pi}{2}\right] \tag{1.8}
\end{equation*}
$$

with

$$
A_{p}^{p-1}=(n-5-\alpha)(n-3-\alpha)(2+\alpha) \alpha\left(:=k_{0}(n-1)\right)
$$

is a singular solution of (1.7) with a singular point at $\theta=0$. (Note that this is a singular solution in one dimension less.) We will construct the inner and outer solution of (1.7) and glue them at some point close to 0 , which gives solutions of (1.7). The main difficulty is the matching of four parameters, which correspond to matching of $u$ and its derivatives up to the third order.

This paper is organized as follows. In Section 2, we present some preliminaries. In Section 3, we construct inner solutions of (1.7) by studying an initial value problem of (1.7) with large initial values at $\theta=0$. In Section 4, we construct outer
solutions of (1.7). We first study an initial value problem of (1.7) with the initial values at $\theta=\frac{\pi}{2}$, then we analyze the asymptotic behaviors of the solutions of this initial value problem near $\theta=0$. Finally, in Section 5 , we match the inner and outer solutions constructed in Sections 3 and 4 to obtain solutions of (1.1). This completes the proof of Theorem 1.1. We leave some computation results in the last section.

## 2. PRELIMINARIES

In this section, we present some known results which will be used in the subsequent sections.

Let $u=u(r)$ be a positive radial solution of (1.1). Using the Emden-Fowler transformation:

$$
\begin{equation*}
u(r)=r^{-\alpha} v(t), \quad t=\ln r \tag{2.1}
\end{equation*}
$$

we see that $v(t)$ satisfies the equation:

$$
\begin{equation*}
v^{(4)}(t)+K_{3} v^{\prime \prime \prime}(t)+K_{2} v^{\prime \prime}(t)+K_{1} v^{\prime}(t)+K_{0} v(t)=v^{p}(t), \quad t \in(-\infty, \infty) \tag{2.2}
\end{equation*}
$$

where the coefficients $K_{0}, K_{1}, K_{2}, K_{3}$ are given in [7]:

$$
\begin{aligned}
K_{0}= & \frac{8}{(p-1)^{4}}\left[(n-2)(n-4)(p-1)^{3}+2\left(n^{2}-10 n+20\right)(p-1)^{2}\right. \\
& -16(n-4)(p-1)+32] \\
K_{1}= & -\frac{2}{(p-1)^{3}}\left[(n-2)(n-4)(p-1)^{3}+4\left(n^{2}-10 n+20\right)(p-1)^{2}\right. \\
& -48(n-4)(p-1)+128] \\
K_{2}= & \frac{1}{(p-1)^{2}}\left[\left(n^{2}-10 n+20\right)(p-1)^{2}-24(n-4)(p-1)+96\right] \\
K_{3}= & \frac{2}{p-1}[(n-4)(p-1)-8] .
\end{aligned}
$$

By direct calculations it is easy to see that $K_{0}=k_{0}$. The characteristic polynomial (linearized at $K_{0}^{\frac{1}{p-1}}$ ) of (2.2) is

$$
\nu \mapsto \nu^{4}+K_{3} \nu^{3}+K_{2} \nu^{2}+K_{1} \nu+(1-p) K_{0}
$$

and the eigenvalues are given by

$$
\begin{aligned}
& \nu_{1}=\frac{N_{1}+\sqrt{N_{2}+4 \sqrt{N_{3}}}}{2(p-1)}, \quad \nu_{2}=\frac{N_{1}-\sqrt{N_{2}+4 \sqrt{N_{3}}}}{2(p-1)}, \\
& \nu_{3}=\frac{N_{1}+\sqrt{N_{2}-4 \sqrt{N_{3}}}}{2(p-1)}, \quad \nu_{4}=\frac{N_{1}-\sqrt{N_{2}-4 \sqrt{N_{3}}}}{2(p-1)}
\end{aligned}
$$

where

$$
\begin{aligned}
N_{1}:= & -(n-4)(p-1)+8, \quad N_{2}:=\left(n^{2}-4 n+8\right)(p-1)^{2} \\
N_{3}:= & (9 n-34)(n-2)(p-1)^{4}+8(3 n-8)(n-6)(p-1)^{3} \\
& +\left(16 n^{2}-288 n+832\right)(p-1)^{2}-128(n-6)(p-1)+256 .
\end{aligned}
$$

Let $\tilde{\nu}_{j}=\nu_{j}-\alpha$ for $j=1,2,3,4$. We have the following proposition ([10]):

Proposition 2.1. For any $n \geq 5$ and $p>\frac{n+4}{n-4}$, we have

$$
\begin{equation*}
\tilde{\nu}_{2}<2-n<0<\tilde{\nu}_{1} . \tag{2.3}
\end{equation*}
$$

(1) For any $5 \leq n \leq 12$ or $n \geq 13$ and $\frac{n+4}{n-4}<p<p_{c}(n)$, we have $\tilde{\nu}_{3}, \tilde{\nu}_{4} \notin \mathbb{R}$ and $\Re\left(\tilde{\nu}_{3}\right)=\Re\left(\tilde{\nu}_{4}\right)=\frac{4-n}{2}<0$.
(2) For any $n \geq 13$ and $p=p_{c}(n)$, we have $\tilde{\nu}_{3}=\tilde{\nu}_{4}=\frac{4-n}{2}$.
(3) For any $n \geq 13$ and $p>p_{c}(n)$, we have

$$
\begin{equation*}
\tilde{\nu}_{2}<4-n<\tilde{\nu}_{4}<\frac{4-n}{2}<\tilde{\nu}_{3}<0<\tilde{\nu}_{1}, \quad \tilde{\nu}_{3}+\tilde{\nu}_{4}=4-n . \tag{2.4}
\end{equation*}
$$

Theorem 2.1. ([7]) The following limits hold:

$$
\begin{equation*}
\lim _{t \rightarrow \infty} v(t)=K_{0}^{1 / p-1}, \quad \lim _{t \rightarrow \infty} v^{(k)}(t)=0 \tag{2.5}
\end{equation*}
$$

for any $k \geq 1$.
Remark: We see that $K_{i}(i=0,1,2,3)$ and $\nu_{j}, \tilde{\nu}_{j}(j=1,2,3,4)$ above depend on $n$ and $p$. In the following, by abuse of notation, we use $K_{i}, \nu_{j}, \tilde{\nu}_{j}$ with the dimension $n$ replaced by $n-1$ and denote $k_{0}=k_{0}(n), k_{1}=k_{1}(n)$.

## 3. InNer solutions

In this section, we construct inner solutions of (1.7).
Let $Q \gg 1$ be a large constant and $\tilde{b}$ be a constant which will be given below. We consider an initial value problem:

$$
\left\{\begin{array}{l}
T_{1} w(\theta)+k_{1} T_{2} w(\theta)+k_{0} w=w^{p}  \tag{3.1}\\
w(0)=Q, \quad w^{\prime}(0)=0, \quad w^{\prime \prime}(0)=(\tilde{b}+\mu) Q^{1+\frac{2}{\alpha}}, \quad w^{\prime \prime \prime}(0)=0
\end{array}\right.
$$

where $\mu>0$ is a small constant. Since $Q \gg 1$, we set $Q=\epsilon^{-\frac{4}{p-1}}\left(:=\epsilon^{-\alpha}\right)$ with $\epsilon>0$ sufficiently small.

Let $w(\theta)=\epsilon^{-\alpha} v\left(\frac{\theta}{\epsilon}\right)$, then we have $v(0)=1, v^{\prime}(0)=0, v^{\prime \prime}(0)=\tilde{b}+\mu, v^{\prime \prime \prime}(0)=0$ and $v(r)$ (for $r=\frac{\theta}{\epsilon}$ ) satisfies the following equation:

$$
\begin{align*}
v^{(4)}(r) & +2(n-2) \epsilon \cot (\epsilon r) v^{\prime \prime \prime}(r)+\left((n-2)(n-4) \frac{\epsilon^{2}}{\sin ^{2}(\epsilon r)}-(n-2)^{2} \epsilon^{2}+k_{1} \epsilon^{2}\right) v^{\prime \prime} \\
& +\left((n-2) k_{1} \epsilon^{3} \cot (\epsilon r)-(n-2)(n-4) \epsilon^{3} \frac{\cot (\epsilon r)}{\sin ^{2}(\epsilon r)}\right) v^{\prime}(r)+k_{0} \epsilon^{4} v(r)=v^{p}(r) \tag{3.2}
\end{align*}
$$

with initial conditions

$$
v(0)=1, \quad v^{\prime}(0)=0, \quad v^{\prime \prime}(0)=\tilde{b}+\mu, \quad v^{\prime \prime \prime}(0)=0 .
$$

For $\epsilon>0$ sufficiently small, we have

$$
\begin{array}{r}
\epsilon \cot (\epsilon r)=\frac{1}{r}-\frac{1}{3} \epsilon^{2} r+\sum_{k=1}^{\infty} l_{k} \epsilon^{2 k+2} r^{2 k+1}, \\
\epsilon^{2} \sin ^{-2}(\epsilon r)=\frac{1}{r^{2}}+\frac{1}{3} \epsilon^{2}+\sum_{k=1}^{\infty} m_{k} \epsilon^{2 k+2} r^{2 k} \\
\epsilon^{3} \cot (\epsilon r) \sin ^{-2}(\epsilon r)=\frac{1}{r^{3}}+\sum_{k=1}^{\infty} n_{k} \epsilon^{2 k+2} r^{2 k-1},
\end{array}
$$

then, (3.2) can be written as the following form

$$
\begin{align*}
& v^{(4)}(r)+\left(\frac{2(n-2)}{r}-\frac{2(n-2)}{3} \epsilon^{2} r+\sum_{k=1}^{\infty} l_{k}^{\prime} \epsilon^{2 k+2} r^{2 k+1}\right) v^{\prime \prime \prime}(r) \\
& \quad+\left(\frac{(n-2)(n-4)}{r^{2}}+\left[\frac{(n-2)(n-4)}{3}-(n-2)^{2}+k_{1}\right] \epsilon^{2}+\sum_{k=1}^{\infty} m_{k}^{\prime} \epsilon^{2 k+2} r^{2 k+1}\right) v^{\prime \prime}(r) \\
& \quad-\left(\frac{(n-2)(n-4)}{r^{3}}-(n-2) k_{1} r^{-1} \epsilon^{2}+\sum_{k=1}^{\infty} n_{k}^{\prime} \epsilon^{2 k+2} r^{2 k-1}\right) v^{\prime}(r)+k_{0} \epsilon^{4} v(r) \\
& =v^{p}(r) \tag{3.3}
\end{align*}
$$

with initial conditions

$$
v(0)=1, \quad v^{\prime \prime}(0)=\tilde{b}+\mu, \quad v^{\prime}(0)=v^{\prime \prime \prime}(0)=0
$$

The first approximation to the solution of $(3.3)$ is the radial solution $v_{0}(r)$ of the problem

$$
\begin{equation*}
\Delta^{2} v=v^{p} \text { in } \mathbb{R}^{n-1}, \quad v(0)=1, \quad v^{\prime}(0)=0, \quad v^{\prime \prime}(0)=\tilde{b}+\mu, \quad v^{\prime \prime \prime}(0)=0 . \tag{3.4}
\end{equation*}
$$

We write $v_{0}=v_{01}+v_{02}$, where $v_{01}$ and $v_{02}$ satisfy the following equations respectively

$$
\begin{equation*}
\Delta^{2} v=v^{p}, \quad v(0)=1, \quad v^{\prime}(0)=0, \quad v^{\prime \prime}(0)=\tilde{b}, \quad v^{\prime \prime \prime}(0)=0 \tag{3.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\Delta^{2} v=v_{0}^{p}-v_{01}^{p}, \quad v(0)=0, \quad v^{\prime}(0)=0, \quad v^{\prime \prime}(0)=\mu, \quad v^{\prime \prime \prime}(0)=0 \tag{3.6}
\end{equation*}
$$

We now choose $\tilde{b}<0$ to be the unique value such that the solution $v_{01}$ is the unique positive radial ground state of (3.5).
Lemma 3.1. Assume that $v_{01}(r)$ and $v_{02}(r)$ are the solutions to (3.5) and (3.6) respectively. For $\frac{n+3}{n-5}<p<p_{c}(n-1)$, there exists $R_{0} \gg 1$ such that for $r \geq R_{0}$, the solution $v_{01}(r)$ satisfies

$$
\begin{equation*}
v_{01}(r)=A_{p} r^{-\alpha}+\frac{a_{0} \cos (\beta \ln r)+b_{0} \sin (\beta \ln r)}{r^{\frac{n-5}{2}}}+O\left(r^{\tilde{\nu}_{2}}\right) \tag{3.7}
\end{equation*}
$$

where $\beta=\frac{\sqrt{4 \sqrt{N_{3}}-N_{2}}}{2(p-1)}\left(\right.$ with $n$ being replaced by $n-1$ in $N_{2}$ and $N_{3}$ ) and $\sqrt{a_{0}^{2}+b_{0}^{2}} \neq$ 0 .

The solution $v_{02}(r)$ satisfies

$$
\begin{equation*}
v_{02}(r)=\mu B_{p} r^{\tilde{\nu}_{1}}+O\left(\mu^{2} r^{\nu_{1}+\tilde{\nu}_{1}}+\mu r^{\tilde{\nu}_{1}+\left(\alpha-\frac{n-5}{2}\right)}\right) \tag{3.8}
\end{equation*}
$$

with $B_{p} \neq 0$ when $\mu=O\left(\frac{1}{r^{\nu_{1}-\sigma}}\right)$ for $r$ in any interval $\left[e^{T}, e^{10 T}\right]$ with $T \gg 1$ and $\sigma=\alpha-\frac{n-5}{2}$.

Proof. The proof of this lemma is divided to two steps. We consider $v_{01}(r)$ in the first step. The main arguments in the proof are similar to those in the proof of Theorem 3.1 of [8].

Using the Emden-Fowler transformation:

$$
\begin{equation*}
v_{01}(r)=r^{-\alpha} v(t), \quad t=\ln r(r>0) \tag{3.9}
\end{equation*}
$$

and letting $v(t)=A_{p}-h(t)$, we see that $h(t)$ satisfies

$$
\begin{equation*}
h^{(4)}(t)+K_{3} h^{\prime \prime \prime}(t)+K_{2} h^{\prime \prime}(t)+K_{1} h^{\prime}(t)+(1-p) K_{0} h(t)+O\left(h^{2}\right)=0, \quad t>1 . \tag{3.10}
\end{equation*}
$$

Note that $r^{\alpha} v_{01}(r) \rightarrow A_{p}$ as $r \rightarrow \infty$ and hence $h(t) \rightarrow 0$ as $t \rightarrow \infty$. It follows from Proposition 2.1 that $\tilde{\nu}_{3}, \tilde{\nu}_{4} \notin \mathbb{R}$ and $\Re\left(\tilde{\nu}_{3}\right)=\Re\left(\tilde{\nu}_{4}\right)=\frac{5-n}{2}<0$ and $\tilde{\nu}_{2}<3-n<$ $0<\tilde{\nu}_{1}$ provided $\frac{n+3}{n-5}<p<p_{c}(n-1)$. Let $\nu_{3}=\sigma+i \beta$, where $\beta=\frac{\sqrt{4 \sqrt{N_{3}}-N_{2}}}{2(p-1)}$ and $\sigma=-\frac{n-5}{2}+\alpha<0$ for $p>\frac{n+3}{n-5}$.

