

Chapter 3

Basic Floquet Theory

3.1 General Results

If we have a problem of the form

$$\mathbf{x}' = \mathbf{A}(t)\mathbf{x} \quad (3.1)$$

where $\mathbf{A}(t)$ is periodic with period T , then \mathbf{x} need not be periodic, however it must be of the form

$$e^{\mu t} \mathbf{p}(t) \quad (3.2)$$

where $\mathbf{p}(t)$ has period T . Additionally, it has n such μ_j and together they satisfy

$$e^{\mu_1 T} e^{\mu_2 T} \dots e^{\mu_n T} = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) ds \right). \quad (3.3)$$

The following theorems prove those results. We follow Ward [28].

Definition (Fundamental Matrix). Let $\mathbf{x}^1(t), \dots, \mathbf{x}^n(t)$ be n solutions of $\mathbf{x}' = \mathbf{A}(t)\mathbf{x}$. Let

$$\mathbf{X}(t) = \left[\begin{bmatrix} \mathbf{x}^1 \end{bmatrix} \dots \begin{bmatrix} \mathbf{x}^n \end{bmatrix} \right] \quad (3.4)$$

so that $\mathbf{X}(t)$ is an $n \times n$ matrix solution of $\mathbf{X}' = \mathbf{A}\mathbf{X}$.

If $\mathbf{x}^1(t), \dots, \mathbf{x}^n(t)$ are linearly independent, then $\mathbf{X}(t)$ is non-singular and is called a *fundamental matrix*. If $\mathbf{X}(t_0) = \mathbf{I}$, then $\mathbf{X}(t)$ is the *principal fundamental matrix*.

Lemma 3.1. *If $\mathbf{X}(t)$ is a fundamental matrix then so is $\mathbf{Y}(t) = \mathbf{X}(t)\mathbf{B}$ for any non-singular constant matrix \mathbf{B} .*

Proof. Since $\mathbf{X}(t)$ and \mathbf{B} are non-singular then the inverse of $\mathbf{Y}(t)$ is $\mathbf{B}^{-1}\mathbf{X}^{-1}(t)$ and so $\mathbf{Y}(t)$ is non-singular. Also,

$$\mathbf{Y}' = \mathbf{X}'\mathbf{B} = \mathbf{A}\mathbf{X}\mathbf{B} = \mathbf{A}\mathbf{Y} \quad (3.5)$$

so that $\mathbf{Y}'(t) = \mathbf{A}\mathbf{Y}(t)$. \square

Lemma 3.2. *Let the Wronskian $W(t)$ of $\mathbf{X}(t)$ be the determinant of $\mathbf{X}(t)$. Then*

$$W(t) = W(t_0) \exp \left(\int_{t_0}^t \text{tr}(\mathbf{A}(s)) ds \right). \quad (3.6)$$

Proof. Let t_0 be some time. Expanding in a Taylor series,

$$\mathbf{X}(t) = \mathbf{X}(t_0) + (t - t_0) \mathbf{X}'(t_0) + O\left((t - t_0)^2\right) \quad (3.7)$$

$$= \mathbf{X}(t_0) + (t - t_0) \mathbf{A}(t_0) \mathbf{X}(t_0) + O\left((t - t_0)^2\right) \quad (3.8)$$

$$= [\mathbf{I} + (t - t_0) \mathbf{A}(t_0)] \mathbf{X}(t_0) + O\left((t - t_0)^2\right) \quad (3.9)$$

so that

$$\det(\mathbf{X}(t)) = \det[\mathbf{I} + (t - t_0) \mathbf{A}(t_0)] \det(\mathbf{X}(t_0)) \quad (3.10)$$

$$W(t) = \det[\mathbf{I} + (t - t_0) \mathbf{A}(t_0)] W(t_0). \quad (3.11)$$

Now since

$$\det(\mathbf{I} + \epsilon \mathbf{C}) = 1 + \epsilon \text{tr}(\mathbf{C}) + O(\epsilon^2), \quad (3.12)$$

we have that

$$W(t) = W(t_0) (1 + (t - t_0) \text{tr}(\mathbf{A}(t_0))). \quad (3.13)$$

Now by expanding $W(t)$ in a Taylor series, we obtain that

$$W(t) = W(t_0) + (t - t_0) W'(t_0) + O\left((t - t_0)^2\right) \quad (3.14)$$

so that

$$W'(t_0) = W(t_0) \text{tr}(\mathbf{A}(t_0)). \quad (3.15)$$

Since we have not made any assumptions about t_0 , we can write

$$W'(t) = W(t) \text{tr}(\mathbf{A}(t)). \quad (3.16)$$

We know that the solution to this equation is

$$W(t) = W(t_0) \exp \left(\int_{t_0}^t \text{tr}(\mathbf{A}(s)) ds \right) \quad (3.17) \quad \square$$

Theorem 3.3. *Let $\mathbf{A}(t)$ be a T -periodic matrix. If $\mathbf{X}(t)$ is a fundamental matrix then so is $\mathbf{X}(t + T)$ and there exists a non-singular constant matrix \mathbf{B} such that*

$$i. \quad \mathbf{X}(t + T) = \mathbf{X}(t) \mathbf{B} \text{ for all } t$$

$$ii. \quad \det(\mathbf{B}) = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) ds \right)$$

Proof. Begin by showing that $\mathbf{X}(t+T)$ is also a fundamental matrix. Let $\mathbf{Y}(t) = \mathbf{X}(t+T)$. Then

$$\mathbf{Y}'(t) = \mathbf{X}'(t+T) = \mathbf{A}(t+T)\mathbf{X}(t+T) = \mathbf{A}(t)\mathbf{X}(t+T) = \mathbf{A}(t)\mathbf{Y}(t) \quad (3.18)$$

and so $\mathbf{X}(t+T)$ is a fundamental matrix.

i. Let $\mathbf{B}(t) = \mathbf{X}^{-1}(t)\mathbf{Y}(t)$. Then

$$\mathbf{Y}(t) = \mathbf{X}(t)\mathbf{X}^{-1}(t)\mathbf{Y}(t) \quad (3.19)$$

$$= \mathbf{X}(t)\mathbf{B}(t) \quad (3.20)$$

Let $\mathbf{B}_0 = \mathbf{B}(t_0)$. We know by lemma 3.1 that $\mathbf{Y}_0(t) = \mathbf{X}(t)\mathbf{B}_0$ is a fundamental matrix, where, by definition, $\mathbf{Y}_0(t_0) = \mathbf{Y}(t_0)$. Since these are both solutions to $\mathbf{X}' = \mathbf{A}\mathbf{X}$, by the uniqueness of the solution, we must then have $\mathbf{Y}_0(t) = \mathbf{Y}(t)$ for all time. As a result, $\mathbf{B}_0 = \mathbf{B}(t)$ and so \mathbf{B} is time-independent.

ii. From lemma 3.2, we have that

$$W(t) = W(t_0) \exp \left(\int_{t_0}^t \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.21)$$

$$W(t+T) = W(t_0) \exp \left(\int_{t_0}^t \text{tr}(\mathbf{A}(s)) \, ds + \int_t^{t+T} \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.22)$$

$$W(t+T) = W(t) \exp \left(\int_t^{t+T} \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.23)$$

$$W(t+T) = W(t) \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) \, ds \right). \quad (3.24)$$

We also know that

$$\mathbf{X}(t+T) = \mathbf{X}(t)\mathbf{B} \quad (3.25)$$

$$\det(\mathbf{X}(t+T)) = \det(\mathbf{X}(t)) \det(\mathbf{B}) \quad (3.26)$$

$$W(t+T) = W(t) \det(\mathbf{B}) \quad (3.27)$$

and so

$$\det(\mathbf{B}) = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.28) \quad \square$$

Remark. Since \mathbf{B} is time-independent, it can be computed by setting $t = 0$, so that $\mathbf{B} = \mathbf{X}^{-1}(0)\mathbf{X}(T)$. If we took the initial conditions $\mathbf{X}(0) = \mathbf{I}$, then $\mathbf{B} = \mathbf{X}(T)$.

Definition (Characteristic Multipliers and Exponents). The eigenvalues ρ_1, \dots, ρ_n of \mathbf{B} are called the *characteristic multipliers* for $\mathbf{X}'(t) = \mathbf{A}(t)\mathbf{X}(t)$. The *characteristic exponents* or *Floquet exponents* are μ_1, \dots, μ_n satisfying

$$\rho_1 = e^{\mu_1 T}, \quad \rho_2 = e^{\mu_2 T}, \quad \dots \quad \rho_n = e^{\mu_n T}. \quad (3.29)$$

Note that μ_j for $j \in \mathbb{N}$ may be complex.

Properties.

- i. The characteristic multipliers (eigenvalues) ρ_1, \dots, ρ_n of $\mathbf{B} = \mathbf{X}(T)$ with $\mathbf{X}(0) = \mathbf{I}$ satisfy

$$\det(\mathbf{B}) = \rho_1 \rho_2 \cdots \rho_n = \exp \left(\int_0^T \operatorname{tr}(\mathbf{A}(s)) \, ds \right). \quad (3.30)$$

This follows from theorem 3.3ii.