We can write (3.10) as

$$
\begin{equation*}
\left(\partial_{t}-\nu_{4}\right)\left(\partial_{t}-\nu_{3}\right)\left(\partial_{t}-\nu_{2}\right)\left(\partial_{t}-\nu_{1}\right) h(t)=H(h(t)), \tag{3.11}
\end{equation*}
$$

where $H(h(t))=O\left(h^{2}\right)$. We claim that for any $T \gg 1$, there exist constants $A_{i}$ and $B_{i}(i=1,2,3,4)$ such that

$$
\begin{aligned}
h(t)= & A_{1} e^{\sigma t} \cos \beta t+A_{2} e^{\sigma t} \sin \beta t+A_{3} e^{\nu_{2} t}+A_{4} e^{\nu_{1} t} \\
& +B_{1} \int_{T}^{t} e^{\sigma(t-s)} \sin \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{2} \int_{T}^{t} e^{\sigma(t-s)} \cos \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{3} \int_{T}^{t} e^{\nu_{2}(t-s)} H(h(s)) \mathrm{d} s \\
& +B_{4} \int_{T}^{t} e^{\nu_{1}(t-s)} H(h(s)) \mathrm{d} s .
\end{aligned}
$$

Moreover, each $A_{i}$ depends on $T$ and $\nu_{i}(i=1,2,3,4)$, while each $B_{i}$ depends only on $\nu_{i}(i=1,2,3,4)$. In fact, it follows from (3.11) and the ODE theory of second order (see [13]) that

$$
\begin{align*}
\left(\partial_{t}-\nu_{2}\right)\left(\partial_{t}-\nu_{1}\right) h(t)= & A_{1}^{\prime} e^{\sigma t} \cos \beta t+A_{2}^{\prime} e^{\sigma t} \sin \beta t \\
& +\frac{1}{\beta} \int_{T}^{t} e^{\sigma(t-s)} \sin \beta(t-s) H(h(s)) \mathrm{d} s \tag{3.12}
\end{align*}
$$

where $A_{1}^{\prime}$ and $A_{2}^{\prime}$ are constants depending on $T, \nu_{3}$ and $\nu_{4}$. Multiplying both sides of (3.12) by $e^{-\nu_{2} t}$ and integrating it from $T$ to $t$, we obtain that

$$
\begin{aligned}
\left(\partial_{t}-\nu_{1}\right) h(t)= & A_{3}^{\prime} e^{\nu_{2} t}+\int_{T}^{t} e^{\nu_{2}(t-s)}\left(A_{1}^{\prime} e^{\sigma s} \cos \beta s+A_{2}^{\prime} e^{\sigma s} \sin \beta s\right) \mathrm{d} s \\
& +\frac{1}{\beta} \int_{T}^{t} e^{\nu_{2}(t-s)} \int_{T}^{s} e^{\sigma(s-\xi)} \sin \beta(s-\xi) H(h(\xi)) \mathrm{d} \xi \mathrm{~d} s
\end{aligned}
$$

We now switch the order of integrations and find that

$$
\begin{aligned}
\left(\partial_{t}-\nu_{1}\right) h(t)= & A_{1}^{\prime \prime} e^{\sigma t} \cos \beta t+A_{2}^{\prime \prime} e^{\sigma t} \sin \beta t+A_{3}^{\prime \prime} e^{\nu_{2} t} \\
& +B_{1}^{\prime} \int_{T}^{t} e^{\sigma(t-s)} \sin \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{2}^{\prime} \int_{T}^{t} e^{\sigma(t-s)} \cos \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{3}^{\prime} \int_{T}^{t} e^{\nu_{2}(t-s)} H(h(s)) \mathrm{d} s
\end{aligned}
$$

where $A_{1}^{\prime \prime}, A_{2}^{\prime \prime}$ and $A_{3}^{\prime \prime}$ depend on $T, \nu_{i}(i=2,3,4), B_{i}^{\prime}(i=1,2,3)$ depend only on $\nu_{i}(i=2,3,4)$. Repeating the same argument once again, we obtain our claim. We can also write $h(t)$ as

$$
\begin{aligned}
h(t)= & A_{1} e^{\sigma t} \cos \beta t+A_{2} e^{\sigma t} \sin \beta t+A_{3} e^{\nu_{2} t}+M_{4} e^{\nu_{1} t} \\
& +B_{1} \int_{T}^{t} e^{\sigma(t-s)} \sin \beta(t-s) H(h(s)) \mathrm{d} s+B_{2} \int_{T}^{t} e^{\sigma(t-s)} \cos \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{3} \int_{T}^{t} e^{\nu_{2}(t-s)} H(h(s)) d s-B_{4} \int_{t}^{\infty} e^{\nu_{1}(t-s)} H(h(s)) \mathrm{d} s
\end{aligned}
$$

by using the fact that $\int_{T}^{t}=\int_{T}^{\infty}-\int_{t}^{\infty}$. Since $h(t) \rightarrow 0$ as $t \rightarrow \infty$, we have $M_{4}=0$ (note $\nu_{1}>0$ ). Setting

$$
h_{1}(t)=A_{1} e^{\sigma t} \cos \beta t+A_{2} e^{\sigma t} \sin \beta t+A_{3} e^{\nu_{2} t}
$$

and

$$
\begin{aligned}
h_{2}(t)= & B_{1} \int_{T}^{t} e^{\sigma(t-s)} \sin \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{2} \int_{T}^{t} e^{\sigma(t-s)} \cos \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{3} \int_{T}^{t} e^{\nu_{2}(t-s)} H(h(s)) \mathrm{d} s \\
& -B_{4} \int_{t}^{\infty} e^{\nu_{1}(t-s)} H(h(s)) \mathrm{d} s,
\end{aligned}
$$

we see from the fact $H(h(t))=O\left(h^{2}(t)\right)$ that

$$
\begin{equation*}
\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\tilde{h}_{2}(t)\right], \tag{3.13}
\end{equation*}
$$

where $C>0$ is independent of $T$ and

$$
\begin{aligned}
& \tilde{h}_{1}(t)=\max \left\{\int_{T}^{t} e^{\sigma(t-s)}\left|h_{1}(s)\right|^{2} \mathrm{~d} s, \int_{T}^{t} e^{\nu_{2}(t-s)}\left|h_{1}(s)\right|^{2} \mathrm{~d} s, \int_{t}^{\infty} e^{\nu_{1}(t-s)}\left|h_{1}(s)\right|^{2} \mathrm{~d} s\right\}, \\
& \tilde{h}_{2}(t)=\max \left\{\int_{T}^{t} e^{\sigma(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s, \int_{T}^{t} e^{\nu_{2}(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s, \int_{t}^{\infty} e^{\nu_{1}(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s\right\} .
\end{aligned}
$$

We now show

$$
\begin{equation*}
\left|h_{2}(t)\right|=o\left(e^{\sigma t}\right) \tag{3.14}
\end{equation*}
$$

There are three cases to be considered:
(1) $\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\int_{T}^{t} e^{\sigma(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s\right]$,
(2) $\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\int_{T}^{t} e^{\nu_{2}(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s\right]$,
(3) $\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\int_{t}^{\infty} e^{\nu_{1}(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s\right]$.

We only consider (1) and (3). Case (2) can be discussed similarly. For case (1), we have

$$
\begin{equation*}
\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\int_{T}^{t} e^{\sigma(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s\right] \tag{3.15}
\end{equation*}
$$

Thus,

$$
\begin{equation*}
\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\max _{t \geq T}\left|h_{2}(t)\right| \int_{T}^{t} e^{\sigma(t-s)}\left|h_{2}(s)\right| \mathrm{d} s\right] \tag{3.16}
\end{equation*}
$$

Let $m(t)=\int_{T}^{t} e^{-\sigma s}\left|h_{2}(s)\right|$, then, it is seen from (3.16) that

$$
\begin{equation*}
m^{\prime}(t) \leq C \tilde{h}_{1}(t) e^{-\sigma t}+C \max _{t \geq T}\left|h_{2}(t)\right| m(t) \tag{3.17}
\end{equation*}
$$

For any $\epsilon>0$ sufficiently small, we can choose $T$ sufficiently large such that $0<$ $d_{T}:=C \max _{t \geq T}\left|h_{2}(t)\right|<\epsilon$. It follows from (3.17) that

$$
\begin{equation*}
m(t) \leq C e^{d_{T} t} \int_{T}^{t} \tilde{h}_{1}(s) e^{-\sigma s} e^{-d_{T} s} \mathrm{~d} s \tag{3.18}
\end{equation*}
$$

Substituting $m(t)$ in (3.18) into (3.16), we see that

$$
\begin{equation*}
\left|h_{2}(t)\right| \leq C \tilde{h}_{1}(t)+C d_{T} e^{\left(\sigma+d_{T}\right) t} \int_{T}^{t} \tilde{h}_{1}(s) e^{-\sigma s} e^{-d_{T} s} d s \tag{3.19}
\end{equation*}
$$

Note that $\sigma+d_{T}<0$ for $T$ sufficiently large. We also know that $\nu_{2}<\sigma$, which implies $\tilde{h}_{1}(t)=o\left(e^{\sigma t}\right)$. On the other hand, from (3.19) we can obtain that $\left|h_{2}(t)\right|=$ $o\left(e^{\left(\sigma+d_{T}\right) t}\right)$. Substituting these into (3.15), we eventually have

$$
\begin{equation*}
\left|h_{2}(t)\right|=o\left(e^{\sigma t}\right) \tag{3.20}
\end{equation*}
$$

For case (3), we have

$$
\begin{equation*}
\left|h_{2}(t)\right| \leq C\left[\tilde{h}_{1}(t)+\int_{t}^{\infty} e^{\nu_{1}(t-s)}\left|h_{2}(s)\right|^{2} \mathrm{~d} s\right] \tag{3.21}
\end{equation*}
$$

Thus,

$$
\begin{equation*}
\left|h_{2}(t)\right| \leq C \tilde{h}_{1}(t)+C \max _{t \geq T}\left|h_{2}(t)\right| \int_{t}^{\infty} e^{\nu_{1}(t-s)}\left|h_{2}(s)\right| \mathrm{d} s \tag{3.22}
\end{equation*}
$$

Let $l(t)=\int_{t}^{\infty} e^{-\nu_{1} s}\left|h_{2}(s)\right| \mathrm{d} s$, it is seen from (3.22) that

$$
\begin{equation*}
-l^{\prime}(t) \leq C \tilde{h}_{1}(t) e^{-\nu_{1} t}+d_{T} l(t) \tag{3.23}
\end{equation*}
$$

It follows from (3.23) that

$$
\begin{equation*}
l(s) \leq C e^{-d_{T} t} \int_{t}^{\infty} \tilde{h}_{1}(s) e^{-\nu_{1} s} e^{d_{T} s} \mathrm{~d} s \tag{3.24}
\end{equation*}
$$

Since $\tilde{h}_{1}(t)=o\left(e^{\sigma t}\right)$, we obtain from (3.24) that

$$
l(s)=o\left(e^{\left(\sigma-\nu_{1}\right) t}\right)
$$

Substituting this into (3.22), we also have

$$
\left|h_{2}(t)\right|=o\left(e^{\sigma t}\right)
$$

We now write $h(t)$ as

$$
\begin{aligned}
h(t)= & M_{1} e^{\sigma t} \cos \beta t+M_{2} e^{\sigma t} \sin \beta t+A_{3} e^{\nu_{2} t} \\
& -B_{1} \int_{t}^{\infty} e^{\sigma(t-s)} \sin \beta(t-s) H(h(s)) \mathrm{d} s \\
& -B_{2} \int_{t}^{\infty} e^{\sigma(t-s)} \cos \beta(t-s) H(h(s)) \mathrm{d} s \\
& +B_{3} \int_{T}^{t} e^{\nu_{2}(t-s)} H(h(s)) \mathrm{d} s \\
& -B_{4} \int_{t}^{\infty} e^{\nu_{1}(t-s)} H(h(s)) \mathrm{d} s
\end{aligned}
$$

Then, it follows from the facts $H(h(t))=O\left(h^{2}(t)\right), h_{1}(t)=O\left(e^{\sigma t}\right), h_{2}(t)=o\left(e^{\sigma t}\right)$ and $\nu_{2}<2 \sigma$ that

$$
\begin{equation*}
h(t)=M_{1} e^{\sigma t} \cos (\beta t)+M_{2} e^{\sigma t} \sin (\beta t)+A_{3} e^{\nu_{2} t}+O\left(e^{2 \sigma t}\right) \tag{3.25}
\end{equation*}
$$

This implies that (3.7) holds for some $a_{0}$ and $b_{0}$. By a similar argument which was used in the proof of $\left[10\right.$, Theorem 3.3], we can show $a_{0}^{2}+b_{0}^{2} \neq 0$. This completes the proof of the first step.

We now proceed the second step. Setting $v_{02}=\mu \tilde{v}_{02}$, we see that $\tilde{v}_{02}(r)$ satisfies the following equation:

$$
\begin{equation*}
\Delta^{2} \tilde{v}_{02}-p v_{01}^{p-1} \tilde{v}_{02}=\mu^{-1}\left[\left(v_{01}+\mu \tilde{v}_{02}\right)^{p}-v_{01}^{p}-p \mu v_{01}^{p-1} \tilde{v}_{02}\right] \tag{3.26}
\end{equation*}
$$

with initial conditions

$$
\tilde{v}_{02}(0)=0, \quad \tilde{v}_{02}^{\prime}(0)=0, \quad \tilde{v}_{02}^{\prime \prime}(0)=1, \quad \tilde{v}_{02}^{\prime \prime \prime}(0)=0
$$

Using the Emden-Fowler transformation:

$$
\tilde{v}_{02}(r)=r^{-\alpha} \hat{v}(t), \quad t=\ln r(r>0)
$$

and the expression obtained for $v_{01}(r)$, we see that $\hat{v}(t)$ satisfies

$$
\begin{equation*}
\hat{v}^{(4)}+K_{3} \hat{v}^{\prime \prime \prime}+K_{2} \hat{v}^{\prime \prime}+K_{1} \hat{v}^{\prime}+(1-p) K_{0} \hat{v}=f(r, \mu, \hat{v}) \tag{3.27}
\end{equation*}
$$

where

$$
f(r, \mu, \hat{v})=O\left(\left[\mu \hat{v}+r^{\left(\alpha-\frac{n-5}{2}\right)}\right]\right) \hat{v}
$$

provided that $\mu \hat{v}=o(1)$ for $t$ sufficiently large. It follows from (3.27) that

$$
\begin{aligned}
\hat{v}(t)= & \hat{A}_{1} e^{\sigma t} \cos \beta t+\hat{A}_{2} e^{\sigma t} \sin \beta t+\hat{A}_{3} e^{\nu_{2} t}+\hat{A}_{4} e^{\nu_{1} t} \\
& +\hat{B}_{1} \int_{T}^{t} e^{\sigma(t-s)} \sin \beta(t-s) f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
& +\hat{B}_{2} \int_{T}^{t} e^{\sigma(t-s)} \cos \beta(t-s) f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
& +\hat{B}_{3} \int_{T}^{t} e^{\nu_{2}(t-s)} f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
& +\hat{B}_{4} \int_{T}^{t} e^{\nu_{1}(t-s)} f(r, \mu, \hat{v}(s)) \mathrm{d} s
\end{aligned}
$$

where $\hat{A}_{i}=\hat{A}_{i}\left(T, \nu_{1}, \nu_{2}, \nu_{3}, \nu_{4}\right)(i=1,2,3,4)$ and $\hat{B}_{i}=\hat{B}_{i}\left(\nu_{1}, \nu_{2}, \nu_{3}, \nu_{4}\right)$. We first show that $\tilde{v}_{02}$ is strictly increasing in $(0, \infty)$. Using the initial values, we can find $R \in(0, \infty)$ such that $\tilde{v}_{02}(r)>0$ for $r \in(0, R)$. Writing (3.26) as

$$
\mu \Delta^{2} \tilde{v}_{02}=\left(v_{01}+\mu \tilde{v}_{02}\right)^{p}-v_{01}^{p}
$$

we obtain that $\left(\Delta \tilde{v}_{02}\right)^{\prime}>0$, and hence $\Delta \tilde{v}_{02}>\Delta \tilde{v}_{02}(0)=n$ for $r \in(0, R)$, which implies that $\left(\tilde{v}_{02}\right)^{\prime}(r)>0$ for $r \in(0, R)$. Moreover, we can deduce that $R=\infty$ and $\tilde{v}_{02}^{\prime}(r)>0$ for $r \in(0, \infty)$. Therefore, $\hat{v}$ is increasing in $(0, \infty)$. Next, we claim that $\hat{A}_{4} \neq 0$ for any $T \gg 1$ sufficiently large. Indeed, for $t \in[T, 10 T]$,

$$
\begin{aligned}
e^{-\nu_{1} t} \hat{v}(t)= & \hat{A}_{4}+\tilde{g}(t) \\
& +\hat{B}_{1} e^{\left(\sigma-\nu_{1}\right) t} \int_{T}^{t} e^{-\sigma s} \sin \beta(t-s) f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
& +\hat{B}_{2} e^{\left(\sigma-\nu_{1}\right) t} \int_{T}^{t} e^{-\sigma s} \cos \beta(t-s) f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
& +\hat{B}_{3} e^{\left(\nu_{2}-\nu_{1}\right) t} \int_{T}^{t} e^{-\nu_{2} s} f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
& +\hat{B}_{4} \int_{T}^{t} e^{-\nu_{1} s} f(r, \mu, \hat{v}(s)) \mathrm{d} s \\
\leq & \left|\hat{A}_{4}\right|+|\tilde{g}(t)|+\left(\Sigma_{j=1}^{4}\left|\hat{B}_{j}\right|\right) \max _{t \in[T, 10 T]}\left[\mu \hat{v}+e^{\left(\alpha-\frac{n-5}{2}\right) t}\right] \int_{T}^{t} e^{-\nu_{1} s} \hat{v}(s),
\end{aligned}
$$

where

$$
\tilde{g}(t)=\hat{A}_{1} e^{\left(\sigma-\nu_{1}\right) t} \cos \beta t+\hat{A}_{2} e^{\left(\sigma-\nu_{1}\right) t} \sin \beta t+\hat{A}_{3} e^{\left(\nu_{2}-\nu_{1}\right) t}
$$

Since

$$
\left(\Sigma_{j=1}^{4}\left|\hat{B}_{j}\right|\right) \max _{t \in[T, 10 T]}\left[\mu \hat{v}+e^{\left(\alpha-\frac{n-5}{2}\right) t}\right]=\tau=o(1)
$$

we have

$$
\begin{equation*}
e^{-\nu_{1} t} \hat{v}(t) \leq\left|\hat{A}_{4}\right|+|\tilde{g}(t)|+\tau \int_{T}^{t} e^{-\nu_{1} s} \hat{v}(s) \mathrm{d} s \tag{3.28}
\end{equation*}
$$

Let $\ell(t)=\int_{T}^{t} e^{-\nu_{1} s} \hat{v}(s) \mathrm{d} s$. We see that

$$
\begin{equation*}
\left(e^{-\tau t} \ell(t)\right)^{\prime} \leq\left(\left|\hat{A}_{4}\right|+|\tilde{g}(t)|\right) e^{-\tau t} \tag{3.29}
\end{equation*}
$$

Integrating (3.29) in $[T, t]$, we obtain that

$$
\ell(t) \leq \frac{\left|\hat{A}_{4}\right|+\max _{t \in[T, 10 T]}|\tilde{g}(t)|}{\tau} e^{\tau(t-T)}
$$

If we choose $\tau(t-T) \leq C$ for $t \in[T, 10 T]$, i.e., $\tau=O\left(\frac{1}{T}\right)$, we see that

$$
\begin{equation*}
\ell(t) \leq \frac{\left(\left|\hat{A}_{4}\right|+\max _{t \in[T, 10 T]}|\tilde{g}(t)|\right) C}{\tau} \tag{3.30}
\end{equation*}
$$

Substituting this into (3.28), we have

$$
\begin{equation*}
e^{-\nu_{1} t} \hat{v}(t) \leq\left|\hat{A}_{4}\right|(1+C)+|\tilde{g}(t)|+C \max _{t \in[T, 10 T]}|\tilde{g}(t)| \tag{3.31}
\end{equation*}
$$

Suppose $\hat{A}_{4}=0$. We see from (3.31) and the expression of $|\tilde{g}(t)|$ that

$$
\hat{v}(t)=o(1), \quad \forall t \in[T, 10 T] .
$$

This contradicts the fact that $\hat{v}$ is increasing in $(0, \infty)$. Therefore, $\hat{A}_{4} \neq 0$ and our claim holds. Moreover, it is known from (3.31) and the expression of $\hat{v}(t)$ that

$$
\begin{equation*}
\hat{v}(t)=B_{p} e^{\nu_{1} t}+O\left(\mu e^{2 \nu_{1} t}+e^{\left(\sigma+\nu_{1}\right) t}\right) \tag{3.32}
\end{equation*}
$$

with $B_{p} \neq 0$ and $\mu=O\left(e^{\left(-\nu_{1}+\sigma\right) t}\right)$. Therefore,

$$
v_{02}(r)=\mu B_{p} r^{\tilde{\nu}_{1}}+O\left(\mu^{2} r^{\nu_{1}+\tilde{\nu}_{1}}+\mu r^{\tilde{\nu}_{1}+\sigma}\right)
$$

with $B_{p} \neq 0$ and $\mu=O\left(\frac{1}{r^{\nu}-\sigma}\right)$. This completes the proof of this lemma.
Lemma 3.2. Let p satisfy the conditions of Lemma 3.1 and $v_{1}(r)$ be the unique solution of the following equation

$$
\left\{\begin{array}{l}
v_{1}^{(4)}(r)+\frac{2(n-2)}{r} v_{1}^{\prime \prime \prime}(r)+\frac{(n-2)(n-4)}{r^{2}} v_{1}^{\prime \prime}(r)-\frac{(n-2)(n-4)}{r^{3}} v_{1}^{\prime}(r)-\frac{2(n-2)}{3} r v_{0}^{\prime \prime \prime}(r)  \tag{3.33}\\
\quad+\left(\frac{(n-2)(n-4)}{3}-(n-2)^{2}+k_{1}\right) v_{0}^{\prime \prime}(r)+\frac{(n-2) k_{1}}{r} v_{0}^{\prime}(r)=p v_{0}^{p-1}(r) v_{1}(r), \\
v_{1}(0)=0, v_{1}^{\prime}(0)=0, v_{1}^{\prime \prime}(0)=0, v_{1}^{\prime \prime \prime}(0)=0 .
\end{array}\right.
$$

Then for $r \in\left[e^{T}, e^{10 T}\right]$ with $T \gg 1$ and $\mu=O\left(\frac{1}{r^{\nu_{1}-\sigma}}\right)$,

$$
\begin{align*}
v_{1}(r)= & C_{p} r^{2-\alpha}+r^{2-\frac{n-5}{2}}\left(a_{1} \cos (\beta \ln r)+b_{1} \sin (\beta \ln r)\right)+\mu D_{p} r^{2+\tilde{\nu}_{1}} \\
& +O\left(\mu^{2} r^{\tilde{\nu}_{1}+\nu_{1}+2}+\mu r^{\tilde{\nu}_{1}+\sigma+2}\right)+o\left(r^{2-\frac{n-5}{2}}\right), \tag{3.34}
\end{align*}
$$

where $C_{p}$ satisfies

$$
\begin{equation*}
E_{1} C_{p}-p A_{p}^{p-1} C_{p}=F_{1} A_{p} \tag{3.35}
\end{equation*}
$$

and

$$
\begin{aligned}
E_{1}= & (1+\alpha)(1-\alpha)(2-\alpha) \alpha-2(n-2)(2-\alpha)(1-\alpha) \alpha-(n-2)(n-4)(2-\alpha) \\
& +(n-2)(n-4)(2-\alpha)(1-\alpha) \\
F_{1}= & \left((n-2)^{2}-k_{1}-\frac{(n-2)(n-4)}{3}\right) \alpha(\alpha+1)-\frac{2(n-2)}{3} \alpha(\alpha+1)(\alpha+2) \\
& +k_{1}(n-2) \alpha
\end{aligned}
$$

$D_{p}$ satisfies

$$
\begin{equation*}
E_{2} D_{p}=F_{2} B_{p} \tag{3.36}
\end{equation*}
$$

where

$$
E_{2}=\left(2+\tilde{\nu}_{1}\right)\left(\tilde{\nu}_{1}+n-1\right)\left(\tilde{\nu}_{1}+n-3\right) \tilde{\nu}_{1}-p A_{p}^{p-1}
$$

and

$$
\begin{aligned}
F_{2}= & \frac{2(n-2)\left(\tilde{\nu}_{1}-1\right)\left(\tilde{\nu}_{1}-2\right) \tilde{\nu}_{1}}{3}+\left((n-2)^{2}-k_{1}-\frac{(n-2)(n-4)}{3}\right)\left(\tilde{\nu}_{1}-1\right) \tilde{\nu}_{1} \\
& -k_{1}(n-2) \tilde{\nu}_{1}+p(p-1) A_{p}^{p-2} C_{p},
\end{aligned}
$$

$\left(a_{1}, b_{1}\right)$ is the solution of

$$
\left\{\begin{array}{l}
A a_{1}-B b_{1}=G \\
B a_{1}+A b_{1}=H
\end{array}\right.
$$

where

$$
\begin{gathered}
A=\frac{n^{4}-12 n^{3}+14 n^{2}+132 n-135}{16}-p A_{p}^{p-1}+\left(\frac{n^{2}-6 n-35}{2}\right) \beta^{2}+\beta^{4} \\
B=\left(2 n^{2}-12 n-6\right) \beta+8 \beta^{3}
\end{gathered}
$$

$$
\begin{aligned}
G= & p(p-1) A_{p}^{p-2} C_{p} a_{0}+\frac{n^{4}-11 n^{3}+41 n^{2}-61 n+30}{12} a_{0}+\frac{n^{2}-6 n+5}{4} k_{1} a_{0} \\
& +\frac{4 n^{2}+3 n-n^{3}-14}{6} b_{0} \beta-2 k_{1} b_{0} \beta+\frac{n^{2}-9 n+14}{3} a_{0} \beta^{2}+a_{0} k_{1} \beta^{2} \\
& -\frac{2(n-2)}{3} b_{0} \beta^{3},
\end{aligned}
$$

and

$$
\begin{aligned}
H= & p(p-1) A_{p}^{p-2} C_{p} b_{0}+\frac{n^{4}-11 n^{3}+41 n^{2}-61 n+30}{12} b_{0}+\frac{n^{2}-6 n+5}{4} k_{1} b_{0} \\
& -\frac{4 n^{2}+3 n-n^{3}-14}{6} a_{0} \beta+2 k_{1} a_{0} \beta+\frac{n^{2}-9 n+14}{3} b_{0} \beta^{2}+b_{0} k_{1} \beta^{2} \\
& +\frac{2(n-2)}{3} a_{0} \beta^{3} .
\end{aligned}
$$

Remark: We need to show $E_{2} \neq 0$ and the following 2 by 2 matrix is invertible,

$$
K=\left[\begin{array}{cc}
A, & -B \\
B, & A
\end{array}\right]
$$

This will be proved in last section.
Proof. Let

$$
\begin{aligned}
v_{1}(r)= & C_{p} r^{2-\alpha}+\tilde{f}(r) r^{2-\frac{n-5}{2}}+\mu D_{p} r^{2+\tilde{\nu}_{1}}+o\left(r^{2-\frac{n-5}{2}}\right) \\
& +O\left(\mu^{2} r^{\tilde{\nu}_{1}+\nu_{1}+2}+\mu r^{\tilde{\nu}_{1}+\sigma+2}\right)
\end{aligned}
$$

where

$$
\tilde{f}(r)=a_{1} \cos (\beta \ln r)+b_{1} \sin (\beta \ln r)
$$

Using the expression in (3.7) and (3.8), we can get (3.34) by direct calculations.
Furthermore, we can obtain the following proposition.
Proposition 3.1. Let $\frac{n+3}{n-5}<p<p_{c}(n-1)$ and $v(r)$ be a solution of (3.2). Then for $\epsilon>0$ sufficiently small,

$$
v(r)=v_{0}(r)+\sum_{k=1}^{\infty} \epsilon^{2 k} v_{k}(r)
$$

Moreover, for $r \in\left[e^{T}, e^{10 T}\right]$ with $T \gg 1$ and, $\mu=o\left(\frac{1}{r^{\nu_{1}-\sigma}}\right)$,

$$
\begin{align*}
v_{k}(r)= & \sum_{j=1}^{k} d_{j}^{k} r^{2 j-\alpha}+\sum_{j=1}^{k} e_{j}^{k} r^{2 j-\frac{n-5}{2}} \sin \left(\beta \ln r+E_{j}^{k}\right)+\sum_{j=1}^{k} \mu f_{j}^{k} r^{2 j+\tilde{\nu}_{1}} \\
& +O\left(\mu^{2} r^{\tilde{\nu}_{1}+\nu_{1}+2 k}+\mu r^{\tilde{\nu}_{1}+\sigma+2 k}\right)+o\left(r^{2 k-\frac{n-5}{2}}\right) \tag{3.37}
\end{align*}
$$

where $d_{j}^{k}, e_{j}^{k}, f_{j}^{k}, E_{j}^{k}(j=1,2, \cdots, k)$ are constants. Moreover,

$$
d_{1}^{1}=C_{p}, e_{1}^{1}=\sqrt{a_{1}^{2}+b_{1}^{2}}, f_{1}^{1}=D_{p}, \sin E_{1}^{1}=\frac{a_{1}}{e_{1}^{1}}, \cos E_{1}^{1}=\frac{b_{1}}{e_{1}^{1}}
$$

where $C_{p}, a_{1}, b_{1}, D_{p}$ are given in Lemma 3.2.

Proof. Using Taylor's expansion of $v^{p}$ and the expressions of $v_{0}(r), v_{1}(r), \cdots, v_{k-1}(r)$, we can obtain this proposition by the induction argument and direct calculations. Note that

$$
O\left(r^{2-\frac{n-5}{2}}\right)=o\left(r^{2-\alpha}\right)
$$

Theorem 3.1. Let $\frac{n+3}{n-5}<p<p_{c}(n-1)$ and $w_{\epsilon, \mu}^{i n n}(\theta)$ be the solution of (1.7) with

$$
w(0)=\epsilon^{-\alpha}, \quad w_{\theta}(0)=0, \quad w_{\theta \theta}(0)=(\tilde{b}+\mu) \epsilon^{-\alpha-2}, \quad w_{\theta \theta \theta}(0)=0
$$

Then for any sufficiently small $\epsilon>0, \frac{\theta}{\epsilon} \in\left[e^{T}, e^{10 T}\right]$ with $T \gg 1$ and $\mu=$ $O\left(\left(\frac{\epsilon}{\theta}\right)^{\nu_{1}-\sigma}\right)$,

$$
\begin{aligned}
w_{\epsilon, \mu}^{i n n}(\theta)= & \frac{A_{p}}{\theta^{\alpha}}+\frac{C_{p}}{\theta^{\alpha-2}}+B_{p} \mu \epsilon^{-\nu_{1}} \theta^{\tilde{\nu}_{1}}+\sum_{k=2}^{\infty} \sum_{j=1}^{k} d_{j}^{k} \epsilon^{2(k-j)} \theta^{2 j-\alpha} \\
& +\epsilon^{\frac{n-5}{2}-\alpha}\left[\frac{a_{0} \cos \left(\beta \ln \frac{\theta}{\epsilon}\right)+b_{0} \sin \left(\beta \ln \frac{\theta}{\epsilon}\right)}{\theta^{\frac{n-5}{2}}}+\frac{a_{1} \cos \left(\beta \ln \frac{\theta}{\epsilon}\right)+b_{1} \sin \left(\beta \ln \frac{\theta}{\epsilon}\right)}{\theta^{\frac{n-5}{2}-2}}\right. \\
& \left.+\sum_{k=2}^{\infty}\left(\sum_{j=1}^{k} e_{j}^{k} \epsilon^{2(k-j)} \theta^{2 j-\frac{n-5}{2}} \sin \left(\beta \ln \frac{\theta}{\epsilon}+E_{j}^{k}\right)+o\left(\theta^{2 k-\frac{n-5}{2}}\right)\right)+O\left(\theta^{2-\frac{n-5}{2}}\right)\right] \\
& +\epsilon^{-\alpha} \sum_{k=1}^{\infty}\left[\sum_{j=1}^{k}\left(\mu f_{j}^{k} \epsilon^{2 k-2 j-\tilde{\nu}_{1}} \theta^{2 j+\tilde{\nu}_{1}}\right)+O\left(\mu^{2} \theta^{\tilde{\nu}_{1}+\nu_{1}+2 k} \epsilon^{-\tilde{\nu}_{1}-\nu_{1}}\right.\right. \\
& \left.\left.+\mu \theta^{\tilde{\nu}_{1}+\sigma+2 k} \epsilon^{-\tilde{\nu}_{1}-\sigma}\right)\right]+O\left(\mu^{2}\left(\frac{\theta}{\epsilon}\right)^{\tilde{\nu}_{1}+\nu_{1}}+\mu\left(\frac{\theta}{\epsilon}\right)^{\tilde{\nu}_{1}+\sigma}\right) .
\end{aligned}
$$

Proof. This is a direct consequence of Proposition 3.1 by setting $r=\frac{\theta}{\epsilon}$.