- ii. Since the trace is the sum of the eigenvalues, we also have

$$\operatorname{tr}(\mathbf{B}) = \rho_1 + \rho_2 + \cdots + \rho_n. \quad (3.31)$$

- iii. The characteristic exponents are not unique since if $\rho_j = e^{\mu_j T}$, then $\rho_j = e^{(\mu_j + 2\pi i/T)T}$.

- iv. The characteristic multipliers ρ_j are an intrinsic property of the equation $\mathbf{X}'(t) = \mathbf{A}\mathbf{X}$ and do not depend on the choice of the fundamental matrix.

Proof. Suppose $\hat{\mathbf{X}}(t)$ is another fundamental matrix. Then

$$\hat{\mathbf{X}}(t+T) = \hat{\mathbf{X}}(t)\hat{\mathbf{B}}. \quad (3.32)$$

We have showed in the proof of theorem 3.3 that since $\mathbf{X}(t)$ and $\hat{\mathbf{X}}(t)$ are fundamental matrices then there is a constant non-singular matrix \mathbf{C} such that

$$\hat{\mathbf{X}}(t) = \mathbf{X}(t)\mathbf{C} \quad (3.33)$$

so that

$$\hat{\mathbf{X}}(t+T) = \mathbf{X}(t+T)\mathbf{C} \quad (3.34)$$

$$(\hat{\mathbf{X}}(t)\hat{\mathbf{B}}) = (\mathbf{X}(t)\mathbf{B})\mathbf{C} \quad (3.35)$$

$$\mathbf{X}(t)\mathbf{C}\hat{\mathbf{B}} = \mathbf{X}(t)\mathbf{B}\mathbf{C} \quad (3.36)$$

$$\mathbf{C}\hat{\mathbf{B}} = \mathbf{B}\mathbf{C} \quad (3.37)$$

$$\mathbf{C}\hat{\mathbf{B}}\mathbf{C}^{-1} = \mathbf{B} \quad (3.38)$$

so the eigenvalues of \mathbf{B} and $\hat{\mathbf{B}}$ are the same. \square

Theorem 3.4. *Let ρ be a characteristic multiplier and let μ be the corresponding characteristic exponent so that $\rho = e^{\mu T}$. Then there exists a solution $\mathbf{x}(t)$ of $\mathbf{x}' = \mathbf{A}(t)\mathbf{x}$ such that*

$$i. \mathbf{x}(t+T) = \rho\mathbf{x}(t)$$

- ii. *There exists a periodic solution $\mathbf{p}(t)$ with period T such that $\mathbf{x}(t) = e^{\mu t}\mathbf{p}(t)$.*

Proof.

- i. Let \mathbf{b} be an eigenvector of \mathbf{B} corresponding to eigenvalue ρ . Let $\mathbf{x}(t) = \mathbf{X}(t)\mathbf{b}$. Then $\mathbf{x}' = \mathbf{A}\mathbf{x}$ and

$$\mathbf{x}(t+T) = \mathbf{X}(t+T)\mathbf{b} \quad (3.39)$$

$$= \mathbf{X}(t)\mathbf{B}\mathbf{b} \quad (3.40)$$

$$= \rho\mathbf{X}(t)\mathbf{b} \quad (3.41)$$

$$= \rho\mathbf{x}(t) \quad (3.42)$$

so that $\mathbf{x}(t+T) = \rho\mathbf{x}(t)$.

- ii. Let $\mathbf{p}(t) = \mathbf{x}(t)e^{-\mu t}$. We now need to show that $\mathbf{p}(t)$ is T -periodic.

$$\mathbf{p}(t+T) = \mathbf{x}(t+T)e^{-\mu(t+T)} \quad (3.43)$$

$$= \rho\mathbf{x}(t)e^{-\mu(t+T)} \quad (3.44)$$

$$= \frac{\rho}{e^{\mu T}}\mathbf{x}(t)e^{-\mu t} \quad (3.45)$$

$$= \mathbf{x}(t)e^{-\mu t} \quad (3.46)$$

$$= \mathbf{p}(t) \quad (3.47)$$

As a result, we have a solution of the form $\mathbf{x}(t) = e^{\mu t}\mathbf{p}(t)$ where $\mathbf{p}(t)$ is periodic with period T . \square

Remarks.

- i. If μ is replaced by $\mu + 2\pi i/T$, then we get

$$\mathbf{x}(t) = e^{\mu t}\mathbf{p}(t)e^{2\pi i t/T} \quad (3.48)$$

where $\mathbf{p}(t)e^{2\pi i t/T}$ is still periodic with period T . As a result, the fact that μ is not unique does not alter our results.

- ii. We have that

$$\mathbf{x}_j(t+T) = \rho_j\mathbf{x}_j(t) \quad (3.49)$$

$$\mathbf{x}_j(t+NT) = \rho_j^N\mathbf{x}_j(t). \quad (3.50)$$

Each characteristic multipliers falls into one of the following categories:

- (a) If $|\rho| < 1$, then $\text{Re}(\mu) < 0$ and so $\mathbf{x}(t) \xrightarrow{t \rightarrow \infty} 0$.
- (b) If $|\rho| = 1$, then $\text{Re}(\mu) = 0$ and so we have a pseudo-periodic solution.
If $\rho = \pm 1$, then the solution is periodic with period T .
- (c) If $|\rho| > 1$, then $\text{Re}(\mu) > 0$ and so $\mathbf{x}(t) \rightsquigarrow \infty$ as $t \rightarrow \infty$.

The entire solution is stable if all the characteristic multipliers satisfy $|\rho_j| \leq 1$.

- iii. As for the general solution, suppose that $\mathbf{b}_1, \dots, \mathbf{b}_n$ are n linearly independent eigenvectors of \mathbf{B} corresponding to distinct eigenvalues ρ_1, \dots, ρ_n . Then there are n linearly independent solutions to $\mathbf{x}' = \mathbf{A}\mathbf{x}$, which by the above theorem are given by

$$\mathbf{x}_j(t) = e^{\mu_j t} \mathbf{p}_j(t) \quad (3.51)$$

where $\mathbf{p}_j(t)$ is T -periodic. As a result, we can define

$$\mathbf{X}_0(t) = \begin{bmatrix} \begin{bmatrix} \mathbf{x}_1 \end{bmatrix} & \cdots & \begin{bmatrix} \mathbf{x}_n \end{bmatrix} \end{bmatrix}, \quad \mathbf{P}_0(t) = \begin{bmatrix} \begin{bmatrix} \mathbf{p}_1 \end{bmatrix} & \cdots & \begin{bmatrix} \mathbf{p}_n \end{bmatrix} \end{bmatrix}, \quad (3.52)$$

$$\mathbf{D}_0(t) = \begin{bmatrix} \mu_1 & & 0 \\ & \ddots & \\ 0 & & \mu_n \end{bmatrix}, \quad \mathbf{Y}_0(t) = \begin{bmatrix} e^{\mu_1 t} & & 0 \\ & \ddots & \\ 0 & & e^{\mu_n t} \end{bmatrix}, \quad (3.53)$$

such that

$$\mathbf{X}_0 = \mathbf{P}_0 \mathbf{Y}_0, \quad \mathbf{Y}'_0 = \mathbf{D}_0 \mathbf{Y}_0. \quad (3.54)$$

- iv. Now consider what happens if $\rho < 0$. Suppose $\rho < 0$ real, so that we can write

$$\rho = e^{(\nu + i\pi/T)T} \quad (3.55)$$

where

$$\rho = -e^{\nu T}. \quad (3.56)$$

Then we obtain

$$\mathbf{x}(t) = e^{\mu t} \mathbf{p}(t) \quad (3.57)$$

$$= e^{\nu t} e^{i\pi t/T} \mathbf{p}(t) \quad (3.58)$$

$$= e^{\nu t} \mathbf{q}(t), \quad (3.59)$$

where $\mathbf{q}(t)$ has period T since $\mathbf{p}(t)$ has period T . Since we can choose \mathbf{x} to be real, without loss of generality, we can also choose \mathbf{q} to be real. For the general solution, if $\rho_j < 0$, we can replace \mathbf{p}_j with \mathbf{q}_j and μ_j with ν_j so that

$$\mathbf{P}_0 = \begin{bmatrix} \begin{bmatrix} \mathbf{p}_1 \end{bmatrix} & \cdots & \begin{bmatrix} \mathbf{q}_j \end{bmatrix} & \cdots & \begin{bmatrix} \mathbf{p}_n \end{bmatrix} \end{bmatrix}, \quad \mathbf{Y}_0 = \begin{bmatrix} e^{\mu_1 T} & & & 0 \\ & \ddots & & \\ & & e^{\nu_j T} & \\ & & & \ddots \\ 0 & & & & e^{\mu_n T} \end{bmatrix} \quad (3.60)$$

and

$$\mathbf{X}_0(t) = \mathbf{P}_0(t) \mathbf{Y}_0(t). \quad (3.61)$$

- v. Suppose now that ρ is complex. Then since ρ is an eigenvalue of the real matrix \mathbf{B} , $\bar{\rho}$ is as well. The characteristic exponents are μ and $\bar{\mu}$. Let

$$\mu = \nu + i\sigma, \quad \mathbf{p}(t) = \mathbf{q}(t) + i\mathbf{r}(t) \quad (3.62)$$

where $\mathbf{q}(t)$ and $\mathbf{r}(t)$ must both have period T since $\mathbf{p}(t)$ does. Since $\mathbf{x}(t) = e^{\mu t}\mathbf{p}(t)$ is a solution to $\mathbf{x}' = \mathbf{A}(t)\mathbf{x}$, then by taking the complex conjugate, so is $\bar{\mathbf{x}}(t) = e^{\bar{\mu}t}\bar{\mathbf{p}}(t)$. We can write these as

$$\mathbf{x}(t) = e^{(\nu+i\sigma)t} (\mathbf{q}(t) + i\mathbf{r}(t)) \quad (3.63)$$

$$= e^{\nu t} [(\mathbf{q} \cos(\sigma t) - \mathbf{r} \sin(\sigma t)) + i(\mathbf{r} \cos(\sigma t) + \mathbf{q} \sin(\sigma t))] \quad (3.64)$$

and

$$\bar{\mathbf{x}}(t) = e^{(\nu-i\sigma)t} (\mathbf{q}(t) - i\mathbf{r}(t)) \quad (3.65)$$

$$= e^{\nu t} [(\mathbf{q} \cos(\sigma t) + \mathbf{r} \sin(\sigma t)) - i(\mathbf{r} \cos(\sigma t) + \mathbf{q} \sin(\sigma t))]. \quad (3.66)$$

We can alternately write the linearly independent real solutions

$$\mathbf{x}_R = \text{Re} [e^{\mu t}\mathbf{p}(t)] = e^{\nu t} [\cos(\sigma t) \mathbf{q}(t) - \sin(\sigma t) \mathbf{r}(t)], \quad (3.67)$$

$$\mathbf{x}_I = \text{Im} [e^{\mu t}\mathbf{p}(t)] = e^{\nu t} [\sin(\sigma t) \mathbf{q}(t) + \cos(\sigma t) \mathbf{r}(t)], \quad (3.68)$$

so that

$$\mathbf{X}_0 = \begin{bmatrix} \begin{bmatrix} \mathbf{x}_1 \\ \vdots \\ \mathbf{x}_R \end{bmatrix} & \begin{bmatrix} \mathbf{x}_I \\ \vdots \\ \mathbf{x}_n \end{bmatrix} \end{bmatrix}, \quad (3.69)$$

$$\mathbf{P}_0 = \begin{bmatrix} \begin{bmatrix} \mathbf{p}_1 \\ \vdots \\ \mathbf{q} \end{bmatrix} & \begin{bmatrix} \mathbf{r} \\ \vdots \\ \mathbf{p}_n \end{bmatrix} \end{bmatrix}, \quad (3.70)$$

$$\mathbf{Y}_0 = \begin{bmatrix} e^{\mu_1 T} & & & & 0 \\ & \ddots & & & \\ & & e^{\nu t} \cos(\sigma t) & e^{\nu t} \sin(\sigma t) & \\ & & -e^{\nu t} \sin(\sigma t) & e^{\nu t} \cos(\sigma t) & \\ & & & & \ddots \\ 0 & & & & & e^{\mu_n T} \end{bmatrix} \quad (3.71)$$

and

$$\mathbf{X}_0(t) = \mathbf{P}_0(t)\mathbf{Y}_0(t). \quad (3.72)$$

3.1.1 Example

For example, consider

$$x_1' = \left(1 + \frac{\cos(t)}{2 + \sin(t)}\right) x_1 \quad (3.73)$$

$$x_2' = x_1 - x_2. \quad (3.74)$$

Here, we know that the solution is in general

$$x_1 = c_1 e^t (2 + \sin(t)) \quad (3.75)$$

$$x_2 = c_1 e^t \left(2 + \frac{1}{2} \sin(t) - \frac{1}{2} \cos(t) \right) + c_2 e^{-t} \quad (3.76)$$

which we can write as

$$\mathbf{x} = c_1 e^t \begin{bmatrix} 2 + \sin(t) \\ 2 + \frac{1}{2} \sin(t) - \frac{1}{2} \cos(t) \end{bmatrix} + c_2 e^{-t} \begin{bmatrix} 0 \\ 1 \end{bmatrix}. \quad (3.77)$$

Using all the above definitions, the fundamental matrix is

$$\mathbf{X}(t) = \begin{bmatrix} e^t (2 + \sin(t)) & 0 \\ e^t (2 + \frac{1}{2} \sin(t) - \frac{1}{2} \cos(t)) & e^{-t} \end{bmatrix} \quad (3.78)$$

so that

$$\mathbf{B} = \mathbf{X}^{-1}(0)\mathbf{X}(2\pi) \quad (3.79)$$

$$= \begin{bmatrix} 2 & 0 \\ \frac{3}{2} & 1 \end{bmatrix}^{-1} \begin{bmatrix} 2e^{2\pi} & 0 \\ \frac{3}{2}e^{2\pi} & e^{-2\pi} \end{bmatrix} \quad (3.80)$$

$$= \frac{1}{2} \begin{bmatrix} 1 & 0 \\ -\frac{3}{2} & 2 \end{bmatrix} \begin{bmatrix} 2e^{2\pi} & 0 \\ \frac{3}{2}e^{2\pi} & e^{-2\pi} \end{bmatrix} \quad (3.81)$$

$$= \begin{bmatrix} e^{2\pi} & 0 \\ 0 & e^{-2\pi} \end{bmatrix} \quad (3.82)$$

As a result $\rho_1 = e^{2\pi}$, $\rho_2 = e^{-2\pi}$ and so $\mu_1 = 1$ and $\mu_2 = -1$. Theorem 3.4 then tells us that there is a solution of the form

$$\mathbf{x}_1(t) = e^t \mathbf{p}_1(t), \quad \mathbf{x}_2(t) = e^{-t} \mathbf{p}_2(t) \quad (3.83)$$

where $\mathbf{p}_1(t)$ and $\mathbf{p}_2(t)$ are periodic with period 2π . We know that in fact

$$\mathbf{p}_1(t) = \begin{bmatrix} 2 + \sin(t) \\ 2 + \frac{1}{2} \sin(t) - \frac{1}{2} \cos(t) \end{bmatrix}, \quad \mathbf{p}_2(t) = \begin{bmatrix} 0 \\ 1 \end{bmatrix}. \quad (3.84)$$

3.1.2 Periodic Solution

Consider a problem of the form $\mathbf{x}' = \mathbf{f}(\mathbf{x})$ with $\mathbf{x} \in \mathbb{R}^n$ where there is a periodic solution $\mathbf{x}(t) = \boldsymbol{\phi}(t)$ with period T . Linearise the solution about $\boldsymbol{\phi}$ by writing $\mathbf{x} = \boldsymbol{\phi} + \mathbf{v}$. We then obtain

$$\mathbf{v}' = \mathbf{A}(t)\mathbf{v} \quad (3.85)$$

where $\mathbf{A}(t)$ is the Jacobian of \mathbf{f} (so $A_{ij}(t) = \frac{\partial f_i}{\partial x_j} \Big|_{\boldsymbol{\phi}(t)}$). Since $\boldsymbol{\phi}(t)$ has period T , so does $\mathbf{A}(t)$. Now let $\mathbf{X}(t)$ be the principal fundamental matrix of $\mathbf{v}' = \mathbf{A}\mathbf{v}$ (so that $\mathbf{X}(0) = \mathbf{I}$). Then $\mathbf{B} = \mathbf{X}(T)$.

Now by definition,

$$\phi'(t) = \mathbf{f}(\phi(t)) \quad (3.86)$$

so

$$\phi''(t) = \left. \frac{\partial f_i}{\partial x_j} \right|_{\phi(t)} \phi'(t) \quad (3.87)$$

$$\phi''(t) = \mathbf{A}(t)\phi'(t) \quad (3.88)$$

If we let $\mathbf{v} = \phi'$, then

$$\mathbf{v}'(t) = \mathbf{A}(t)\mathbf{v}(t) \quad (3.89)$$

where, since $\phi(t)$ has period T by assumption, $\mathbf{v}(t)$ must also, and so the corresponding characteristic multiplier is 1. As a result, for a nonlinear system with a periodic solution, one characteristic multiplier is always $\rho = 1$.