We now obtain the following lemmas which will be useful in the following proofs.
Lemma 3.3. Let $\frac{n+3}{n-5}<p<p_{c}(n-1)$ and

$$
v(Q, \mu, \theta)=Q v_{0}\left(Q^{\frac{p-1}{4}} \theta\right)
$$

Then for $Q^{\frac{p-1}{4}} \theta \in\left[e^{T}, e^{10 T}\right]$ with $T \gg 1 ; \mu=O\left(\frac{1}{\left(Q^{(p-1) / 4} \theta\right)^{\nu_{1}-\sigma}}\right)$ and $n=0,1,2$, $v(Q, \mu, \theta)$ satisfies

$$
\begin{aligned}
\frac{\partial^{n}}{\partial Q^{n}}(v(Q, \mu, \theta))= & \frac{\partial^{n}}{\partial Q^{n}}\left(\frac{A_{p}}{\theta^{\alpha}}\right) \\
& +\frac{\partial^{n}}{\partial Q^{n}}\left\{C \theta^{-\frac{n-5}{2}} Q^{-\left(\frac{(p-1)(n-5)}{8}-1\right)} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)+\kappa\right)\right\} \\
& +Q^{\frac{\tilde{\nu}_{2}}{\alpha}+1-n} O\left(\theta^{\tilde{\nu}_{2}}\right)+\mu B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}} \\
& +O\left(\mu^{2} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}+\nu_{1}}+\mu Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1-n} \theta^{\sigma+\tilde{\nu}_{1}}\right),
\end{aligned}
$$

$$
\begin{aligned}
& \frac{\partial^{n}}{\partial Q^{n}}\left(v_{\theta}^{\prime}(Q, \mu, \theta)\right)=\frac{\partial^{n}}{\partial Q^{n}}\left(-\alpha \frac{A_{p}}{\theta^{\alpha+1}}\right) \\
& +\frac{\partial^{n+1}}{\partial Q^{n} \partial \theta}\left\{C \theta^{-\frac{n-5}{2}} Q^{-\left(\frac{(p-1)(n-5)}{8}-1\right)} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)+\kappa\right)\right\} \\
& +Q^{\frac{\tilde{\nu}_{2}}{\alpha}+1-n} O\left(\theta^{\tilde{\nu}_{2}-1}\right)+\mu \tilde{\nu}_{1} B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}-1} \\
& +O\left(\mu^{2} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}+\nu_{1}-1}+\mu Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1-n} \theta^{\sigma+\tilde{\nu}_{1}-1}\right), \\
& \frac{\partial^{n}}{\partial Q^{n}}\left(\frac{\partial^{2}}{\partial \theta^{2}} v(Q, \mu, \theta)\right)=\frac{\partial^{n}}{\partial Q^{n}}\left(\alpha(\alpha+1) \frac{A_{p}}{\theta^{\alpha+2}}\right) \\
& +\frac{\partial^{n+2}}{\partial Q^{n} \partial \theta^{2}}\left\{C \theta^{-\frac{n-5}{2}} Q^{-\left(\frac{(p-1)(n-5)}{8}-1\right)} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)+\kappa\right)\right\} \\
& +Q^{\frac{\tilde{\nu}_{2}}{\alpha}+1-n} O\left(\theta^{\tilde{\nu}_{2}-2}\right)+\mu \tilde{\nu}_{1}\left(\tilde{\nu}_{1}-1\right) B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}-2} \\
& +O\left(\mu^{2} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}+\nu_{1}-2}+\mu Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1-n} \theta^{\sigma+\tilde{\nu}_{1}-2}\right), \\
& \frac{\partial^{n}}{\partial Q^{n}}\left(\frac{\partial^{3}}{\partial \theta^{3}} v(Q, \mu, \theta)\right)=\frac{\partial^{n}}{\partial Q^{n}}\left(-\alpha(\alpha+1)(\alpha+2) \frac{A_{p}}{\theta^{\alpha+3}}\right) \\
& +\frac{\partial^{n+3}}{\partial Q^{n} \partial \theta^{3}}\left\{C \theta^{-\frac{n-5}{2}} Q^{-\left(\frac{(p-1)(n-5)}{8}-1\right)} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)+\kappa\right)\right\} \\
& +Q^{{\frac{\tilde{\nu}_{2}}{\alpha}}_{\alpha}+1-n} O\left(\theta^{\tilde{\tilde{2}}_{2}-3}\right)+\mu \tilde{\nu}_{1}\left(\tilde{\nu}_{1}-1\right)\left(\tilde{\nu}_{1}-2\right) B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1-n} \theta^{\tilde{\nu}_{1}-3} \\
& +O\left(\mu^{2} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1-n} \theta^{\tilde{\tilde{\nu}}_{1}+\nu_{1}-3}+\mu Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1-n} \theta^{\sigma+\tilde{\nu}_{1}-3}\right) .
\end{aligned}
$$

For $n=0,1$, we have

$$
\begin{aligned}
& \frac{\partial^{n}}{\partial \mu^{n}}(v(Q, \mu, \theta))= \mu^{1-n} B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}} \\
&+O\left(\mu^{2-n} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}}+1\right. \\
&\left.\theta^{\tilde{\nu}_{1}+\nu_{1}}+\mu^{1-n} Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1} \theta^{\sigma+\tilde{\nu}_{1}}\right), \\
& \frac{\partial^{n}}{\partial \mu^{n}}\left(\frac{\partial}{\partial \theta} v(Q, \mu, \theta)\right)= \mu^{1-n} \tilde{\nu}_{1} B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}-1} \\
&+O\left(\mu^{2-n} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}+\nu_{1}-1}+\mu^{1-n} Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1} \theta^{\sigma+\tilde{\nu}_{1}-1}\right), \\
& \frac{\partial^{n}}{\partial \mu^{n}}\left(\frac{\partial^{2}}{\partial \theta^{2}} v(Q, \mu, \theta)\right)= \mu^{1-n} \tilde{\nu}_{1}\left(\tilde{\nu}_{1}-1\right) B_{p} Q^{\frac{\tilde{\nu}_{1}}{\alpha}+1} \theta^{\tilde{1}_{1}-2} \\
&+O\left(\mu^{2-n} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}+\nu_{1}-2}+\mu^{1-n} Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1} \theta^{\sigma+\tilde{\nu}_{1}-2}\right), \\
& \frac{\partial^{n}}{\partial \mu^{n}}\left(\frac{\partial^{3}}{\partial \theta^{2}} v(Q, \mu, \theta)\right)= \mu^{1-n} \tilde{\nu}_{1}\left(\tilde{\nu}_{1}-1\right)\left(\tilde{\nu}_{1}-2\right) B_{p} Q^{\tilde{\nu}_{1}} \alpha 1 \\
& \theta^{\tilde{\nu}_{1}-3} \\
&+O\left(\mu^{2-n} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}+\nu_{1}-3}+\mu^{1-n} Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}+1} \theta^{\sigma+\tilde{\nu}_{1}-3}\right),
\end{aligned}
$$

while for $n=2$, we have

$$
\frac{\partial^{2}}{\partial \mu^{2}}\left(\frac{\partial^{m}}{\partial \theta^{m}} v(Q, \mu, \theta)\right)=O\left(Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}+\nu_{1}-m}\right), m=0,1,2,3,
$$

where $\kappa=\tan ^{-1} \frac{b_{0}}{a_{0}}, C=\sqrt{a_{0}^{2}+b_{0}^{2}}$.

Proof. These estimates are obtained by expansions of $v_{01}(r)$ and $v_{02}(r)$ given above and direct calculations.
Lemma 3.4. In the region $\theta=\left|O\left(Q^{\frac{\sigma}{(2-\sigma) \alpha}}\right)\right|, \mu=O\left(\theta^{2-\frac{2 \nu_{1}}{\sigma}}\right), \sigma=-\frac{1}{2}(n-5-2 \alpha)$, the solution $w(Q, \mu, \theta)$ of (1.7) with $w(Q, \mu, 0)=Q, w_{\theta}(Q, \mu, 0)=0, w_{\theta \theta}(Q, \mu, 0)=$ $(\tilde{b}+\mu) Q^{1+\frac{2}{\alpha}}, w_{\theta \theta \theta}(Q, \mu, 0)=0$ satisfies
(1) $\left|\frac{\partial^{m+n}}{\partial Q^{n} \partial \theta^{m}} w(Q, \mu, \theta)-\frac{\partial^{m+n}}{\partial Q^{n} \partial \theta^{m}} v(Q, \mu, \theta)\right|=Q^{-\frac{(n-5)(p-1)}{8}-(n-1)}\left|o\left(\theta^{-\frac{n-5}{2}-m}\right)\right|$,
(2) $\left|\frac{\partial^{m+n}}{\partial \mu^{n} \partial \theta^{m}} w(Q, \theta)-\frac{\partial^{n+m}}{\partial \mu^{n} \partial \theta^{m}} v(Q, \theta)\right|=\left|O\left(\mu^{2-n} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}+1} \theta^{\tilde{\nu}_{1}+\nu_{1}-m}\right)\right|$.

Proof. This lemma can be obtained from Lemma 3.3 and Theorem 3.1. Note that

$$
\epsilon=Q^{-\frac{1}{\alpha}},-\frac{\sigma}{\alpha}=\frac{(p-1)(n-5)}{8}-1
$$

Moreover

$$
Q^{\frac{p-1}{4}} \theta \in\left[e^{T}, e^{10 T}\right]
$$

provided that $Q$ is sufficiently large.
Now we write the inner solutions obtained in Theorem 3.1 in the form of parameter $Q$ and $\mu$ :
Theorem 3.2. Let $\frac{n+3}{n-5}<p<p_{c}(n-1)$ and $w_{Q, \mu}^{i n n}(\theta)$ be an inner solution of (1.7) with $w(0)=Q, w_{\theta}(0)=0, w_{\theta \theta}(0)=(\tilde{b}+\mu) Q^{1+\frac{2}{\alpha}}, w_{\theta \theta \theta}(0)=0$. Then for any sufficiently large $Q>0$ and $\theta=\left|O\left(Q^{\frac{\sigma}{(2-\sigma) \alpha}}\right)\right|=\left|O\left(\mu^{\frac{\sigma}{2 \sigma-2 \nu_{1}}}\right)\right|$,

$$
\begin{aligned}
w_{Q, \mu}^{i n n}(\theta)= & \frac{A_{p}}{\theta^{\alpha}}+\frac{C_{p}}{\theta^{\alpha-2}}+B_{p} \mu Q^{\frac{\nu_{1}}{\alpha}} \theta^{\tilde{\nu}_{1}}+\sum_{k=2}^{\infty} \sum_{j=1}^{k} d_{j}^{k} Q^{-\frac{(p-1)(k-j)}{2}} \theta^{2 j-\alpha} \\
& +Q^{\frac{\sigma}{\alpha}}\left[\frac{a_{0} \cos \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)\right)+b_{0} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)\right)}{\theta^{\frac{n-5}{2}}}\right. \\
& +\frac{a_{1} \cos \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)\right)+b_{1} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)\right)}{\theta^{\frac{n-5}{2}-2}}+O\left(\theta^{2-\frac{n-5}{2}}\right) \\
& \left.+\sum_{k=2}^{\infty}\left(\sum_{j=1}^{k} e_{j}^{k} Q^{-\frac{(p-1)(k-j)}{2}} \theta^{2 j-\frac{n-5}{2}} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)+E_{j}^{k}\right)+o\left(\theta^{2 k-\frac{n-5}{2}}\right)\right)\right] \\
& +Q \sum_{k=1}^{\infty}\left[\sum_{j=1}^{k}\left(\mu f_{j}^{k} Q^{-\frac{2 k-2 j-\tilde{\nu}_{1}}{\alpha}} \theta^{2 j+\tilde{\nu}_{1}}\right)+O\left(\mu^{2} Q^{\frac{\tilde{\nu}_{1}+\nu_{1}}{\alpha}} \theta^{\tilde{\nu}_{1}+\nu_{1}+2 k}\right.\right. \\
& \left.\left.+\mu Q^{\frac{\tilde{\nu}_{1}+\sigma}{\alpha}} \theta^{\tilde{\nu}_{1}+\sigma+2 k}\right)\right] .
\end{aligned}
$$

## 4. Outer solutions

In this section we construct outer solutions for (1.7). Let $w_{*}(\theta)$ be the singular solution given in (1.8), and we have the following lemma.
Lemma 4.1. Equation

$$
\begin{equation*}
T_{1} \phi(\theta)+k_{1} T_{2} \phi(\theta)+k_{0} \phi=p w_{*}^{p-1}(\theta) \phi(\theta), \quad 0<\theta<\frac{\pi}{2} \tag{4.1}
\end{equation*}
$$

admits a solution which can be written as

$$
\begin{equation*}
\phi(\theta)=\theta^{-\frac{n-5}{2}}\left[c_{1} \cos \left(\beta \ln \frac{\theta}{2}\right)+c_{2} \sin \left(\beta \ln \frac{\theta}{2}\right)\right]+O\left(\theta^{2-\frac{n-5}{2}}\right) \text { as } \theta \rightarrow 0 \tag{4.2}
\end{equation*}
$$

where $c_{1}$, $c_{2}$ are constants such that $c_{1}^{2}+c_{2}^{2} \neq 0$, and also admits another solution which can be written as

$$
\begin{equation*}
\psi(\theta)=c_{0} \theta^{\tilde{\nu}_{2}}+O\left(\theta^{\tilde{\nu}_{2}+2}\right) \text { as } \theta \rightarrow 0 \tag{4.3}
\end{equation*}
$$

where $c_{0}$ is a nonzero constant. Here $T_{1}$ and $T_{2}$ are differential operators defined in (1.7).

Proof. For the following equations,

$$
\left\{\begin{array}{l}
T_{1} \phi_{1}(\theta)+k_{1} T_{2} \phi_{1}(\theta)+k_{0} \phi_{1}(\theta)=p w_{*}^{p-1}(\theta) \phi_{1}(\theta), \quad 0<\theta<\frac{\pi}{2}  \tag{4.4}\\
\phi_{1}\left(\frac{\pi}{2}\right)=1, \quad \phi_{1}^{\prime}\left(\frac{\pi}{2}\right)=0, \quad \phi_{1}^{\prime \prime}\left(\frac{\pi}{2}\right)=0, \quad \phi_{1}^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

and

$$
\left\{\begin{array}{l}
T_{1} \phi_{2}(\theta)+k_{1} T_{2} \phi_{2}(\theta)+k_{0} \phi_{2}(\theta)=p w_{*}^{p-1}(\theta) \phi_{2}(\theta), \quad 0<\theta<\frac{\pi}{2},  \tag{4.5}\\
\phi_{2}\left(\frac{\pi}{2}\right)=0, \quad \phi_{2}^{\prime}\left(\frac{\pi}{2}\right)=0, \quad \phi_{2}^{\prime \prime}\left(\frac{\pi}{2}\right)=1, \quad \phi_{2}^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0,
\end{array}\right.
$$

we claim that both $\left.\phi_{1} \theta\right)$ and $\phi_{2}(\theta)$ are strictly decreasing for $\theta \in\left(0, \frac{\pi}{2}\right)$. We only show the case of $\phi_{2}(\theta)$, the case of $\phi_{1}(\theta)$ can be treated similarly.