3.2 General Results for $n = 2$

3.2.1 Stability of Periodic Solution

Consider a problem of the form $\mathbf{x}' = \mathbf{f}(\mathbf{x})$ with $\mathbf{x} \in \mathbb{R}^2$ where there is a periodic solution $\mathbf{x}(t) = \phi(t)$ with period T . We know from §3.1.2 that we must have $\rho_1 = 1$ and we know from theorem 3.3ii that

$$\rho_1 \rho_2 = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.90)$$

$$\rho_2 = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) \, ds \right). \quad (3.91)$$

From remark (ii) on page 53, we know that for the perturbation to be bounded and hence for the solution to be stable, we must have $\rho_1 \leq 1$ and $\rho_2 \leq 1$ and so, since we know $\rho_1 = 1$ and we wish ρ_1 and ρ_2 to be distinct, we must have

$$0 > \int_0^T \text{tr}(\mathbf{A}(s)) \, ds \quad (3.92)$$

$$0 > \int_0^T \text{tr} \left(\left. \frac{\partial f_i}{\partial x_j} \right|_{\phi(s)} \right) \, ds \quad (3.93)$$

$$0 > \int_0^T \left(\left. \frac{\partial f_1}{\partial x_1} + \frac{\partial f_2}{\partial x_2} \right|_{\phi(s)} \right) \, ds \quad (3.94)$$

$$0 > \int_0^T \nabla \cdot \mathbf{f}|_{\mathbf{x}=\phi} \, ds. \quad (3.95)$$

We get instability when

$$0 < \int_0^T \nabla \cdot \mathbf{f}|_{\mathbf{x}=\phi} \, ds. \quad (3.96)$$

3.2.2 Example

Consider

$$x' = x - y - x(x^2 + y^2) \quad (3.97)$$

$$y' = x + y - y(x^2 + y^2). \quad (3.98)$$

Let

$$x = r(t) \cos(\theta(t)) \quad (3.99)$$

$$y = r(t) \sin(\theta(t)) \quad (3.100)$$

so that our problem becomes

$$\sin(\theta)(r - r\theta') = \cos(\theta)(r - r^3 - r') \quad (3.101)$$

$$\cos(\theta)(r - r\theta') = -\sin(\theta)(r - r^3 - r'). \quad (3.102)$$

By squaring and adding these equations, we obtain that

$$(r - r\theta')^2 = (r - r^3 - r')^2 \quad (3.103)$$

so we can write

$$a = r - r\theta' \quad (3.104)$$

$$sa = r - r^3 - r' \quad (3.105)$$

where $s = \pm 1$. Our equations then become

$$a \sin(\theta) = sa \cos(\theta) \quad (3.106)$$

$$a \cos(\theta) = -sa \sin(\theta) \quad (3.107)$$

which can be rewritten as

$$a \sin(\theta) = sa \cos(\theta) \quad (3.108)$$

$$-s^2 a \sin(\theta) = sa \cos(\theta) \quad (3.109)$$

so that we must have

$$a \sin(\theta) = -a \sin(\theta) \quad (3.110)$$

$$a \sin(\theta) = 0. \quad (3.111)$$

As a result, we have that

$$a \sin(\theta) = sa \cos(\theta) = 0 \quad (3.112)$$

so that we must have $a = 0$. This means that

$$r - r\theta' = r - r^3 - r' = 0. \quad (3.113)$$

We have that

$$r' = r(1 - r^3) \quad (3.114)$$

and so we have a solution of constant radius when $r = 0$ (the trivial case) and $r = \pm 1$. Without loss of generality, choose $r = 1$. Then since

$$r\theta' = r, \quad (3.115)$$

we have that $\theta' = 1$, so $\theta = t + C$. As a result, our solution has period $T = 2\pi$.

Now

$$\nabla \cdot \mathbf{f}|_{r=1} = \left[\frac{\partial f_1}{\partial x} + \frac{\partial f_2}{\partial y} \right]_{r=1} \quad (3.116)$$

$$= [(1 - 3x^2 - y^2) + (1 - x^2 - 3y^2)]_{r=1} \quad (3.117)$$

$$= [2 - 4r^2]_{r=1} \quad (3.118)$$

$$= -2 \quad (3.119)$$

so that

$$\rho_2 = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.120)$$

$$= \exp \left(\int_0^{2\pi} -2 \, ds \right) \quad (3.121)$$

$$= e^{-4\pi} \quad (3.122)$$

$$< 1. \quad (3.123)$$

As a result, the limit cycle with radius $r = 1$ is stable.

3.2.3 Stability of Second-Order ODE

Consider the second-order ODE

$$x'' + a(t)x = 0 \quad (3.124)$$

where $a(t)$ is periodic with period T . Letting $x_1 = x$ and $x_2 = x'_1$, this can be rewritten as

$$\begin{bmatrix} x'_1 \\ x'_2 \end{bmatrix} = \begin{bmatrix} 0 & 1 \\ -a(t) & 0 \end{bmatrix} \begin{bmatrix} x_1 \\ x_2 \end{bmatrix} \quad (3.125)$$

By choosing the initial condition

$$\begin{bmatrix} x_1(0) \\ x_2(0) \end{bmatrix} = \begin{bmatrix} 1 \\ 0 \end{bmatrix} \quad (3.126)$$

we obtain a solution of the form

$$\begin{bmatrix} x_1^{(1)}(t) \\ x_1'^{(1)}(t) \end{bmatrix}. \quad (3.127)$$

Likewise by choosing the initial condition

$$\begin{bmatrix} x_1(0) \\ x_2(0) \end{bmatrix} = \begin{bmatrix} 0 \\ 1 \end{bmatrix} \quad (3.128)$$

we obtain a solution of the form

$$\begin{bmatrix} x_1^{(2)}(t) \\ x_1'^{(2)}(t) \end{bmatrix}. \quad (3.129)$$

As a result, we have chosen $\mathbf{X}(0) = \mathbf{I}$ so that

$$\mathbf{B} = \mathbf{X}(T) = \begin{bmatrix} x_1^{(1)}(T) & x_1^{(2)}(T) \\ x_1'^{(1)}(T) & x_1'^{(2)}(T) \end{bmatrix}. \quad (3.130)$$

Now we have from property (i) on page 52 that

$$\rho_1 \rho_2 = \exp \left(\int_0^T \text{tr}(\mathbf{A}(s)) \, ds \right) \quad (3.131)$$

$$= \exp \left(\int_0^T 0 \, ds \right) \quad (3.132)$$

$$= 1 \quad (3.133)$$

and from property (ii) that

$$\rho_1 + \rho_2 = \text{tr}(\mathbf{B}) \quad (3.134)$$

$$= x_1^{(1)}(T) + x_1'^{(2)}(T). \quad (3.135)$$

Let $\phi = \text{tr}(\mathbf{B})/2$ so that

$$\rho_1 \rho_2 = 1 \quad (3.136)$$

$$\rho_1 + \rho_2 = 2\phi. \quad (3.137)$$

Solving these, we obtain that

$$\rho = \phi \pm \sqrt{\phi^2 - 1}. \quad (3.138)$$

We can rewrite ρ_i as $\exp(\mu_i T)$, so that

$$\mu_1 + \mu_2 = 0 \quad (3.139)$$

and so

$$e^{\mu_1 T} + e^{\mu_2 T} = 2\phi \quad (3.140)$$

$$e^{\mu_1 T} + e^{-\mu_1 T} = 2\phi \quad (3.141)$$

$$\frac{e^{\mu_1 T} + e^{-\mu_1 T}}{2} = \phi \quad (3.142)$$

$$\cosh(\mu_1 T) = \phi. \quad (3.143)$$

Consider the following cases.

- I. Let $-1 < \phi < 1$. We can then define σ by $\phi = \cos(\sigma T)$, where, without loss of generality, $0 < \sigma T < \pi$, so that

$$\rho = \phi \pm \sqrt{\phi^2 - 1} \quad (3.144)$$

$$= \cos(\sigma T) \pm i \sin(\sigma T) \quad (3.145)$$

$$= e^{\pm i\sigma T} \quad (3.146)$$

As in remark (v) on page 55, we can write the general solution as

$$\mathbf{x}(t) = c_1 \operatorname{Re}(e^{i\sigma t} \mathbf{p}(t)) + c_2 \operatorname{Im}(e^{i\sigma t} \mathbf{p}(t)) \quad (3.147)$$

and since $|\rho_1| = 1$ and $|\rho_2| = 1$, then from remark (ii) on page 53, the solution is stable and pseudo-periodic.

Now $e^{i\sigma t}$ has period $\hat{T} = \frac{2\pi}{\sigma}$. Now since $\phi \neq 1$ and $\phi \neq -1$, we must have

$$\sigma T \neq m\pi \quad (3.148)$$

$$\frac{2\pi}{\hat{T}} T \neq m\pi \quad (3.149)$$

$$\frac{2T}{m} \neq \hat{T} \quad (3.150)$$

so that $\hat{T} \neq 2T, T, \frac{2}{3}T, \dots$

Note that for \hat{T} to equal nT , we must have

$$\sigma = \frac{2\pi}{nT} \quad (3.151)$$

for $n \neq 1, 2$ from above.

- II. Let $\phi > 1$. Then since $\rho = \phi \pm \sqrt{\phi^2 - 1}$, we must have $\rho_1 > 1$ and since $\rho_1 \rho_2 = 1$, we must have $\rho_1 > 1 > \rho_2 > 0$ and $\rho_2 = \frac{1}{\rho_1}$ means $\mu_2 = -\mu_1$. Our solution must therefore be of the form

$$\mathbf{x}(t) = c_1 e^{\mu_1 t} \mathbf{p}_1(t) + c_2 e^{-\mu_1 t} \mathbf{p}_2(t) \quad (3.152)$$

where $\mathbf{p}_1(t)$ and $\mathbf{p}_2(t)$ are both periodic with period T . As a result, the solution is unstable.