Let us set $A(\theta)=\frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d \phi_{2}(\theta)}{d \theta}\right)$. Before proving that $\phi_{2}(\theta)$ is decreasing, we first present a useful fact that will be used in the following proof. If $A(\theta)>0$ for $\theta \in\left(\theta_{0}, \frac{\pi}{2}\right)$, where $\theta_{0} \in\left(0, \frac{\pi}{2}\right)$, then $\phi_{2}^{\prime}(\theta)<0$ for $\theta \in\left(\theta_{0}, \frac{\pi}{2}\right)$ and $\phi_{2}(\theta)>0$ for $\theta \in\left(\theta_{0}, \frac{\pi}{2}\right)$. The proof of this fact is simple, thus we omit it here. Next, we start to show that $\phi_{2}(\theta)$ is decreasing. By using the boundary condition of $\phi_{2}$ at $\theta=\frac{\pi}{2}$, we have $A\left(\frac{\pi}{2}\right)=1$ and find $\theta_{1} \in\left(0, \frac{\pi}{2}\right)$ such that $A(\theta)>0$ for $\theta \in\left(\theta_{1}, \frac{\pi}{2}\right)$, then $\phi_{2}(\theta)>0$ for $\theta \in\left(\theta_{1}, \frac{\pi}{2}\right)$. Using the fact $k_{1}(n)<0$ and the second conclusion in Lemma 6.1, we have

$$
T_{1} \phi_{2}(\theta)=\left[p w_{*}^{p-1}-k_{0}\right] \phi_{2}(\theta)-k_{1} \frac{A(\theta)}{\sin ^{n-2} \theta}>0, \quad \text { for } \theta \in\left(\theta_{1}, \frac{\pi}{2}\right)
$$

Now we are going to show that $\theta_{1}=0$. If not, $\theta_{1} \in\left(0, \frac{\pi}{2}\right)$ and $A\left(\theta_{1}\right)=0$. For $\theta \in\left(\theta_{1}, \frac{\pi}{2}\right)$, we have

$$
\frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d}{d \theta}\left(\frac{A(\theta)}{\sin ^{n-2} \theta}\right)\right)>0
$$

using this inequality and the following fact

$$
\left.\frac{d}{d \theta}\left(\frac{A(\theta)}{\sin ^{n-2} \theta}\right)\right|_{\theta=\frac{\pi}{2}}=0
$$

we have

$$
\begin{equation*}
\frac{d}{d \theta}\left(\frac{A(\theta)}{\sin ^{n-2} \theta}\right)<0 \quad \text { for } \theta \in\left(\theta_{1}, \frac{\pi}{2}\right) \tag{4.6}
\end{equation*}
$$

It follows from (4.6) that

$$
\begin{equation*}
\frac{A(\theta)}{\sin ^{n-2} \theta}>1 \quad \text { for } \theta \in\left(\theta_{1}, \frac{\pi}{2}\right) \tag{4.7}
\end{equation*}
$$

which contradicts the fact that $A\left(\theta_{1}\right)=0$. Thus, $A(\theta)>0$ and $\phi_{2}^{\prime}(\theta)<0$ for $\theta \in\left(0, \frac{\pi}{2}\right)$. Hence, we prove the claim.

We now prove that there are $D_{1} \neq 0$ and $D_{2} \neq 0$ such that for $\theta$ near 0 ,

$$
\begin{equation*}
\phi_{1}(\theta)=D_{1} \theta^{\tilde{\nu}_{2}}+O\left(\theta^{2+\tilde{\nu}_{2}}\right) \tag{4.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\phi_{2}(\theta)=D_{2} \theta^{\tilde{\nu}_{2}}+O\left(\theta^{2+\tilde{\nu}_{2}}\right) \tag{4.9}
\end{equation*}
$$

We also only show (4.9). The proof of (4.8) is similar. Using Emden-Fowler transformations:

$$
\tilde{\phi}(t)=(\sin \theta)^{\alpha} \phi_{2}(\theta), t=\ln \left(\tan \frac{\theta}{2}\right)
$$

we obtain that for $t \in(-\infty, 0), \tilde{\phi}(t)$ satisfies the following homogeneous equation

$$
\begin{equation*}
\tilde{\phi}^{(4)}(t)+a_{3}(t) \tilde{\phi}^{\prime \prime \prime}(t)+a_{2}(t) \tilde{\phi}^{\prime \prime}(t)+a_{1}(t) \tilde{\phi}^{\prime}(t)+a_{0}(t) \tilde{\phi}(t)=0 \tag{4.10}
\end{equation*}
$$

where
$a_{3}(t)=K_{3}+O\left(e^{2 t}\right), a_{2}(t)=K_{2}+O\left(e^{2 t}\right), a_{1}(t)=K_{1}+O\left(e^{2 t}\right), a_{0}(t)=(1-p) K_{0}$.
Therefore,

$$
\begin{array}{r}
\tilde{\phi}^{(4)}(t)+K_{3} \tilde{\phi}^{\prime \prime \prime}(t)+K_{2} \tilde{\phi}^{\prime \prime}(t)+K_{1} \tilde{\phi}^{\prime}(t)+(1-p) K_{0} \tilde{\phi}(t) \\
=O\left(e^{2 t}\left(\tilde{\phi}^{\prime \prime \prime}(t)+\tilde{\phi}^{\prime \prime}(t)+\tilde{\phi}^{\prime}(t)\right)\right) \tag{4.11}
\end{array}
$$

Following the arguments in the proof of Lemma 3.1, we can write the solutions of (4.11) as (for any $T \ll-1$ ):

$$
\begin{align*}
\tilde{\phi}(t)= & A_{5} e^{\sigma t} \cos \beta t+A_{6} e^{\sigma t} \sin \beta t+A_{7} e^{\nu_{2} t}+A_{8} e^{\nu_{1} t} \\
& +B_{5} \int_{-\infty}^{t} e^{\sigma(t-s)} \sin \beta(t-s) g(s, \tilde{\phi}(s)) \mathrm{d} s \\
& +B_{6} \int_{-\infty}^{t} e^{\sigma(t-s)} \cos \beta(t-s) g(s, \tilde{\phi}(s)) \mathrm{d} s \\
& +B_{7} \int_{-\infty}^{t} e^{\nu_{2}(t-s)} g(s, \tilde{\phi}(s)) \mathrm{d} s \\
& +B_{8} \int_{T}^{t} e^{\nu_{1}(t-s)} g(s, \tilde{\phi}(s)) \mathrm{d} s \tag{4.12}
\end{align*}
$$

where $g(t, \tilde{\phi}(t))$ is the right hand side of (4.11), $A_{8}$ depends on $T$ and each $B_{i+4}$ depends only on $\nu_{i}(i=1,2,3,4)$. It is known from (4.12) that if $A_{7}=0$, then for $|t|$ sufficiently large,

$$
\begin{equation*}
\tilde{\phi}(t)=A_{5} e^{\sigma t} \cos \beta t+A_{6} e^{\sigma t} \sin \beta t+O\left(e^{(2+\sigma) t}\right) \tag{4.13}
\end{equation*}
$$

with $A_{5}^{2}+A_{6}^{2} \neq 0$ or

$$
\begin{equation*}
\tilde{\phi}(t)=A_{8} e^{\nu_{1} t}+O\left(e^{\left(2+\nu_{1}\right) t}\right) \tag{4.14}
\end{equation*}
$$

with $A_{8} \neq 0$. Otherwise, if $A_{5}^{2}+A_{6}^{2}=0$ and $A_{8}=0$, we know that $\tilde{\phi}(t)=$ $O\left(e^{\left(2+\nu_{1}\right) t}\right)$. Substituting this into (4.12), we see that $\tilde{\phi}(t)=O\left(e^{\left(4+\nu_{1}\right) t}\right)$, repeating this procedure, we eventually obtain that $\tilde{\phi}(t) \equiv 0$. This is impossible. Therefore, for $\theta$ near 0 ,

$$
\phi_{2}(\theta)=A_{5} \theta^{-\frac{n-5}{2}} \cos \left(\beta \ln \frac{\theta}{2}\right)+A_{6} \theta^{-\frac{n-5}{2}} \sin \left(\beta \ln \frac{\theta}{2}\right)+O\left(\theta^{2-\frac{n-5}{2}}\right)
$$

or

$$
\phi_{2}(\theta)=A_{8} \theta^{\tilde{\nu}_{1}}+O\left(\theta^{2+\tilde{\nu}_{1}}\right)
$$

But these contradict the fact that $\phi_{2}(\theta)$ is strictly decreasing for $\theta \in\left(0, \frac{\pi}{2}\right)$. Thus, we prove the claim and get (4.9).

Let $\phi(\theta)=\phi_{1}(\theta)-\frac{D_{1}}{D_{2}} \phi_{2}(\theta)$. Then $\phi(\theta)$ satisfies the problem

$$
\left\{\begin{array}{l}
T_{1} \phi(\theta)+k_{1} T_{2} \phi(\theta)+k_{0} \phi(\theta)=p w_{*}^{p-1}(\theta) \phi(\theta), \quad 0<\theta<\frac{\pi}{2},  \tag{4.15}\\
\phi\left(\frac{\pi}{2}\right)=1, \quad \phi^{\prime}\left(\frac{\pi}{2}\right)=0, \quad \phi^{\prime \prime}\left(\frac{\pi}{2}\right)=-\frac{D_{1}}{D_{2}}, \quad \phi^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0 .
\end{array}\right.
$$

We claim that for $\theta$ near 0 ,

$$
\begin{equation*}
\phi(\theta)=\theta^{-\frac{n-5}{2}}\left[c_{1} \cos \left(\beta \ln \frac{\theta}{2}\right)+c_{2} \sin \left(\beta \ln \frac{\theta}{2}\right)\right]+O\left(\theta^{2-\frac{n-5}{2}}\right) \tag{4.16}
\end{equation*}
$$

with $c_{1}^{2}+c_{2}^{2} \neq 0$. Using Emden-Fowler transformations:

$$
\begin{equation*}
\hat{\phi}(t)=(\sin \theta)^{\alpha} \phi(\theta), t=\ln \left(\tan \frac{\theta}{2}\right) \tag{4.17}
\end{equation*}
$$

and (4.8), (4.9), we obtain that for $t$ near $-\infty$,

$$
\begin{equation*}
\hat{\phi}(t)=e^{\sigma t}\left[c_{1} \cos (\beta t)+c_{2} \sin (\beta t)\right]+c_{3} e^{\nu_{1} t}+O\left(e^{(2+\sigma) t}\right) \tag{4.18}
\end{equation*}
$$

provided $c_{1}^{2}+c_{2}^{2} \neq 0$ or

$$
\begin{equation*}
\hat{\phi}(t)=c_{3} e^{\nu_{1} t}+O\left(e^{\left(2+\nu_{1}\right) t}\right) \tag{4.19}
\end{equation*}
$$

provided $c_{1}^{2}+c_{2}^{2}=0$ and $c_{3} \neq 0$. (Note that if both $c_{1}^{2}+c_{2}^{2}=0$ and $c_{3}=0$, we can obtain $\hat{\phi}(t) \equiv 0$. This is impossible.) We now show that (4.19) can not occur. On the contrary, we see that for $\theta$ near 0 ,

$$
\phi(\theta)=c_{3} \theta^{\tilde{\nu}_{1}}+O\left(\theta^{2+\tilde{\nu}_{1}}\right)
$$

This implies that $\phi(\theta) \rightarrow 0$ as $\theta \rightarrow 0$. Since

$$
\hat{\phi}(t)=O\left(e^{\nu_{1} t}\right), \quad \hat{\phi}^{\prime}(t)=O\left(e^{\nu_{1} t}\right), \quad \hat{\phi}^{\prime \prime}(t)=O\left(e^{\nu_{1} t}\right), \quad \hat{\phi}^{\prime \prime \prime}(t)=O\left(e^{\nu_{1} t}\right)
$$

we obtain from (4.17) that
$\phi^{\prime}(\theta)=O\left(\theta^{\tilde{\nu}_{1}-1}\right), \sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}=O\left(\theta^{n-3+\tilde{\nu}_{1}}\right), \quad \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}\right)=O\left(\theta^{n-4+\tilde{\nu}_{1}}\right)$.
The similar arguments imply that

$$
\sin ^{n-2} \theta \frac{d}{d \theta}\left(\frac{1}{\sin ^{n-2} \theta} \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}\right)\right)=O\left(\theta^{n-5+\tilde{\nu}_{1}}\right) .
$$

If we define $e(\theta)=\sin ^{n-2} \theta \frac{d}{d \theta}\left(\frac{1}{\sin ^{n-2} \theta} \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}\right)\right)$, we see that $e(0)=0$. Then, we claim that $\phi$ changes sign in $\left(0, \frac{\pi}{2}\right)$. Suppose that this is not true, without loss generality, we assume $\phi>0$ in $\left(0, \frac{\pi}{2}\right)$. Then it follows from the equation of $\phi$ that for $\theta \in\left(0, \frac{\pi}{2}\right)$,

$$
\begin{equation*}
\frac{d}{d \theta}\left[e(\theta)-k_{1}\left(\sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}\right)\right]=\sin ^{n-2} \theta\left[p w_{*}^{p-1}-k_{0}\right] \phi(\theta)>0 \tag{4.20}
\end{equation*}
$$

But integrating both sides of (4.20) in ( $0, \frac{\pi}{2}$ ) and using the boundary conditions : $\phi^{\prime}\left(\frac{\pi}{2}\right)=\phi^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0$, we obtain that

$$
\int_{0}^{\frac{\pi}{2}} \sin ^{n-2} \theta\left[p w_{*}^{p-1}-k_{0}\right] \phi(\theta) d \theta=0
$$

This is clearly impossible. Noticing that $\phi \neq 0$ for $\theta$ near 0 , we see that there is a minimal zero point $\hat{\theta} \in\left(0, \frac{\pi}{2}\right)$ of $\phi$. Without loss generality, we assume that $\phi>0$ in $(0, \hat{\theta})$. It follows from (4.20) that $E(\theta):=e(\theta)+k_{1}\left(\sin ^{n-2} \theta \frac{\mathrm{~d} \phi(\theta)}{\mathrm{d} \theta}\right)$ is increasing
for $\theta \in(0, \hat{\theta})$. Noticing $E(0)=0$, we then obtain that $E(\theta)>0$ for $\theta \in(0, \hat{\theta})$. Therefore,

$$
\begin{equation*}
\frac{d}{d \theta}\left[\frac{1}{\sin ^{n-2} \theta} \frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}\right)+k_{1} \phi(\theta)\right]>0 \quad \text { for } \theta \in(0, \hat{\theta}) \tag{4.21}
\end{equation*}
$$

Moreover, by a similar argument, we have

$$
\begin{equation*}
\frac{d}{d \theta}\left(\sin ^{n-2} \theta \frac{d \phi(\theta)}{d \theta}\right)>0 \quad \text { for } \theta \in(0, \hat{\theta}) \tag{4.22}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{d \phi(\theta)}{d \theta}>0 \quad \text { for } \theta \in(0, \hat{\theta}) \tag{4.23}
\end{equation*}
$$

But (4.23) implies that $\phi(\hat{\theta})>0$, which contradicts the fact that $\phi(\hat{\theta})=0$. This contradiction implies that (4.19) can not occur and thus (4.18) holds. As a consequence, (4.16) holds and hence (4.2) holds.

Let $\psi(\theta)=\phi_{1}(\theta)$. We easily know that (4.3) can be obtained from (4.8). This completes the proof of this lemma.

For any sufficiently small $\delta>\eta>0$, we set $\psi_{1}(\theta)$ to be the solution of the problem

$$
\left\{\begin{array}{l}
T_{1} \psi_{1}(\theta)+k_{1} T_{2} \psi_{1}(\theta)+k_{0} \psi_{1}(\theta)=\eta^{-2}\left[\left(w_{*}+\Phi+\Psi\right)^{p}-w_{*}^{p}-p w_{*}^{p-1}\left(\Phi+\eta^{2} \psi\right)\right]  \tag{4.24}\\
\left(\psi_{1}+\psi\right)\left(\frac{\pi}{2}\right)=2, \quad\left(\psi_{1}+\psi\right)^{\prime}\left(\frac{\pi}{2}\right)=0, \quad\left(\psi_{1}+\psi\right)^{\prime \prime}\left(\frac{\pi}{2}\right)=\frac{D_{1} \delta^{2}}{D_{2} \eta^{2}}, \quad\left(\psi_{1}+\psi\right)^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

where $\psi(\theta)$ is given in Lemma 4.1, $\Phi=\delta^{2} \phi(\theta)$ and $\Psi=\eta^{2}\left(\psi_{1}(\theta)+\psi(\theta)\right)$. We can see that $\Psi$ satisfies the problem

$$
\left\{\begin{array}{l}
T_{1} \Psi(\theta)+k_{1} T_{2} \Psi(\theta)+k_{0} \Psi(\theta)=\left(w_{*}+\Phi+\Psi\right)^{p}-w_{*}^{p}-p w_{*}^{p-1} \Phi  \tag{4.25}\\
\Psi\left(\frac{\pi}{2}\right)=2 \eta^{2}, \quad \Psi^{\prime}\left(\frac{\pi}{2}\right)=0, \quad \Psi^{\prime \prime}\left(\frac{\pi}{2}\right)=\frac{D_{1} \delta^{2}}{D_{2}}, \quad \Psi^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

This implies that

$$
\left\{\begin{array}{l}
T_{1}(\Psi+\Phi)+k_{1} T_{2}(\Psi+\Phi)+k_{0}(\Psi+\Phi)=\left(w_{*}+\Phi+\Psi\right)^{p}-w_{*}^{p}  \tag{4.26}\\
(\Psi+\Phi)\left(\frac{\pi}{2}\right)=2 \eta^{2}+\delta^{2}, \quad(\Psi+\Phi)^{\prime}\left(\frac{\pi}{2}\right)=0, \quad(\Psi+\Phi)^{\prime \prime}\left(\frac{\pi}{2}\right)=0, \quad(\Psi+\Phi)^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

Arguments similar to those in the proof of Lemma 4.1 imply that $\Psi(\theta)+\Phi(\theta)$ is strictly decreasing, then

$$
\begin{equation*}
\Psi(\theta)+\Phi(\theta)>0 \quad \text { for } \theta \in\left(0, \frac{\pi}{2}\right) \tag{4.27}
\end{equation*}
$$

Setting $\psi_{2}(\theta)=\psi(\theta)+\psi_{1}(\theta)$, we easily see that $\psi_{2}$ satisfies the problem

$$
\left\{\begin{array}{c}
T_{1} \psi_{2}(\theta)+k_{1} T_{2} \psi_{2}(\theta)+k_{0} \psi_{2}(\theta)=p w_{*}^{p-1} \psi_{2}  \tag{4.28}\\
\quad+\eta^{-2}\left[\left(w_{*}+\Phi+\eta^{2} \psi_{2}\right)^{p}-w_{*}^{p}-p w_{*}^{p-1}\left(\Phi+\eta^{2} \psi_{2}\right)\right] \\
\psi_{2}\left(\frac{\pi}{2}\right)=2, \quad \psi_{2}^{\prime}\left(\frac{\pi}{2}\right)=0, \quad \psi_{2}^{\prime \prime}\left(\frac{\pi}{2}\right)=\frac{D_{1} \delta^{2}}{D_{2} \eta^{2}}, \quad \psi_{2}^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

By the Emden-Fowler transformation:

$$
\tilde{\psi}_{2}(t)=(\sin \theta)^{\alpha} \psi_{2}(\theta), \quad t=\ln \tan \frac{\theta}{2}
$$

we see that $\tilde{\psi}_{2}(t)$ satisfies the problem

$$
\left\{\begin{array}{l}
\tilde{\psi}_{2}^{(4)}(t)+a_{3}(t) \tilde{\psi}_{2}^{\prime \prime \prime}(t)+a_{2}(t) \tilde{\psi}_{2}^{\prime \prime}(t)+a_{1}(t) \tilde{\psi}_{2}^{\prime}(t)+a_{0}(t) \tilde{\psi}_{2}(t)=G\left(\tilde{\psi}_{2}(t)\right),-\infty<t<0,  \tag{4.29}\\
\tilde{\psi}_{2}^{\prime}(0)=0, \tilde{\psi}_{2}^{\prime \prime \prime}(0)=0,
\end{array}\right.
$$

where $a_{0}(t), a_{1}(t), a_{2}(t), a_{3}(t)$ are defined in (4.10), and
$G\left(\tilde{\psi}_{2}(t)\right)=(\sin \theta)^{4+\alpha} \eta^{-2}\left\{\left[w_{*}+\Phi+\eta^{2} \sin ^{-\alpha} \theta \tilde{\psi}_{2}\right]^{p}-w_{*}^{p}-p w_{*}^{p-1}\left(\Phi+\eta^{2} \sin ^{-\alpha} \theta \tilde{\psi}_{2}\right)\right\}$.
Moreover, we can rewrite (4.29) as the following form: (see the proof of Lemma 4.1)

$$
\begin{equation*}
\tilde{\psi}_{2}^{(4)}(t)+K_{3} \tilde{\psi}_{2}^{\prime \prime \prime}(t)+K_{2} \tilde{\psi}_{2}^{\prime \prime}(t)+K_{1} \tilde{\psi}_{2}^{\prime}(t)+(1-p) K_{0} \tilde{\psi}_{2}(t)=G\left(\tilde{\psi}_{2}(t)\right)+g\left(t, \tilde{\psi}_{2}(t)\right), \tag{4.30}
\end{equation*}
$$

where $g\left(t, \tilde{\psi}_{2}(t)\right)=O\left(e^{2 t}\left(\tilde{\psi}_{2}^{\prime \prime \prime}(t)+\tilde{\psi}_{2}^{\prime \prime}(t)+\tilde{\psi}_{2}^{\prime}(t)\right)\right)$ for $t \ll-1$. Therefore, for $t<T$ with any $T \ll-1$,

$$
\begin{align*}
\tilde{\psi}_{2}(t)= & D_{5} e^{\nu_{2} t}+D_{6} e^{\sigma t} \cos \beta t+D_{7} e^{\sigma t} \sin \beta t+D_{8} e^{\nu_{1} t} \\
& +B_{5} \int_{-\infty}^{t} e^{\sigma(t-s)} \sin \beta(t-s)\left(G\left(\tilde{\psi}_{2}(s)\right)+g\left(s, \tilde{\psi}_{2}(s)\right)\right) \mathrm{d} s \\
& +B_{6} \int_{-\infty}^{t} e^{\sigma(t-s)} \cos \beta(t-s)\left(G\left(\tilde{\psi}_{2}(s)\right)+g\left(s, \tilde{\psi}_{2}(s)\right)\right) \mathrm{d} s \\
& +B_{7} \int_{-\infty}^{t} e^{\nu_{2}(t-s)}\left(G\left(\tilde{\psi}_{2}(s)\right)+g\left(s, \tilde{\psi}_{2}(s)\right)\right) \mathrm{d} s \\
& +B_{8} \int_{T}^{t} e^{\nu_{1}(t-s)}\left(G\left(\tilde{\psi}_{2}(s)\right)+g\left(s, \tilde{\psi}_{2}(s)\right)\right) \mathrm{d} s, \tag{4.31}
\end{align*}
$$

where $B_{5}, B_{6}, B_{7}, B_{8}$ depend only on $\nu_{i}(i=1,2,3,4)$. Using the fact $\Psi(\theta)+\Phi(\theta)$ is strictly decreasing in ( $0, \frac{\pi}{2}$ ) and equation (4.2), we conclude that $D_{5} \neq 0$. Let $\phi(\theta)=$ $\sin ^{-\alpha} \theta \tilde{\phi}(t)$, we see that for $t \in[10 T, 2 T]$ and $\delta^{2}=O\left(e^{(2-\sigma) t}\right), \eta^{2}=O\left(e^{\left(2-\nu_{2}\right) t}\right)$,

$$
\begin{equation*}
G\left(\tilde{\psi}_{2}(t)\right)=\eta^{-2} O\left(\left(\delta^{2} \tilde{\phi}(t)+\eta^{2} \tilde{\psi}_{2}(t)\right)^{2}\right)=O\left(e^{\left(2+\nu_{2}\right) t}\right) . \tag{4.32}
\end{equation*}
$$

Note that $\tilde{\phi}(t)=e^{\sigma t}\left(c_{1} \cos (\beta t)+c_{2} \sin (\beta t)\right)+O\left(e^{(2+\sigma) t}\right)$ and $\tilde{\psi}_{2}(t)=D_{5} e^{\nu_{2} t}+$ $O\left(e^{\left(2+\nu_{2}\right) t}\right)$. Then

$$
\delta^{2} \tilde{\phi}(t)+\eta^{2} \tilde{\psi}_{2}(t)=O\left(e^{2 t}\right) .
$$

Therefore, it follows from (4.31) and (4.32) that

$$
\begin{equation*}
\tilde{\psi}_{2}(t)=D_{5} e^{\nu_{2} t}+D_{6} e^{\sigma t} \cos \beta t+D_{7} e^{\sigma t} \sin \beta t+O\left(e^{\left(2+\nu_{2}\right) t}\right) \tag{4.33}
\end{equation*}
$$

provided $\delta^{2}=O\left(e^{(2-\sigma) t}\right)$ and $\eta^{2}=O\left(e^{\left(2-\nu_{2}\right) t}\right)$. Hence, for $\theta$ near 0 ,

$$
\begin{equation*}
\Psi(\theta)=\eta^{2}\left[D_{5} \theta^{\tilde{\nu}_{2}}+\theta^{-\frac{n-5}{2}}\left(D_{6} \cos \left(\beta \ln \frac{\theta}{2}\right)+D_{7} \sin \left(\beta \ln \frac{\theta}{2}\right)\right)+O\left(\theta^{2+\tilde{\nu}_{2}}\right)\right] \tag{4.34}
\end{equation*}
$$

with $D_{5} \neq 0$ provided that $\theta=O\left(\delta^{\frac{2}{2-\sigma}}\right)=O\left(\eta^{\frac{2}{2-\nu_{2}}}\right)$. Since $\tilde{\nu}_{2}<3-n$, we easily know that $\tilde{\nu}_{2}+2<-(n-5)<-(n-5) / 2$. Thus, $\theta^{-\frac{n-5}{2}}=o\left(\theta^{2+\tilde{\nu}_{2}}\right)$.

Now we can obtain the following theorem.

Theorem 4.1. For any $\delta>\eta>0$ sufficiently small, problem (1.7) admits outer solutions $w_{\delta, \eta}^{\text {out }} \in C^{4}\left(0, \frac{\pi}{2}\right)$ satisfying

$$
\begin{equation*}
w_{\delta, \eta}^{\text {out }}(\theta)=w_{*}(\theta)+\Phi(\theta)+\Psi(\theta), \quad \theta \in\left(0, \frac{\pi}{2}\right) \tag{4.35}
\end{equation*}
$$

with $\left(w_{\delta, \eta}^{\text {out }}\right)^{\prime}\left(\frac{\pi}{2}\right)=\left(w_{\delta, \eta}^{\text {out }}\right)^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0$. Moreover,

$$
\begin{align*}
w_{\delta, \eta}^{\text {out }}(\theta)= & \frac{A_{p}}{\theta^{\alpha}}+\frac{2 A_{p}}{3(p-1)} \frac{1}{\theta^{\alpha-2}}+\delta^{2}\left[\frac{\vartheta_{1} \cos \left(\beta \ln \frac{\theta}{2}\right)+\vartheta_{2} \sin \left(\beta \ln \frac{\theta}{2}\right)}{\theta^{\frac{n-5}{2}}}+O\left(\frac{1}{\theta^{\frac{n-5}{2}-2}}\right)\right] \\
& +\eta^{2}\left[\vartheta_{3} \theta^{\tilde{\nu}_{2}}+O\left(\theta^{\tilde{\nu}_{2}+2}\right)\right] \tag{4.36}
\end{align*}
$$

provided that $\theta=O\left(\delta^{\frac{2}{2-\sigma}}\right)=O\left(\eta^{\frac{2}{2-\nu_{2}}}\right)$, where $\vartheta_{1}, \vartheta_{2}, \vartheta_{3}$ are constants independent of $\delta, \eta$ such that $\vartheta_{1}^{2}+\vartheta_{2}^{2} \neq 0, \vartheta_{3} \neq 0$.

Proof. The proof can be obtained from the expressions of $w_{*}(\theta), \Phi(\theta)$ and $\Psi(\theta)$ given in (1.8), (4.16) and (4.34).

## 5. Infinitely many solutions of (1.7) and Proof of Theorem 1.1

In this section, we construct infinitely many regular solutions for (1.7) by matching the inner and outer solutions.

We construct solutions of the problem

$$
\left\{\begin{array}{l}
T_{1} w+k_{1} T_{2} w+k_{0} w=w^{p}, \quad w(\theta)>0,0<\theta<\frac{\pi}{2},  \tag{5.1}\\
w(0)=Q\left(:=\epsilon^{-\alpha}\right), w^{\prime}\left(\frac{\pi}{2}\right)=0, w^{\prime \prime}(0)=(\tilde{b}+\mu) \epsilon^{-\alpha-2}, w^{\prime \prime \prime}\left(\frac{\pi}{2}\right)=0
\end{array}\right.
$$

by matching the inner and outer solutions given in Theorems 3.2 and 4.1. To do so, we will find $\Theta \in(0, \pi / 2)$ with

$$
\Theta=O\left(Q^{\frac{\sigma}{(2-\sigma) \alpha}}\right)(Q \gg 1)
$$

such that the following identities hold:

$$
\begin{gather*}
{\left.\left[w_{Q, \mu}^{\text {inn }}(\theta)-w_{\delta, \eta}^{\text {out }}(\theta)\right]\right|_{\theta=\Theta}=0}  \tag{5.2}\\
{\left.\left[w_{Q, \mu}^{i n n}(\theta)-w_{\delta, \eta}^{o u t}(\theta)\right]_{\theta}^{\prime}\right|_{\theta=\Theta}=0}  \tag{5.3}\\
{\left.\left[w_{Q, \mu}^{i n n}(\theta)-w_{\delta, \eta}^{\text {out }}(\theta)\right]_{\theta}^{\prime \prime}\right|_{\theta=\Theta}=0}  \tag{5.4}\\
{\left.\left[w_{Q, \mu}^{i n n}(\theta)-w_{\delta, \eta}^{o u t}(\theta)\right]_{\theta}^{\prime \prime \prime}\right|_{\theta=\Theta}=0} \tag{5.5}
\end{gather*}
$$

These will be done by arguments similar to those in the proof of Lemma 6.1 of [1] and Theorem 1.1 of [4]. Then, we obtain a $C^{4}$ function $w(\theta)$ defined by $w(\theta)=w_{Q, \mu}^{i n n}(\theta)$ for $\theta \leq \Theta$ and $w(\theta)=w_{\delta, \eta}^{o u t}(\theta)$ for $\theta \geq \Theta$ which is a solution to (5.1).

First, we observe that

$$
\begin{equation*}
\frac{2 A_{p}}{3(p-1)}=C_{p} \tag{5.6}
\end{equation*}
$$

by (3.35), where $A_{p}, C_{p}$ are given in Section 3.
Define $Q_{*}, \mu_{*}, \delta_{*}^{2}$ and $\eta_{*}^{2}$ by

$$
\begin{gather*}
\beta \ln Q_{*}^{\frac{p-1}{4}}+\kappa=\beta \ln 2^{-1}+\omega+2 m \pi  \tag{5.7}\\
\delta_{*}^{2}=\sqrt{\frac{a_{0}^{2}+b_{0}^{2}}{\vartheta_{1}^{2}+\vartheta_{2}^{2}}} Q_{*}^{\frac{\sigma}{\alpha}} \tag{5.8}
\end{gather*}
$$

$$
\begin{gather*}
\eta_{*}^{2}=O\left(Q_{*}^{\frac{\left(2-\nu_{2}\right) \sigma}{(2-\sigma) \alpha}}\right), \mu_{*}=O\left(Q_{*}^{\frac{2 \sigma-2 \nu_{1}}{(2-\sigma) \alpha}}\right),  \tag{5.9}\\
\mu_{*} B_{p} Q_{*}^{\frac{\nu_{1}}{\alpha}}=\vartheta_{3} \eta_{*}^{2} \Theta_{*}^{\tilde{\nu}_{2}-\tilde{\nu}_{1}} \tag{5.10}
\end{gather*}
$$

where $\kappa, \omega$ are given by

$$
\kappa=\tan ^{-1}\left(\frac{b_{0}}{a_{0}}\right), \omega=\tan ^{-1}\left(\frac{\vartheta_{2}}{\vartheta_{1}}\right)
$$

and $m \gg 1$ is an integer. The integer $m$ is chosen such that the results in Section 3 and Section 4 hold.