- III. Let $\phi = 1$. Then $\rho_1 = \rho_2 = 1$. Here, theorem 3.4 only guarantees that we will have *one* solution $\mathbf{x}(t)$ of the form $e^{\mu t} \mathbf{p}(t)$. If \mathbf{B} has two linearly independent eigenvectors, we can find two linearly independent $\mathbf{p}_1(t)$ and $\mathbf{p}_2(t)$ so that the two solutions are both in the standard form. However, if \mathbf{B} only has one eigenvector, we will end up with one solution of the form $\mathbf{p}_1(t)$ (since $\rho = 1$ in this case) and the other of the form $t\mathbf{p}_1(t) + \mathbf{p}_2(t)$. To see this, we replace

$$\begin{bmatrix} \lambda_1 & 0 \\ 0 & \lambda_2 \end{bmatrix} \quad (3.153)$$

with the Jordan block

$$\begin{bmatrix} \lambda & 1 \\ 0 & \lambda \end{bmatrix}. \quad (3.154)$$

As a result, instead of our solution being of the form

$$\mathbf{X}(t) = \mathbf{P}(t) \exp \left(\begin{bmatrix} \lambda_1 & 0 \\ 0 & \lambda_2 \end{bmatrix} t \right) \quad (3.155)$$

$$= \mathbf{P}(t) \begin{bmatrix} e^{\lambda_1 t} & 0 \\ 0 & e^{\lambda_2 t} \end{bmatrix} \quad (3.156)$$

$$= \begin{bmatrix} P_1 e^{\lambda_1 t} & P_2 e^{\lambda_2 t} \\ P_3 e^{\lambda_1 t} & P_4 e^{\lambda_2 t} \end{bmatrix}, \quad (3.157)$$

it will be of the form

$$\mathbf{X}(t) = \mathbf{P}(t) \exp \left(\begin{bmatrix} \lambda & 1 \\ 0 & \lambda \end{bmatrix} t \right) \quad (3.158)$$

$$= \mathbf{P}(t) \begin{bmatrix} e^{\lambda t} & t e^{\lambda t} \\ 0 & e^{\lambda t} \end{bmatrix} \quad (3.159)$$

$$= \begin{bmatrix} P_1 e^{\lambda t} & P_1 t e^{\lambda t} + P_2 e^{\lambda t} \\ P_3 e^{\lambda t} & P_3 t e^{\lambda t} + P_4 e^{\lambda t} \end{bmatrix}. \quad (3.160)$$

See the papers by Akhmedov [1] and Wiesel and Pohlen [30].

- IV. Let $\phi < -1$. Since $\rho = \phi \pm \sqrt{\phi^2 - 1}$, we must have $\rho_1 < -1$ and since $\rho_1 \rho_2 = 1$, we must have $\rho_1 < -1 < \rho_2 < 0$ and $\rho_2 = \frac{1}{\rho_1}$ means $\mu_2 = -\mu_1$. Now we can write $\mu_1 = \frac{i\pi}{T} + \gamma$ so that our solution must be of the form

$$\mathbf{x}(t) = c_1 e^{\gamma t} e^{i\pi t/T} \mathbf{p}_1(t) + c_2 e^{-\gamma t} e^{i\pi t/T} \mathbf{p}_2(t) \quad (3.161)$$

where $\mathbf{p}_1(t)$ and $\mathbf{p}_2(t)$ are both periodic with period T and so $e^{i\pi t/T} \mathbf{p}_1(t)$ and $e^{i\pi t/T} \mathbf{p}_2(t)$ are both periodic with period $2T$. As a result, the solution is unstable.

- V. Let $\phi = -1$. Then $\rho_1 = \rho_2 = -1$. As in the case when $\phi = 1$, we have one solution which is periodic (this time with period $2T$),

$$\mathbf{x}_1(t) = e^{i\pi t/T} \mathbf{p}_1(t) \quad (3.162)$$

and the other which grows linearly with time,

$$\mathbf{x}_2(t) = t e^{i\pi t/T} \mathbf{p}_1(t) + e^{i\pi t/T} \mathbf{p}_2(t). \quad (3.163)$$

We summarise these results in figure 3.1. For $\phi > 1$, we have an unstable solution of the form

$$\mathbf{x}(t) = c_1 e^{\mu_1 t} \mathbf{p}_1(t) + c_2 e^{-\mu_1 t} \mathbf{p}_2(t). \quad (3.164)$$

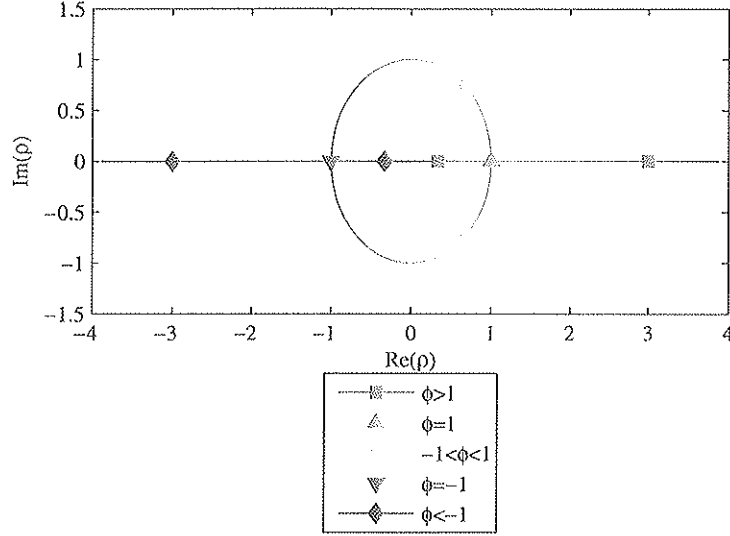


Figure 3.1: The range of ρ_1, ρ_2 for different values of ϕ real. In the region $\phi > 1$, the sample point has $\rho = 1/3, 3$; for $\phi = 1$, we have $\rho = 1$. In $-1 < \phi < 1$, the sample point shown is $\rho = 2/3 \pm i\sqrt{5}/3$; for $\phi = -1$, we have $\rho = -1$ and in the region $\phi < -1$, we show $\rho = -1/3, -3$.

For $\phi = 1$, we have an unstable solution of the form

$$\mathbf{x}(t) = (c_1 + tc_2) \mathbf{p}_1(t) + c_2 \mathbf{p}_2(t). \quad (3.165)$$

For $-1 < \phi < 1$, we have a stable pseudo-periodic solution of the form

$$\mathbf{x}(t) = c_1 \operatorname{Re} \left(e^{i\sigma t} \mathbf{p}(t) \right) + c_2 \operatorname{Im} \left(e^{i\sigma t} \mathbf{p}(t) \right). \quad (3.166)$$

For $\phi = -1$, we have an unstable solution of the form

$$\mathbf{x}(t) = (c_1 + tc_2) \mathbf{q}_1(t) + c_2 \mathbf{q}_2(t). \quad (3.167)$$

Finally, for $\phi < -1$, we have an unstable solution of the form

$$\mathbf{x}(t) = c_1 e^{\gamma t} \mathbf{q}_1(t) + c_2 e^{-\gamma t} \mathbf{q}_2(t) \quad (3.168)$$

where $\mathbf{p}_i(t)$ represents a function that has period T and $\mathbf{q}_i(t)$ represents a function that has period $2T$.

3.2.4 Application to Hill's Equation

Consider Hill's equation

$$x'' + (\delta + \epsilon b(t)) = 0 \quad (3.169)$$

where $b(t)$ has period T . If $\epsilon = 0$, the solution is stable, however, there are some values of δ for which the solution is only marginally stable, according to the above criteria. As a result, we expect that for ϵ small but nonzero near those values of δ , we will get the beginning of a region of instability. We wish to find those values of δ .

For $\epsilon = 0$, if $\mathbf{X}(0) = \mathbf{I}$, then

$$\mathbf{X}(t) = \begin{bmatrix} \cos(\sqrt{\delta}t) & \frac{1}{\sqrt{\delta}} \sin(\sqrt{\delta}t) \\ -\sqrt{\delta} \sin(\sqrt{\delta}t) & \cos(\sqrt{\delta}t) \end{bmatrix} \quad (3.170)$$

and so

$$\mathbf{B} = \mathbf{X}(T) = \begin{bmatrix} \cos(\sqrt{\delta}T) & \frac{1}{\sqrt{\delta}} \sin(\sqrt{\delta}T) \\ -\sqrt{\delta} \sin(\sqrt{\delta}T) & \cos(\sqrt{\delta}T) \end{bmatrix}. \quad (3.171)$$

As a result,

$$\phi = \frac{\text{tr}(\mathbf{B})}{2} = \cos(\sqrt{\delta}T). \quad (3.172)$$

If $\phi = 1$, then

$$\sqrt{\delta}T = 2m\pi \quad (3.173)$$

$$\delta = \left(2m \frac{\pi}{T}\right)^2 \quad (3.174)$$

where m is a positive integer since $\sqrt{\delta} > 0$. If $\phi = -1$, then

$$\sqrt{\delta}\pi = (2m+1)\pi \quad (3.175)$$

$$\delta = \left((2m+1) \frac{\pi}{T}\right)^2. \quad (3.176)$$

Now we have from the previous section that $\phi = 1$ corresponds to the existence of a periodic solution of period T and $\phi = -1$ corresponds to the existence of a periodic solution of period $2T$. As a result, we will have the border between stability and instability breaking off from $\epsilon = 0$ at

$$\delta = \left(2m \frac{\pi}{T}\right)^2 \quad (3.177)$$

corresponding to solutions with period T and breaking off from $\epsilon = 0$ at

$$\delta = \left((2m+1) \frac{\pi}{T}\right)^2 \quad (3.178)$$

corresponding to solutions with period $2T$.

3.3 Stability Boundary of Mathieu's Equation

3.3.1 Undamped Case

We have from §3.2.3 and §3.2.4 that on the edge of the region of stability, we have either $\phi = 1$ or $\phi = -1$. The former corresponds to the existence of a periodic solution with period T and the latter to a periodic solution with period $2T$. In order to determine the region of stability of the Mathieu equation in the δ - ϵ plane, we then need to determine the conditions on δ and ϵ required in order to have a solution which is periodic with either period π or 2π . We follow McLachlan [17] and Ward [28].

Functions of Period π

We can write a general function of period π as

$$x = \sum_{n=0}^{\infty} a_n \cos(2nt) + \sum_{n=1}^{\infty} b_n \sin(2nt). \quad (3.179)$$

We then obtain

$$0 = x'' + (\delta + \epsilon \cos(2t))x \quad (3.180)$$

$$\begin{aligned} 0 = & \sum_{n=0}^{\infty} (\delta - 4n^2) a_n \cos(2nt) + \sum_{n=1}^{\infty} (\delta - 4n^2) b_n \sin(2nt) \\ & + \epsilon \sum_{n=0}^{\infty} a_n \cos(2nt) \cos(2t) + \epsilon \sum_{n=1}^{\infty} b_n \sin(2nt) \cos(2t). \end{aligned} \quad (3.181)$$

Using the identities

$$\cos(A) \cos(B) = \frac{1}{2} (\cos(A-B) + \cos(A+B)) \quad (3.182)$$

$$\sin(A) \cos(B) = \frac{1}{2} (\sin(A-B) + \sin(A+B)) \quad (3.183)$$

this becomes

$$\begin{aligned} 0 = & \sum_{n=0}^{\infty} (\delta - 4n^2) a_n \cos(2nt) + \sum_{n=1}^{\infty} (\delta - 4n^2) b_n \sin(2nt) \\ & + \frac{\epsilon}{2} \sum_{n=0}^{\infty} a_n (\cos(2(n+1)t) + \cos(2(n-1)t)) \\ & + \frac{\epsilon}{2} \sum_{n=1}^{\infty} b_n (\sin(2(n+1)t) + \sin(2(n-1)t)) \end{aligned} \quad (3.184)$$

and so we must have

$$0 = \sum_{n=0}^{\infty} (\delta - 4n^2) a_n \cos(2nt) + \frac{\epsilon}{2} \sum_{n=0}^{\infty} a_n (\cos(2(n+1)t) + \cos(2(n-1)t)) \quad (3.185)$$

$$0 = \left(\delta a_0 + \frac{\epsilon}{2} a_1 \right) \cos(0) + \left((\delta - 4) a_1 + \frac{\epsilon}{2} (2a_0 + a_2) \right) \cos(2t) + \sum_{n=2}^{\infty} \left((\delta - 4n^2) a_n + \frac{\epsilon}{2} (a_{n-1} + a_{n+1}) \right) \cos(2nt) \quad (3.186)$$

and

$$0 = \sum_{n=1}^{\infty} (\delta - 4n^2) b_n \sin(2nt) + \frac{\epsilon}{2} \sum_{n=1}^{\infty} b_n (\sin(2(n+1)t) + \sin(2(n-1)t)) \quad (3.187)$$

$$0 = \left((\delta - 4) b_1 + \frac{\epsilon}{2} b_2 \right) \sin(2t) + \sum_{n=2}^{\infty} \left((\delta - 4n^2) b_n + \frac{\epsilon}{2} (b_{n-1} + b_{n+1}) \right) \sin(2nt). \quad (3.188)$$

By orthogonality of the sine and cosine, these can be rewritten as

$$\begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ \vdots \end{bmatrix} = \begin{bmatrix} \delta & \frac{\epsilon}{2} & & & 0 \\ \epsilon & \delta - 4 \cdot 1^2 & \frac{\epsilon}{2} & & \\ & \frac{\epsilon}{2} & \delta - 4 \cdot 2^2 & \frac{\epsilon}{2} & \\ & & \frac{\epsilon}{2} & \delta - 4 \cdot 3^2 & \frac{\epsilon}{2} \\ 0 & & & \ddots & \ddots & \ddots \end{bmatrix} \begin{bmatrix} a_0 \\ a_1 \\ a_2 \\ a_3 \\ \vdots \end{bmatrix} \quad (3.189)$$

and

$$\begin{bmatrix} 0 \\ 0 \\ 0 \\ \vdots \end{bmatrix} = \begin{bmatrix} \delta - 4 \cdot 1^2 & \frac{\epsilon}{2} & & & 0 \\ & \frac{\epsilon}{2} & \delta - 4 \cdot 2^2 & \frac{\epsilon}{2} & \\ & & \frac{\epsilon}{2} & \delta - 4 \cdot 3^2 & \frac{\epsilon}{2} \\ 0 & & & \ddots & \ddots & \ddots \end{bmatrix} \begin{bmatrix} b_1 \\ b_2 \\ b_3 \\ \vdots \end{bmatrix} \quad (3.190)$$

In order to have a non-zero solution, the determinant of at least one of these (infinite) matrices must be zero. This gives us the requirement that ϵ and δ must satisfy in order to be on the borderline between stability and instability. We can approximate the determinants of these matrices by the determinants of the finite $n \times n$ matrices of the same form. The resultant curves in the δ - ϵ plane for different values of n are shown in figure 3.2.

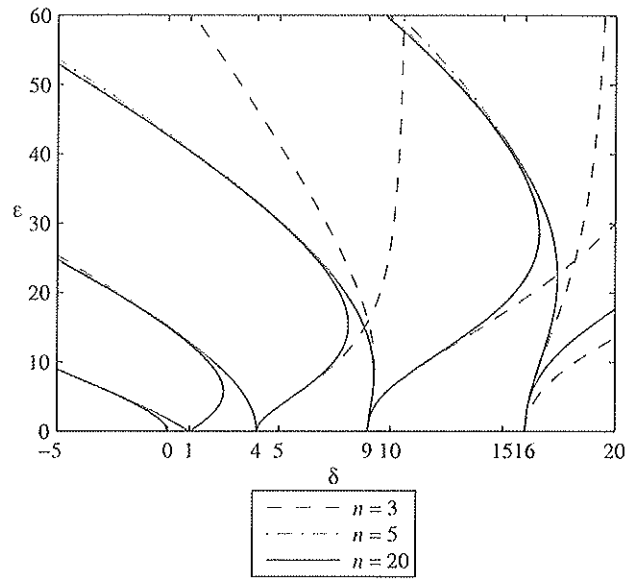


Figure 3.2: The approximation to the border of the region of stability of the Mathieu equation (determined by equations 3.189, 3.190, 3.198, 3.199) where each infinite matrix is approximated by its $n \times n$ counterpart.

Functions of Period 2π

We now perform a similar analysis for functions of period 2π . We can write a general function of period 2π as

$$x = \sum_{n=0}^{\infty} a_n \cos(nt) + \sum_{n=1}^{\infty} b_n \sin(nt). \quad (3.191)$$

We then remove from this all the terms which also have period π since we have already dealt with those. If we included them, we would obtain the lines in the δ - ϵ plane where we obtain solutions that either have period π or have period 2π . As a result, we have

$$x = \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} a_n \cos(nt) + \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} b_n \sin(nt). \quad (3.192)$$

so that we obtain

$$0 = x'' + (\delta + \epsilon \cos(2t)) x \quad (3.193)$$

$$\begin{aligned} 0 = & \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} (\delta - n^2) a_n \cos(nt) + \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} (\delta - n^2) b_n \sin(nt) \\ & + \epsilon \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} a_n \cos(nt) \cos(2t) + \epsilon \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} b_n \sin(nt) \cos(2t) \end{aligned} \quad (3.194)$$

$$\begin{aligned} 0 = & \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} (\delta - n^2) a_n \cos(nt) + \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} (\delta - n^2) b_n \sin(nt) \\ & + \frac{\epsilon}{2} \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} a_n (\cos((n+2)t) + \cos((n-2)t)) \\ & + \frac{\epsilon}{2} \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} b_n (\sin((n+2)t) + \sin((n-2)t)). \end{aligned} \quad (3.195)$$