Note that

$$
\begin{gathered}
O\left(\delta_{*}^{\frac{2}{2-\sigma}}\right)=O\left(Q_{*}^{\frac{\sigma}{\alpha(2-\sigma)}}\right), \\
a_{0} \cos \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)\right)+b_{0} \sin \left(\beta \ln \left(Q^{\frac{p-1}{4}} \theta\right)\right)=\sqrt{a_{0}^{2}+b_{0}^{2}} \sin \left(\beta \ln \theta+\beta \ln Q^{\frac{p-1}{4}}+\kappa\right), \\
\vartheta_{1} \cos \left(\beta \ln \frac{\theta}{2}\right)+\vartheta_{2} \sin \left(\beta \ln \frac{\theta}{2}\right)=\sqrt{\vartheta_{1}^{2}+\vartheta_{2}^{2}} \sin \left(\beta \ln \theta+\beta \ln 2^{-1}+\omega\right) .
\end{gathered}
$$

We will see that $Q, \mu, \delta^{2}$ and $\eta^{2}$ required to satisfy the matching conditions (5.2)(5.5) can be obtained as small perturbations of $Q_{*}, \mu_{*}, \delta_{*}^{2}$ and $\eta_{*}^{2}$ given in (5.7)(5.10), i.e.,

$$
\begin{align*}
& Q=Q_{*}\left(1+O\left(Q_{*}^{\frac{2 \sigma}{(2-\sigma) \alpha}}\right)\right),  \tag{5.11}\\
& \mu=\mu_{*}\left(1+O\left(Q_{*}^{\frac{2 \sigma}{(2-\sigma) \alpha}}\right)\right),  \tag{5.12}\\
& \delta^{2}=\delta_{*}^{2}\left(1+O\left(Q_{*}^{\frac{2 \sigma}{(2-\sigma) \alpha}}\right)\right),  \tag{5.13}\\
& \eta^{2}=\eta_{*}^{2}\left(1+O\left(Q_{*}^{\frac{2 \sigma}{(2-\sigma) \alpha}}\right)\right) . \tag{5.14}
\end{align*}
$$

To show this we define the function $\mathbf{F}(Q, \mu, \delta, \eta)$ by

$$
\mathbf{F}\left(Q, \mu, \delta^{2}, \eta^{2}\right)=\left[\begin{array}{c}
\Theta^{\frac{n-5}{2}}\left(w_{Q, \mu}^{i n n}(\Theta)-w_{\delta, \eta}^{o u t}(\Theta)\right) \\
\left.\Theta\left[\theta^{\frac{n-5}{2}}\left(w_{Q, \mu}^{i n n}(\theta)-w_{\delta, \eta}^{o u t}(\theta)\right)\right]_{\theta}^{\prime}\right|_{\theta=\Theta} \\
\left.\Theta^{2}\left[\theta^{\frac{n-5}{2}}\left(w_{Q, \mu}^{i n n}(\theta)-w_{\delta, \eta}^{o u t}(\theta)\right)\right]_{\theta}^{\prime \prime}\right|_{\theta=\Theta} \\
\left.\Theta^{3}\left[\theta^{\frac{n-5}{2}}\left(w_{Q, \mu}^{i n n}(\theta)-w_{\delta, \eta}^{o u t}(\theta)\right)\right]_{\theta}^{\prime \prime \prime}\right|_{\theta=\Theta}
\end{array}\right]^{T} .
$$

Now, we regard $\delta^{2}, \eta^{2}$ as new variables. Taking $Q_{*}, \mu_{*}, \delta_{*}^{2}$ and $\eta_{*}^{2}$, we find a bound for $\mathbf{F}\left(Q_{*}, \mu_{*}, \delta_{*}^{2}, \eta_{*}^{2}\right)$ by making use of the behaviors of $w_{Q, \mu}^{i n n}(\theta)$ and $w_{\delta, \eta}^{o u t}(\theta)$ given in Theorems 3.2 and Theorem 4.1 respectively. Accordingly we find for some $M>1$ suitably large,

$$
\begin{equation*}
\left|\Theta^{-\frac{n-5}{2}} \mathbf{F}\left(Q_{*}, \mu_{*}, \delta_{*}^{2}, \eta_{*}^{2}\right)\right| \leq M \Theta^{4-\sigma-\frac{n-5}{2}}+\text { small terms } \tag{5.15}
\end{equation*}
$$

We now seek values of $Q, \mu, \delta^{2}, \eta^{2}$ which are small perturbations of $Q_{*}, \mu_{*}, \delta_{*}^{2}, \eta_{*}^{2}$ and such that $\mathbf{F}\left(Q, \mu, \delta^{2}, \eta^{2}\right)=0$. As in [4], we need to evaluate the Jacobian of $\mathbf{F}$ at $\left(Q_{*}, \mu_{*}, \delta_{*}^{2}, \eta_{*}^{2}\right)$ :

$$
\frac{\partial \mathbf{F}\left(Q, \mu, \delta^{2}, \eta^{2}\right)}{\partial\left(Q, \mu, \delta^{2}, \eta^{2}\right)}=\left[\begin{array}{cccc}
I_{1}+I_{3} & I_{4} & -D \sin \tau & I_{5} \\
\beta I_{2}+q_{1} I_{3} & q_{1} I_{4} & -\beta D \cos \tau & q_{4} I_{5} \\
I_{6} & q_{2} I_{4} & I_{8} & q_{5} I_{5} \\
I_{7} & q_{3} I_{4} & I_{9} & q_{6} I_{5}
\end{array}\right]+\text { h.o.t., }
$$

where

$$
I_{1}=C\left(\frac{\sigma}{\alpha} \sin \tau+\frac{\beta(p-1)}{4} \cos \tau\right) Q_{*}^{\frac{\sigma}{\alpha}-1}, I_{2}=C\left(\frac{\sigma}{\alpha} \cos \tau-\frac{\beta(p-1)}{4} \sin \tau\right) Q_{*}^{\frac{\sigma}{\alpha}-1}
$$

$$
\begin{gathered}
I_{3}=\frac{\nu_{1}}{\alpha} B_{p} \mu_{*} \Theta^{\tilde{\nu}_{1}+\frac{n-5}{2}} Q_{*}^{\frac{\nu_{1}}{\alpha}-1}, I_{4}=B_{p} Q_{*}^{\frac{\nu_{1}}{\alpha}} \Theta^{\tilde{\nu}_{1}+\frac{n-5}{2}}, I_{5}=-\vartheta_{3} \Theta^{\tilde{\nu}_{2}+\frac{n-5}{2}}, \\
I_{6}=-\beta^{2} I_{1}-\beta I_{2}+q_{2} I_{3}, I_{7}=-\beta^{3} I_{2}+3 \beta^{2} I_{1}+2 \beta I_{2}+q_{3} I_{3}, \\
I_{8}=\beta^{2} D \sin \tau+\beta D \cos \tau, I_{9}=\beta^{3} D \cos \tau-3 \beta^{2} D \sin \tau-2 \beta D \cos \tau \\
q_{1}=\tilde{\nu}_{1}+\frac{n-5}{2}, q_{2}=\left(\tilde{\nu}_{1}+\frac{n-7}{2}\right) q_{1}, q_{3}=\left(\tilde{\nu}_{1}+\frac{n-9}{2}\right) q_{2}, \\
q_{4}=\tilde{\nu}_{2}+\frac{n-5}{2}, q_{5}=\left(\tilde{\nu}_{2}+\frac{n-7}{2}\right) q_{4}, q_{6}=\left(\tilde{\nu}_{1}+\frac{n-9}{2}\right) q_{5}, \\
C=\sqrt{a_{0}^{2}+b_{0}^{2}}, \quad D=\sqrt{\vartheta_{1}^{2}+\vartheta_{2}^{2}}
\end{gathered}
$$

and

$$
\tau=\beta \ln \Theta+\beta \ln Q_{*}^{\frac{p-1}{4}}+\kappa=\beta \ln \Theta+\beta \ln 2^{-1}+\omega+2 m \pi
$$

We define the function $\mathbf{G}(x, y)$ by

$$
\mathbf{G}(x, y, z, w)=\mathbf{F}\left(Q_{*}+x Q_{*}^{1-\frac{\sigma}{\alpha}}, \mu_{*}+\Theta^{-\tilde{\nu}_{1}-\frac{n-5}{2}} Q_{*}^{-\frac{\nu_{1}}{\alpha}} y, \delta_{*}^{2}+z, \eta_{*}^{2}+\Theta^{-\tilde{\nu}_{2}-\frac{n-5}{2}} w\right)
$$

Using (5.15), (4.36) and the results in Lemmas 3.3, 3.4, we express $\mathbf{G}(x, y, z, w)$ in the form

$$
\begin{aligned}
\mathbf{G}(x, y, z, w)= & \mathbf{C}+\left[\begin{array}{cccc}
I_{1}^{\prime}+I_{3}^{\prime} & I_{4}^{\prime} & -D \sin \tau & I_{5}^{\prime} \\
\beta I_{2}^{\prime}+q_{1} I_{3}^{\prime} & q_{1} I_{4}^{\prime} & -\beta D \cos \tau & q_{4} I_{5}^{\prime} \\
I_{6}^{\prime} & q_{2} I_{4}^{\prime} & I_{8}^{\prime} & q_{5} I_{5}^{\prime} \\
I_{7}^{\prime} & q_{3} I_{4}^{\prime} & I_{9}^{\prime} & q_{6} I_{5}^{\prime}
\end{array}\right]\left(\begin{array}{c}
x \\
y \\
y \\
z
\end{array}\right) \\
& +\mathbf{E}\left(x, y, z, w, Q_{*}, \mu_{*}, \delta_{*}^{2}, \eta_{*}^{2}\right),
\end{aligned}
$$

where

$$
\begin{gathered}
I_{1}^{\prime}=C\left(\frac{\sigma}{\alpha} \sin \tau+\frac{\beta(p-1)}{4} \cos \tau\right), I_{2}^{\prime}=C\left(\frac{\sigma}{\alpha} \cos \tau-\frac{\beta(p-1)}{4} \sin \tau\right) \\
I_{3}^{\prime}=\frac{\nu_{1}}{\alpha} B_{p} \mu_{*} \Theta^{\tilde{\nu}_{1}+\frac{n-5}{2}} Q_{*}^{\frac{\nu_{1}-\sigma}{\alpha}}, I_{4}^{\prime}=B_{p}, I_{5}^{\prime}=-\vartheta_{3} \\
I_{6}^{\prime}=-\beta^{2} I_{1}^{\prime}-\beta I_{2}^{\prime}+q_{2} I_{3}^{\prime}, I_{7}^{\prime}=-\beta^{3} I_{2}^{\prime}+3 \beta^{2} I_{1}^{\prime}+2 \beta I_{2}^{\prime}+q_{3} I_{3}^{\prime} \\
I_{8}^{\prime}=\beta^{2} D \sin \tau+\beta D \cos \tau, I_{9}^{\prime}=\beta^{3} D \cos \tau-3 \beta^{2} D \sin \tau-2 \beta D \cos \tau
\end{gathered}
$$

$\mathbf{C}$ is a constant vector independent of $(x, y, z, w)$ which is bounded above by $M \delta_{*}^{4} \Theta^{\sigma}$, and $|\mathbf{E}|$ is bounded independently of $x, y,, z, w, Q, \mu, \delta$ and $\eta$. Thus,

$$
\mathbf{G}(x, y, z, w)=\mathbf{C}+L\left(\begin{array}{l}
x \\
y \\
z \\
w
\end{array}\right)+\mathbf{T}(x, y, z, w)
$$

where $L$ is a linear operator which is invertible, we shall prove this fact in Lemma 6.1). If we define the operator $\mathbf{J}$ mapping $\mathbb{R}^{4}$ into itself by

$$
\mathbf{J}(x, y, z, w)=-\left(L^{-1} \mathbf{C}+L^{-1} \mathbf{T}(x, y, z, w)\right)
$$

then, provided that $Q_{*}$ is sufficiently large, a direct calculation shows that $\mathbf{J}$ maps the set $I$ into itself, where $I$ is the ball

$$
\begin{equation*}
I=\left\{(x, y, z, w):\left(x^{2}+y^{2}+z^{2}+w^{2}\right)^{\frac{1}{2}} \leq 4 M(\operatorname{det} L)^{-1} \Theta^{4-\sigma}\right\} \tag{5.16}
\end{equation*}
$$

and $\operatorname{det} L$ is the determindent of $L$ which depends on $\sqrt{a_{0}^{2}+b_{0}^{2}}, \beta, D, \alpha, B_{p}, \vartheta_{3}$ and $\nu_{i}, i=1,2,3,4$. An application of the Brouwer Fixed Point Theorem to conclude that $\mathbf{J}$ has a fixed point in $I$. This point $(x, y, z, w)$ satisfies $\mathbf{G}(x, y, z, w)=0$ and

$$
\left(x^{2}+y^{2}+z^{2}+w^{2}\right)^{\frac{1}{2}} \leq M^{\prime} \Theta^{4-\sigma}
$$

where $M^{\prime}$ is a constant defined in (5.16) and is independent of $Q_{*}, \mu_{*}, \delta_{*}, \eta_{*}$ and $\Theta$. By substituting for $Q, \mu \delta$ and $\eta$, then taking $\Theta$ to have the upper limiting value of $Q_{*}^{\frac{\sigma}{(2-\sigma) \alpha}}$, we obtain (5.11)- (5.14). Therefore, we can find a solution to (5.1) such that (5.2)-(5.5) holds.

We have shown that (5.2)-(5.5) has a solution for each fixed $m$ large. This yields a solution of (5.1) and also gives the proof of Theorem 1.1. Hence we have
Theorem 5.1. For $m \gg 1$ large and $Q, \mu, \delta$ and $\eta$ being given in (5.11)-(5.14), problem ${ }_{\sigma}$ (5.1) admits a classical solution $w_{Q, \mu, \delta, \eta}(\theta)$. Moreover, there is $\Theta=$ $\left|O\left(Q^{\frac{\sigma}{(2-\sigma) \alpha}}\right)\right|$ such that (5.2)-(5.5) hold.

As a consequence, problem (1.7) admits infinitely many nonconstant positive solutions. Hence, we prove Theorem 1.1.

## 6. Appendix

In the following lemma, we will show the results which were used in the previous sections.

Lemma 6.1. For the terms $E_{2}, k_{0}(n)$, and matrixes $K$ and $L$, which were defined in previous sections, we have
(1) $E_{2} \neq 0$,
(2) If $p \in\left(\frac{n+3}{n-5}, p_{c}(n-1)\right)$, we have $p k_{0}(n-1) \geq k_{0}(n)$,
(3) $\operatorname{det} K \neq 0$,
(4) $\operatorname{det} L \neq 0$.

Proof. First, we show that $E_{2} \neq 0$, it is known that

$$
\begin{equation*}
E_{2}=\left(\tilde{\nu}_{1}+2\right) \tilde{\nu}_{1}\left(\tilde{\nu}_{1}+n-3\right)\left(\tilde{\nu}_{1}+n-1\right)-p(n-5-\alpha)(n-3-\alpha)(2+\alpha) \alpha \tag{6.1}
\end{equation*}
$$

For convenience, we use $n$ instead of $n-1$ and $\tilde{\nu}_{1}(n)$ instead of $\tilde{\nu}_{1}(n-1)$, i.e. we study the following term

$$
\begin{equation*}
E_{2}=\left(\tilde{\nu}_{1}+2\right) \tilde{\nu}_{1}\left(\tilde{\nu}_{1}+n-2\right)\left(\tilde{\nu}_{1}+n\right)-p(n-4-\alpha)(n-2-\alpha)(2+\alpha) \alpha \tag{6.2}
\end{equation*}
$$

Let $f(\alpha)=p(n-4-\alpha)(n-2-\alpha)(2+\alpha) \alpha$, through a simple computation, we get $f(\alpha)$ and its derivative $f^{\prime}(\alpha)$
$f(\alpha)=\alpha^{4}+(12-2 n) \alpha^{3}+\left(n^{2}-18 n+52\right) \alpha^{2}+\left(6 n^{2}-52 n+96\right) \alpha+8(n-2)(n-4)$,
and

$$
f^{\prime}(\alpha)=4 \alpha^{3}+(36-6 n) \alpha^{2}+\left(2 n^{2}-36 n+104\right) \alpha+\left(6 n^{2}-52 n+96\right)
$$

We compute the roots of $f^{\prime}(\alpha)$ and get all its zero points $\frac{1}{2}\left(n-6 \pm \sqrt{n^{2}+4}\right), \frac{1}{2}(n-6)$. It is easy to see that $f(\alpha)$ is strictly increasing in $\alpha \in\left(0, \frac{1}{2}(n-6)\right)$ and decreasing in $\alpha \in\left(\frac{1}{2}(n-6), \frac{1}{2}\left(n-6+\sqrt{n^{2}+4}\right)\right)$. We know that $\alpha=\frac{4}{p-1}<\frac{n-4}{2}$ and $\frac{n-4}{2} \in$ $\left(\frac{1}{2}(n-6), \frac{1}{2}\left(n-6+\sqrt{n^{2}+4}\right)\right)$. As a consequence, we can conclude

$$
f(\alpha) \leq f\left(\frac{n-6}{2}\right)=\frac{n^{4}}{16}-\frac{n^{2}}{2}+1 \text { for all } p \in\left(\frac{n+4}{n-4}, p_{c}(n)\right)
$$

Let $g(x)=x(x+2)(x+n)(x+n-2)=x^{4}+2 n x^{3}+\left(n^{2}+2 n-4\right) x^{2}+\left(2 n^{2}-4 n\right) x$, we compute its derivative $g^{\prime}(x)=4 x^{3}+6 n x^{2}+\left(2 n^{2}+4 n-8\right) x+\left(2 n^{2}-4 n\right)$ and find $g^{\prime}(x)>0$ for $x>0$ when $n \geq 5$. On the other hand, using $4 \sqrt{N_{3}}>N_{2}$ for $p \in\left(\frac{n+4}{n-4}, p_{c}(n)\right)$, we find

$$
\tilde{\nu}_{1}>\frac{\sqrt{2\left(n^{2}-4 n+8\right)}-(n-4)}{2}
$$

therefore,

$$
\begin{align*}
g\left(\tilde{\nu}_{1}\right) \geq g\left(\frac{\sqrt{2\left(n^{2}-4 n+8\right)}-(n-4)}{2}\right)= & 96-40 n+11 n^{2}-\frac{n^{3}}{2}+\frac{n^{4}}{16} \\
& +\sqrt{2}\left(24-4 n+n^{2}\right) \sqrt{8-4 n+n^{2}} . \tag{6.3}
\end{align*}
$$

Comparing $\frac{n^{4}}{16}-\frac{n^{2}}{2}+1$ and the right hand side of (6.3), by direct computation, we can get

$$
g\left(\frac{\sqrt{2\left(n^{2}-4 n+8\right)}-(n-4)}{2}\right)>\frac{n^{4}}{16}-\frac{n^{2}}{2}+1 \text { for } n \in(0, \infty)
$$

As a result, $g\left(\tilde{\nu}_{1}\right)>f(\alpha)$. Hence, $E_{2}$ is nonzero.
Next, we prove the inequality $p k_{0}(n-1) \geq k_{0}(n)$ for $p \in\left(\frac{n+3}{n-5}, p_{c}(n-1)\right)$. According to the definition of $k_{0}(n)$, it is enough for us to show the following

$$
\begin{equation*}
p(n-5-\alpha)(n-3-\alpha) \geq(n-4-\alpha)(n-2-\alpha) \tag{6.4}
\end{equation*}
$$

Using the relation $p=\frac{4}{\alpha}+1$, it is equivalent to show the following (after computation)

$$
\begin{equation*}
6 \alpha^{2}+(39-10 n) \alpha+4 n^{2}-32 n+60 \geq 0 \tag{6.5}
\end{equation*}
$$

It is known that (6.5) holds provided

$$
\alpha \geq \frac{10 n-39+\sqrt{4 n^{2}-12 n+81}}{12} \text { or } \alpha \leq \frac{10 n-39-\sqrt{4 n^{2}-12 n+81}}{12} .
$$

On the other hand, since $p \in\left(\frac{n+3}{n-5}, p_{c}(n-1)\right)$, we have $\alpha<\frac{n-5}{2}$. It is easy to show $\frac{n-5}{2} \leq \frac{10 n-39-\sqrt{4 n^{2}-12 n+81}}{12}$ when $n \geq 5$. Hence, (6.5) holds. Therefore (6.4) holds.

Then, we show $K$ is invertible, it is enough for us to show $B \neq 0$ or $A \neq 0$. we recall

$$
B=\left(2 n^{2}-12 n-6\right) \beta+8 \beta^{3}=\left(2(n-3)^{2}-24\right) \beta+8 \beta^{3}
$$

It is known that $2(n-3)^{2}-24<0$ only when $n=6$. Since $\beta>0$, therefore $B \neq 0$ when $n \geq 7$. When $n=6$, we find

$$
A=\beta^{4}-\frac{35}{2} \beta^{2}-\frac{135}{2}-(1-\alpha)(3-\alpha)(2+\alpha) \alpha, B=-6 \beta+8 \beta^{3}
$$

If $B \neq 0$ for $n=6$, we have $K$ is invertible, while if $B=0$ for $n=6$, then $A=-21-(1-\alpha)(3-\alpha)(2+\alpha) \alpha<0$ for $\alpha \in(0,1 / 2)$ and $K$ is also invertible. Therefore, we proved the third conclusion.

Finally, we show the matrix $L$ is invertible. Recall that the matrix $L$ is given by

$$
L:=\left[\begin{array}{cccc}
I_{1}^{\prime}+I_{3}^{\prime} & I_{4}^{\prime} & -D \sin \tau & I_{5}^{\prime}  \tag{6.6}\\
\beta I_{2}^{\prime}+q_{1} I_{3}^{\prime} & q_{1} I_{4}^{\prime} & -\beta D \cos \tau & q_{4} I_{5}^{\prime} \\
I_{6}^{\prime} & q_{2} I_{4}^{\prime} & I_{8}^{\prime} & q_{5} I_{5}^{\prime} \\
I_{7}^{\prime} & q_{3} I_{4}^{\prime} & I_{9}^{\prime} & q_{6} I_{5}^{\prime}
\end{array}\right]
$$

where

$$
\begin{gathered}
I_{1}^{\prime}=C\left(\frac{\sigma}{\alpha} \sin \tau+\frac{\beta(p-1)}{4} \cos \tau\right), I_{2}^{\prime}=C\left(\frac{\sigma}{\alpha} \cos \tau-\frac{\beta(p-1)}{4} \sin \tau\right), \\
I_{3}^{\prime}=\frac{\nu_{1}}{\alpha} B_{p} \mu_{*} \Theta^{\tilde{\nu}_{1}+\frac{n-5}{2}} Q_{*}^{\frac{\nu_{1}-\sigma}{\alpha}}, I_{4}^{\prime}=B_{p}, I_{5}^{\prime}=\vartheta_{3} \\
I_{6}^{\prime}=-\beta^{2} I_{1}^{\prime}-\beta I_{2}^{\prime}+q_{2} I_{3}^{\prime}, I_{7}^{\prime}=-\beta^{3} I_{2}^{\prime}+3 \beta^{2} I_{1}^{\prime}+2 \beta I_{2}^{\prime}+q_{3} I_{3}^{\prime} \\
I_{8}^{\prime}=\beta^{2} D \sin \tau+\beta D \cos \tau, I_{9}^{\prime}=\beta^{3} D \cos \tau-3 \beta^{2} D \sin \tau-2 \beta D \cos \tau
\end{gathered}
$$

Using simple linear transformations, we see that

$$
\begin{aligned}
& {\left[\begin{array}{cccc}
I_{1}^{\prime}+I_{3}^{\prime} & I_{4}^{\prime} & -D \sin \tau & I_{5}^{\prime} \\
\beta I_{2}^{\prime}+q_{1} I_{3}^{\prime} & q_{1} I_{4}^{\prime} & -\beta \cos \tau & q_{4} I_{5}^{\prime} \\
I_{6}^{\prime} & q_{2} I_{4}^{\prime} & I_{8}^{\prime} & q_{5} I_{5}^{\prime} \\
I_{7}^{\prime} & q_{3} I_{4}^{\prime} & I_{9}^{\prime} & q_{6} I_{5}^{\prime}
\end{array}\right] \sim\left[\begin{array}{cccc}
I_{1}^{\prime} & I_{4}^{\prime} & -D \sin \tau & I_{5}^{\prime} \\
\beta I_{2}^{\prime} & q_{1} I_{4}^{\prime} & -\beta D \cos \tau & q_{4} I_{5}^{\prime} \\
I_{6}^{\prime}-q_{2} I_{3}^{\prime} & q_{2} I_{4}^{\prime} & I_{8}^{\prime} & q_{5} I_{5}^{\prime} \\
I_{7}^{\prime}-q_{3} I_{3}^{\prime} & q_{3} I_{4}^{\prime} & I_{9}^{\prime} & q_{6} I_{5}^{\prime}
\end{array}\right]} \\
& \sim\left[\begin{array}{cccc}
I_{1}^{\prime} & -D \sin \tau & I_{4}^{\prime} & I_{5}^{\prime} \\
\beta I_{2}^{\prime} & -\beta D \cos \tau & q_{1} I_{4}^{\prime} & q_{4} I_{5}^{\prime} \\
I_{6}^{\prime}-q_{2} I_{3}^{\prime} & I_{8}^{\prime} & q_{2} I_{4}^{\prime} & q_{5} I_{5}^{\prime} \\
I_{7}^{\prime}-q_{3} I_{3}^{\prime} & I_{9}^{\prime} & q_{3} I_{4}^{\prime} & q_{6} I_{5}^{\prime}
\end{array}\right] \sim\left[\begin{array}{cccc}
I_{1}^{\prime} & -D \sin \tau & I_{4}^{\prime} & -I_{5}^{\prime} \\
\beta I_{2}^{\prime} & -\beta D \cos \tau & q_{1} I_{4}^{\prime} & -q_{4} I_{5}^{\prime} \\
0 & 0 & I_{10} & I_{11}^{\prime} \\
0 & 0 & I_{12}^{\prime} & I_{13}^{\prime},
\end{array}\right],
\end{aligned}
$$

where

$$
\begin{gathered}
I_{10}^{\prime}=q_{2} B_{p}+q_{1} B_{p}+\beta^{2} B_{p}, I_{12}^{\prime}=q_{3} B_{p}+\beta^{2} q_{1} B_{p}-3 \beta^{2} B_{p}-2 q_{1} B_{p}, \\
I_{11}^{\prime}=q_{5} \vartheta_{3}+q_{4} \vartheta_{3}+\beta^{2} \vartheta_{3}, I_{13}^{\prime}=q_{6} \vartheta_{3}+\beta^{2} q_{4} \vartheta_{3}-3 \beta^{2} \vartheta_{3}-2 q_{4} \vartheta_{3} .
\end{gathered}
$$

Since

$$
\operatorname{det}\left[\begin{array}{cc}
I_{1}^{\prime} & -D \sin \tau \\
\beta I_{2}^{\prime} & -\beta D \cos \tau
\end{array}\right] \neq 0
$$

showing that $L$ is invertible is equivalent to proving that the following two by two matrix is invertible:

$$
\left[\begin{array}{cc}
q_{2}+q_{1}+\beta^{2} & q_{5}+q_{4}+\beta^{2}  \tag{6.7}\\
q_{3}+\beta^{2} q_{1}-3 \beta^{2}-2 q_{1} & q_{6}+\beta^{2} q_{4}-3 \beta^{2}-2 q_{4}
\end{array}\right] .
$$

It follows from the definitions of $q_{i},(i=1,2,3,4,5,6)$ and $\beta$ that $q_{2}+q_{1}+\beta^{2}=$ $q_{5}+q_{4}+\beta^{2} \neq 0$. Let

$$
\chi_{1}=q_{3}+\beta^{2} q_{1}-3 \beta^{2}-2 q_{1}, \quad \chi_{2}=q_{6}+\beta^{2} q_{4}-3 \beta^{2}-2 q_{4} .
$$

Then

$$
\begin{aligned}
\chi_{1}-\chi_{2} & =q_{3}-q_{6}-\left(q_{1}-q_{4}\right)\left(2-\beta^{2}\right) \\
& =\left(\tilde{\nu}_{1}-\tilde{\nu}_{2}\right)\left[\left(\tilde{\nu}_{1}+\tilde{\nu}_{2}\right)^{2}-\tilde{\nu}_{1} \tilde{\nu}_{2}+\frac{3 n-21}{2}\left(\tilde{\nu}_{1}+\tilde{\nu}_{2}\right)+\frac{3 n^{2}-42 n+135}{4}+\beta^{2}\right] \\
& =\left(\tilde{\nu}_{1}-\tilde{\nu}_{2}\right)\left[\frac{n^{2}-10 n+25}{4}-\tilde{\nu}_{1} \tilde{\nu}_{2}+\beta^{2}\right]
\end{aligned}
$$

where we are using the fact $\tilde{\nu}_{1}+\tilde{\nu}_{2}=-(n-5)$. It is known that (from Section.2)

$$
\tilde{\nu}_{1} \tilde{\nu}_{2}=\frac{n^{2}-10 n+25}{4}-\frac{N_{2}+4 \sqrt{N_{3}}}{4(p-1)^{2}}
$$

and

$$
\beta^{2}=\frac{4 \sqrt{N_{3}}-N_{2}}{4(p-1)^{2}}
$$

where $N_{2}$ and $N_{3}$ (with the dimension $n$ being replaced by $n-1$ ) are defined in Section 2. Therefore,

$$
\chi_{1}-\chi_{2}=\left(\tilde{\nu}_{1}-\tilde{\nu}_{2}\right) \frac{2 \sqrt{N_{3}}}{(p-1)^{2}} \neq 0
$$

Hence, the above two by two matrix (6.7) is invertible.

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## References

[1] C. Budd and J. Norbury, Semilinear elliptic equations and supercritical growth, J. Differential Equations 68 (1987), no.2, 169-197.
[2] M.F. Bidaut-Veron and L. Veron, Nonlinear elliptic equations on compact Riemannian manifolds and asymptotics of Emden equations, Invent. Math. 106 (1991), 489-539.
[3] E.N. Dancer, Y.H. Du and Z.M. Guo, Finite Morse index solutions of an elliptic equation with supercritical exponent, J. Differential Equations 250(2011), 3281-3310.
[4] E.N. Dancer, Z.M. Guo and J.C. Wei, Non-radial singular solutions of Lane-Emden equation in $\mathbb{R}^{n}$, Indiana Univ. Math. J. 61 (2012), no. 5, 19711996.
[5] J. Dàvila, L. Dupaigne, K.L. Wang and J.C. Wei, A monotonicity formula and a Liouvilletype theorem for a fourth order supercritical problem, Adv. Math. 258 (2014), 240285.
[6] A. Farina, On the classification of solutions of Lane-Emden equation on unbounded domains in $\mathbb{R}^{N}$, J. Math. Pures Appl. 87 (2007), 537-561.
[7] F. Gazzola and H.Ch. Grunau, Radial entire solutions for supercritical biharmonic equations, Math. Ann. 334 (2006), 905-936.
[8] Z.M. Guo, Further study of entire solutions of a biharmonic equation with exponential nonlinearity, Ann. Mat. Pura Appl. (in press).
[9] Z.M. Guo, On the symmetry of positive solutions of the Lane-Emden equation with supercritical exponent, Adv. Differential Equations 7(2002), 641-666.
[10] Z. M. Guo and J. Wei, Qualitative properties of entire radial solutions for biharmonic equations with supercritical nonlinearity, Proc. Amer. Math. Soc. 138 (2010), 3957-3964.
[11] B. Gidas and J. Spruck, Global and local behavior of positive solutions of nonlinear elliptic equations, Comm. Pure Appl. Math. 24(1981), 525-598.
[12] C.F. Gui, W.M. Ni and X.F. Wang, On the stability and instability of positive steady states of a semilinear heat equation in $\mathbb{R}^{n}$, Comm. Pure Appl. Math. 45 (1992), 1153-1181.
[13] P. Hartman, Ordinary Differential Equations. 2nd edn. Wiley, New York (1982).
[14] R.A. Johnson, X.B. Pan and Y.F. Yi, Positive solutions of super-critical elliptic equations and asymptotics, Comm. PDEs 18 (1993), 977-1019.
[15] D.D. Joseph and T.S. Lundgren, Quasilinear Dirichlet problems driven by positive sources, Arch. Rational Mech. Anal. 49 (1973), 241-269.
[16] P. Karageorgis, Stability and intersection properties of solutions to the nonlinear biharmonic equation, Nonlinearity 22 (2009) 1653-1661.
[17] N. Korevaar, R. Mazzeo, F. Pacard and R. Schoen, Refined asmptotics for constant scalar curvature metrics with isolated singularities, Invent. Math. 135(1999), no.2, 233-272.
[18] C.S. Lin, A classification of solutions of a conformally invariant fourth order equation in $\mathbb{R}^{n}$, Comment. Math. Helv. 73(1998), 206-231.
[19] J.C. Wei and X.W. Xu, Classification of solutions of higher order conformally invariant equations, Math. Ann. 313(1999), no.2, 207-228.
[20] J.C. Wei, X.W. Xu and W. Yang, Classification of the stable solution to biharmonic problems in large dimensions, Pacific Journal of Math. 263 (2013), no.2, 495-512.
[21] J.C. Wei and D. Ye, Liouville theorems for finite morse index solutions of biharmonic problem, Math. Annalen 356 (2013), no. 4, 1599-1612.
[22] H.H. Zou, Symmetry of positive solutions of $\Delta u+u^{p}=0$ in $\mathbb{R}^{n}$. J. Differential Equations, 120(1995), 46-88.

Department of Mathematics, Henan Normal University, Xinxiang, 453007, P.R. China
E-mail address: gzm@htu.cn
Department of Mathematics, University of British Columbia, Vancouver, BC V6T 1Z2, CANADA

E-mail address: jcwei@math.ubc.ca
Department of Mathematics, University of British Columbia, Vancouver, BC V6T 1Z2, Canada

E-mail address: wyang@math.ubc.ca