We must then have

$$\begin{aligned} 0 = & \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} \left((\delta - 1) a_1 + \frac{\epsilon}{2} (a_1 + a_3) \right) \cos(t) \\ & + \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} \left((\delta - n^2) a_n + \frac{\epsilon}{2} (a_{n-2} + a_{n+2}) \right) \cos(nt) \end{aligned} \quad (3.196)$$

and

$$0 = \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} \left((\delta - 1) b_1 + \frac{\epsilon}{2} (-b_1 + b_3) \right) \sin(t) \\ + \sum_{\substack{n=1 \\ n \text{ odd}}}^{\infty} \left((\delta - n^2) b_n + \frac{\epsilon}{2} (b_{n-2} + b_{n+2}) \right) \sin(nt) \quad (3.197)$$

which we can write as

$$\begin{bmatrix} 0 \\ 0 \\ 0 \\ \vdots \end{bmatrix} = \begin{bmatrix} \delta - 1^2 + \frac{\epsilon}{2} & \frac{\epsilon}{2} & & 0 \\ & \frac{\epsilon}{2} & \delta - 3^2 & \frac{\epsilon}{2} \\ & & \frac{\epsilon}{2} & \delta - 5^2 & \frac{\epsilon}{2} \\ 0 & & & \ddots & \ddots & \ddots \end{bmatrix} \begin{bmatrix} a_1 \\ a_3 \\ a_5 \\ \vdots \end{bmatrix} \quad (3.198)$$

and

$$\begin{bmatrix} 0 \\ 0 \\ 0 \\ \vdots \end{bmatrix} = \begin{bmatrix} \delta - 1^2 - \frac{\epsilon}{2} & \frac{\epsilon}{2} & & 0 \\ & \frac{\epsilon}{2} & \delta - 3^2 & \frac{\epsilon}{2} \\ & & \frac{\epsilon}{2} & \delta - 5^2 & \frac{\epsilon}{2} \\ 0 & & & \ddots & \ddots & \ddots \end{bmatrix} \begin{bmatrix} b_1 \\ b_3 \\ b_5 \\ \vdots \end{bmatrix} \quad (3.199)$$

As before, in order to obtain a nonzero solution, we must have the determinant of at least one of the matrices being zero. This constrains δ and ϵ .

The resultant region of stability is shown in figure 3.3

3.3.2 Undamped Case with ϵ small

Consider now when ϵ is small. We have from §3.2.4 that for ϵ small, we will have the border between stability and instability near

$$\delta = (2m)^2 \quad (3.200)$$

and

$$\delta = (2m + 1)^2. \quad (3.201)$$

As a result, we seek periodic solutions near $\delta = n^2$ to the equation

$$x'' + (\delta + \epsilon \cos(2t)) x = 0. \quad (3.202)$$

Let

$$x = x_0(t) + \epsilon x_1(t) + \epsilon^2 x_2(t) + \dots, \quad (3.203)$$

$$\delta = n^2 + \epsilon \delta_1 + \epsilon^2 \delta_2 + \dots \quad (3.204)$$

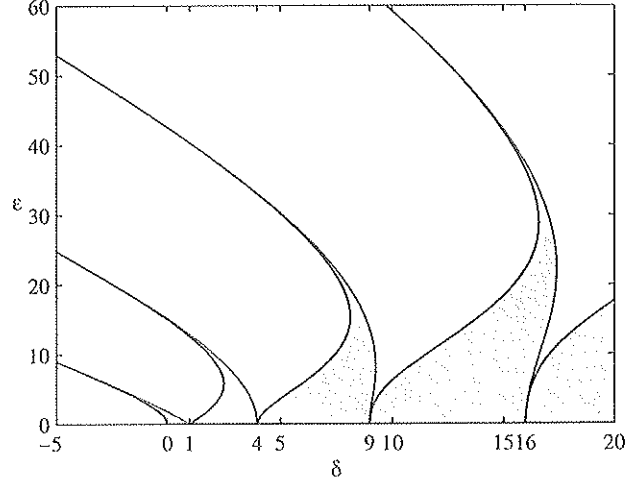


Figure 3.3: The region of stability of the Mathieu equation.

Substituting these into Mathieu's equation, we obtain

$$x_0'' + n^2 x_0 = 0 \quad (3.205)$$

$$x_1'' + n^2 x_1 = -\delta_1 x_0 - x_0 \cos(2t) \quad (3.206)$$

$$x_2'' + n^2 x_2 = -\delta_1 x_1 - \delta_2 x_0 - x_1 \cos(2t). \quad (3.207)$$

For $n \neq 0$, the solution to equation 3.205 is

$$x_0 = a \cos(nt) + b \sin(nt). \quad (3.208)$$

Inserting this into equation 3.206, we obtain

$$x_1'' + n^2 x_1 = -\delta_1 x_0 - x_0 \cos(2t) \quad (3.209)$$

$$\begin{aligned} &= -\delta_1 (a \cos(nt) + b \sin(nt)) \\ &\quad - (a \cos(nt) + b \sin(nt)) \cos(2t) \end{aligned} \quad (3.210)$$

$$\begin{aligned} &= -\delta_1 a \cos(nt) - \delta_1 b \sin(nt) \\ &\quad - \frac{a}{2} \cos((n+2)t) - \frac{a}{2} \cos((n-2)t) \\ &\quad - \frac{b}{2} \sin((n+2)t) - \frac{b}{2} \sin((n-2)t) \end{aligned} \quad (3.211)$$

Under the assumption that $n \neq 1$, in order to eliminate secular terms, we must have

$$-\delta_1 a = 0, \quad -\delta_1 b = 0. \quad (3.212)$$

As a result, in order to avoid x_0 being the zero solution, we must have $\delta_1 = 0$. We then have

$$\begin{aligned} x_1'' + n^2 x_1 = & -\frac{a}{2} \cos((n+2)t) - \frac{a}{2} \cos((n-2)t) \\ & - \frac{b}{2} \sin((n+2)t) - \frac{b}{2} \sin((n-2)t). \end{aligned} \quad (3.213)$$

Letting

$$x_1 = \sum_{i=0}^{\infty} c_i \sin(it) + d_i \cos(it), \quad (3.214)$$

this becomes

$$\begin{aligned} - \sum_{i=1}^{\infty} c_i i^2 \sin(it) + d_i i^2 \cos(it) + \sum_{i=0}^{\infty} c_i n^2 n^2 \sin(it) + d_i n^2 \cos(it) \\ = -\frac{a}{2} \cos((n+2)t) - \frac{a}{2} \cos((n-2)t) \\ - \frac{b}{2} \sin((n+2)t) - \frac{b}{2} \sin((n-2)t). \end{aligned} \quad (3.215)$$

Equating coefficients of the sines and cosines, we obtain that

$$c_{n-2} = \frac{b}{8(-n+1)}, \quad d_{n-2} = \frac{a}{8(-n+1)} \quad (3.216)$$

$$c_{n+2} = \frac{b}{8(n+1)}, \quad d_{n+2} = \frac{a}{8(n+1)}. \quad (3.217)$$

We can assume that all the $\sin(nt)$ and $\cos(nt)$ component is already in x_0 , so we can choose $c_n = 0$, $d_n = 0$. All remaining c_i and d_i are zero. As a result,

$$\begin{aligned} x_1 = & \frac{b}{8(-n+1)} \sin((n-2)t) + \frac{b}{8(n+1)} \sin((n+2)t) \\ & + \frac{a}{8(-n+1)} \cos((n-2)t) + \frac{a}{8(n+1)} \cos((n+2)t). \end{aligned} \quad (3.218)$$

Finally, inserting this into equation 3.207, we obtain that

$$\begin{aligned} x_2'' + n^2 x_2 = & -\delta_2 (a \cos(nt) + b \sin(nt)) \\ & - \frac{b}{16(-n+1)} (\sin(nt) + \sin((n-4)t)) \\ & - \frac{b}{16(n+1)} (\sin((n+4)t) + \sin(nt)) \\ & - \frac{a}{16(-n+1)} (\cos(nt) + \cos((n-4)t)) \\ & - \frac{a}{16(n+1)} (\cos((n+4)t) + \cos(nt)). \end{aligned} \quad (3.219)$$

Under the assumption that $n \neq 2$, in order to eliminate the secular terms, we must have

$$0 = -\delta_2 a - \frac{a}{16(-n+1)} - \frac{a}{16(n+1)}, \quad (3.220)$$

$$0 = -\delta_2 b - \frac{b}{16(-n+1)} - \frac{b}{16(n+1)}, \quad (3.221)$$

which can be rewritten as

$$0 = -a \left(\delta_2 - \frac{1}{8(n^2-1)} \right), \quad (3.222)$$

$$0 = -b \left(\delta_2 - \frac{1}{8(n^2-1)} \right). \quad (3.223)$$

As a result, in order to avoid a nonzero x_0 (i.e., making sure that we don't simultaneously have $a = 0$ and $b = 0$), we must have

$$\delta_2 = \frac{1}{8(n^2-1)}. \quad (3.224)$$

Case $n = 2$

In the case $n = 2$, eliminating the secular terms in equation 3.219 tells us that

$$0 = -\delta_2 a + \frac{a}{8} - \frac{a}{48}, \quad (3.225)$$

$$0 = -\delta_2 b - 0 - \frac{b}{48}, \quad (3.226)$$

which become

$$0 = -a \left(\delta_2 - \frac{5}{48} \right), \quad (3.227)$$

$$0 = -b \left(\delta_2 + \frac{1}{48} \right). \quad (3.228)$$

As a result, for $n = 2$ we must have either

$$a = 0, \quad \delta_2 = -\frac{1}{48} \quad (3.229)$$

or

$$b = 0, \quad \delta_2 = \frac{5}{48}. \quad (3.230)$$

As a result, for $n = 2$, we either have

$$\delta = 4 - \epsilon^2 \frac{1}{48} + O(\epsilon^3) \quad (3.231)$$

or

$$\delta = 4 + \epsilon^2 \frac{5}{48} + O(\epsilon^3). \quad (3.232)$$

We also have either

$$x = b \sin(2t) + \epsilon \frac{b}{24} \sin(4t) + O(\epsilon^2) \quad (3.233)$$

or

$$x = a \cos(2t) + \epsilon \left(-\frac{a}{8} + \frac{a}{24} \cos(4t) \right) + O(\epsilon^2), \quad (3.234)$$

which both have period π , as expected.

Case $n = 1$

In the case $n = 1$, eliminating the secular terms in equation 3.211 tells us that

$$0 = -\delta_1 a - \frac{a}{2}, \quad (3.235)$$

$$0 = -\delta_1 b + \frac{b}{2} \quad (3.236)$$

and so we must either have

$$\delta_1 = -\frac{1}{2}, \quad b = 0 \quad (3.237)$$

or

$$\delta_1 = \frac{1}{2}, \quad a = 0. \quad (3.238)$$

In either of these cases, equation 3.211 becomes

$$x_1'' + x_1 = -\frac{a}{2} \cos(3t) - \frac{b}{2} \sin(3t). \quad (3.239)$$

As before, we let

$$x_1 = \sum_{i=1}^{\infty} c_i \sin(it) + d_i \cos(it) \quad (3.240)$$

and find that

$$c_3 = \frac{b}{16}, \quad d_3 = \frac{a}{16}. \quad (3.241)$$

As a result,

$$x_1 = \frac{b}{16} \sin(3t) + \frac{a}{16} \cos(3t). \quad (3.242)$$

Then equation 3.207 becomes

$$\begin{aligned} x_2'' + x_2 = & -\delta_1 \left(\frac{b}{16} \sin(3t) + \frac{a}{16} \cos(3t) \right) \\ & - \delta_2 (a \cos(t) + b \sin(t)) \\ & - \left(\frac{b}{16} \sin(3t) + \frac{a}{16} \cos(3t) \right) \cos(2t) \end{aligned} \quad (3.243)$$

$$\begin{aligned} = & -\delta_1 \left(\frac{b}{16} \sin(3t) + \frac{a}{16} \cos(3t) \right) - \delta_2 a \cos(t) - \delta_2 b \sin(t) \\ & - \frac{b}{32} \sin(t) - \frac{b}{32} \sin(5t) - \frac{a}{32} \cos(t) - \frac{a}{32} \cos(5t). \end{aligned} \quad (3.244)$$

In order to eliminate the secular terms, we must have

$$0 = -a \left(\delta_2 + \frac{1}{32} \right) \quad (3.245)$$

$$0 = -b \left(\delta_2 + \frac{1}{32} \right). \quad (3.246)$$

As a result, $\delta_2 = -1/32$, so that either

$$\delta = 1 - \epsilon \frac{1}{2} - \epsilon^2 \frac{1}{32} + O(\epsilon^3) \quad (3.247)$$

or

$$\delta = 1 + \epsilon \frac{1}{2} - \epsilon^2 \frac{1}{32} + O(\epsilon^3). \quad (3.248)$$

We also have either

$$x = a \cos(t) + \epsilon \frac{a}{16} \cos(3t) + O(\epsilon^2) \quad (3.249)$$

or

$$x = b \sin(t) + \epsilon \frac{b}{16} \sin(3t) + O(\epsilon^2), \quad (3.250)$$

which are periodic with period 2π , as expected.

Case $n = 0$

In the case $n = 0$, we get

$$x_0 = a + bt. \quad (3.251)$$

Now we expect a periodic solution, so $b = 0$. As a result, equation 3.206 becomes

$$x_1'' = -\delta_1 a - a \cos(2t). \quad (3.252)$$

In analogy with before, when we eliminated secular terms, we must have $\delta_1 = 0$.

As a result, we have

$$x_1'' = -a \cos(2t) \quad (3.253)$$

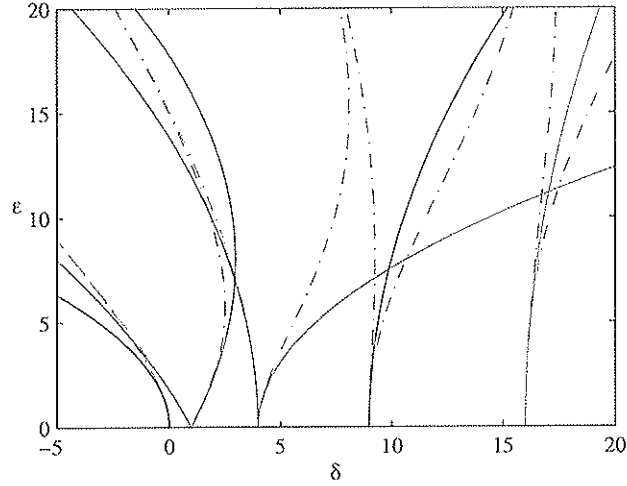


Figure 3.4: The quadratic approximations to the boundary between stability and instability of the Mathieu equation, in comparison with the approximation from §3.3.1, with $n = 20$.

so that

$$x_1 = \frac{a}{4} \cos(2t) \quad (3.254)$$

and equation 3.207 becomes

$$x_2'' = -\delta_2 a - \frac{a}{4} \cos(2t) \cos(2t) \quad (3.255)$$

$$= -\delta_2 a - \frac{a}{8} - \frac{a}{8} \cos(4t) \quad (3.256)$$

so that we must have

$$0 = -a \left(\delta_2 + \frac{1}{8} \right). \quad (3.257)$$

so that $\delta_2 = -1/8$ and

$$\delta = 0 - \epsilon^2 \frac{1}{8} \quad (3.258)$$

with

$$x = a + \epsilon \frac{a}{4} \cos(2t) + O(\epsilon^2), \quad (3.259)$$

which is again periodic with period π , as expected.

These approximations to $\delta(\epsilon)$ for ϵ small are compared to the approximation in the previous section (which is valid for both small and large ϵ) in figure 3.4.

3.3.3 Damped Case

We follow Richards [24]. Our equation is

$$x'' + kx' + (\delta + \epsilon \cos(2t))x = 0. \quad (3.260)$$

If we let

$$y(t) = e^{\frac{k}{2}t}x(t), \quad (3.261)$$

we obtain that

$$y'' + (a + \epsilon \cos(2t))y = 0 \quad (3.262)$$

where

$$a = \delta - \frac{k^2}{4}. \quad (3.263)$$

Now equation 3.260 isn't of the form of equation 3.124 (§3.2.3), but equation 3.262 is. As a result, we know that the solution to equation 3.262 is of the form

$$y(t) = e^{\mu_1 t}p_1(t) + e^{\mu_2 t}p_2(t) \quad (3.264)$$

where μ_1 and μ_2 satisfy

$$e^{\mu\pi} = \rho = \phi \pm \sqrt{\phi^2 - 1} \quad (3.265)$$

where ϕ is half of the trace of \mathbf{B} for $y(t)$ above when we use the initial conditions $\mathbf{X}(0) = \mathbf{I}$. As a result, the largest μ (the one most likely to cause instability) satisfies

$$e^{\mu\pi} = \rho = \phi + \sqrt{\phi^2 - 1} \quad (3.266)$$

so that

$$\mu\pi = \ln(\phi + \sqrt{\phi^2 - 1}) \quad (3.267)$$

$$\mu\pi = \cosh^{-1}(\phi) \quad (3.268)$$

$$\mu = \frac{\cosh^{-1}(\phi)}{\pi}. \quad (3.269)$$

Now in order for $x(t)$ to be stable, we must have

$$0 \geq \operatorname{Re}\left(\mu - \frac{k}{2}\right) \quad (3.270)$$

$$\frac{k}{2} \geq \operatorname{Re}(\mu) \quad (3.271)$$

with μ as above. This can be used to numerically determine the stability of the damped equation. The result for $k = 0.2$ is shown in figure 3.5.

3.3.4 Damped Case with ϵ small

Consider the damped Mathieu equation

$$x'' + kx' + (\delta + \epsilon \cos(2t))x = 0. \quad (3.272)$$

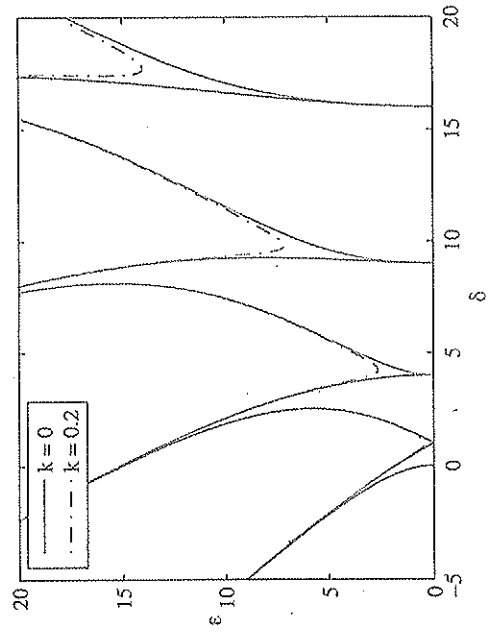


Figure 3.5: The border of the region of stability of the Mathieu equation, in the damped case.